

# Taking physics forward with



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30 May 2023

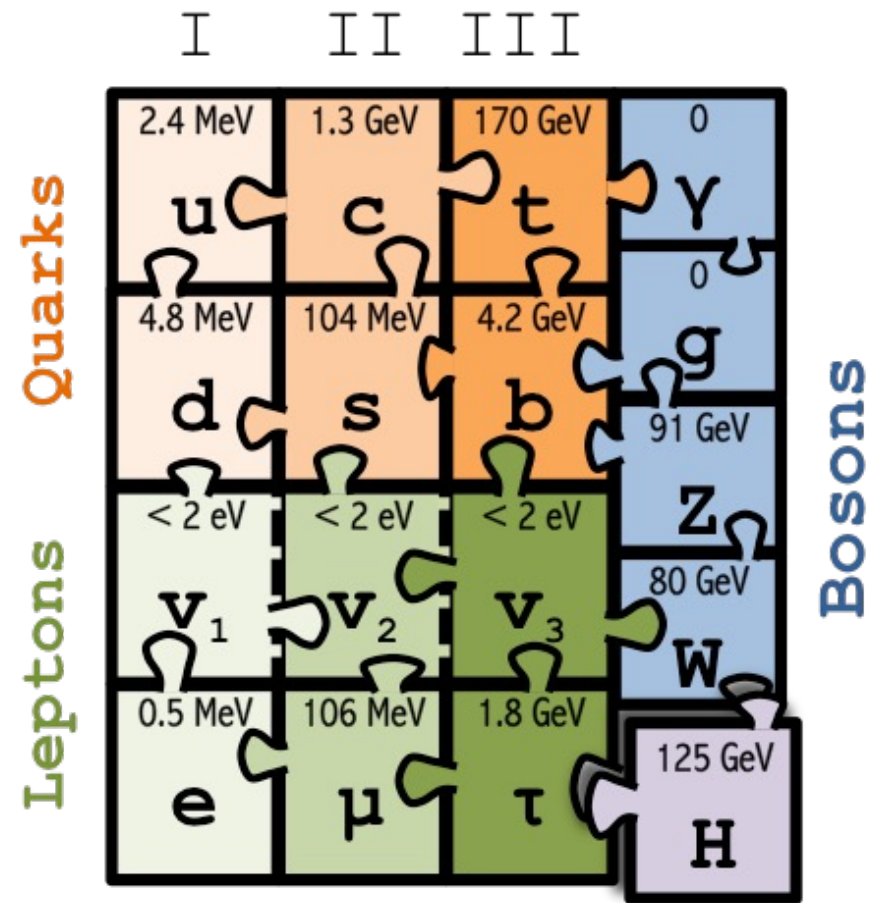
Anna Sfyrla



**UNIVERSITÉ  
DE GENÈVE**

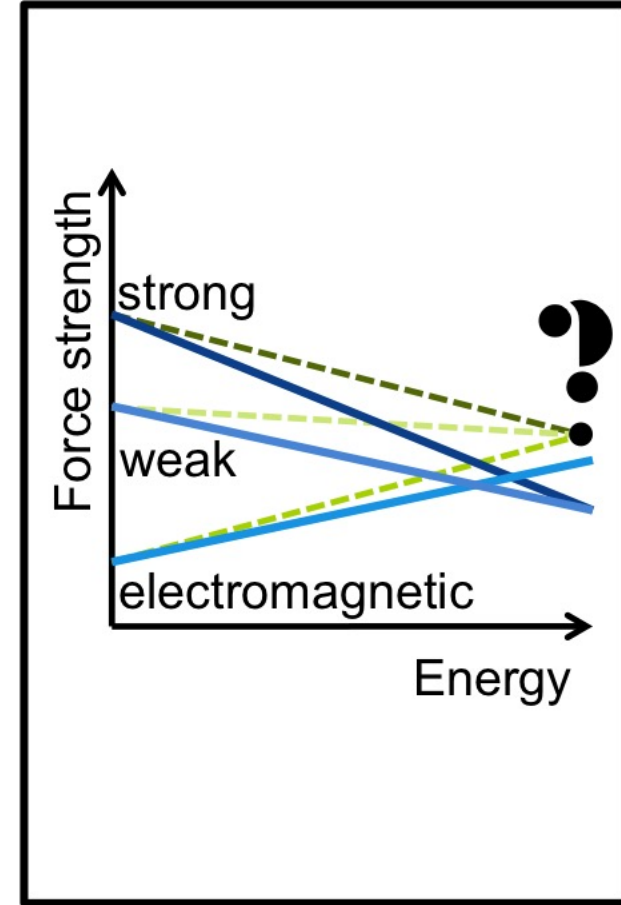
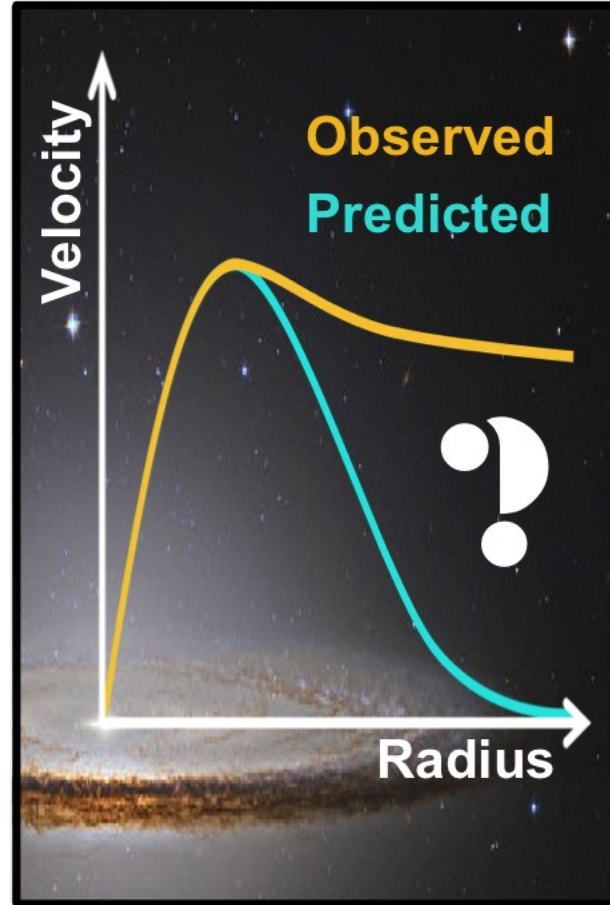
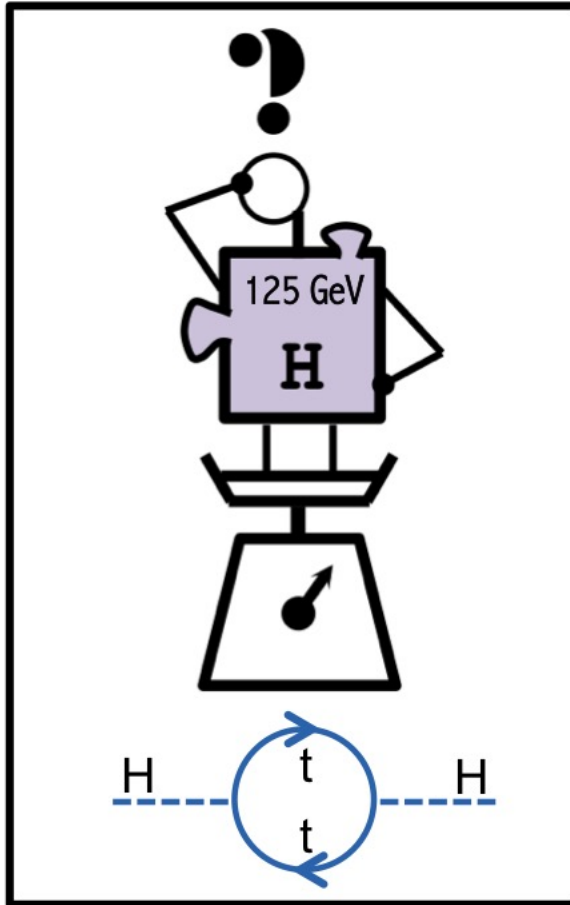
FACULTY OF SCIENCE

# THE STANDARD MODEL





# ...ITS BIGGEST OPEN QUESTIONS ...



# ...AND ITS MORE SUBTLE ONES!

The “strong CP problem”: Why does QCD preserve CP symmetry?

Within the SM, the QCD vacuum structure introduces a CP violating term in the Lagrangian:

$$L_\theta = \theta \frac{g^2}{32\pi^2} F_a^{\mu\nu} \tilde{F}_{a\mu\nu}$$

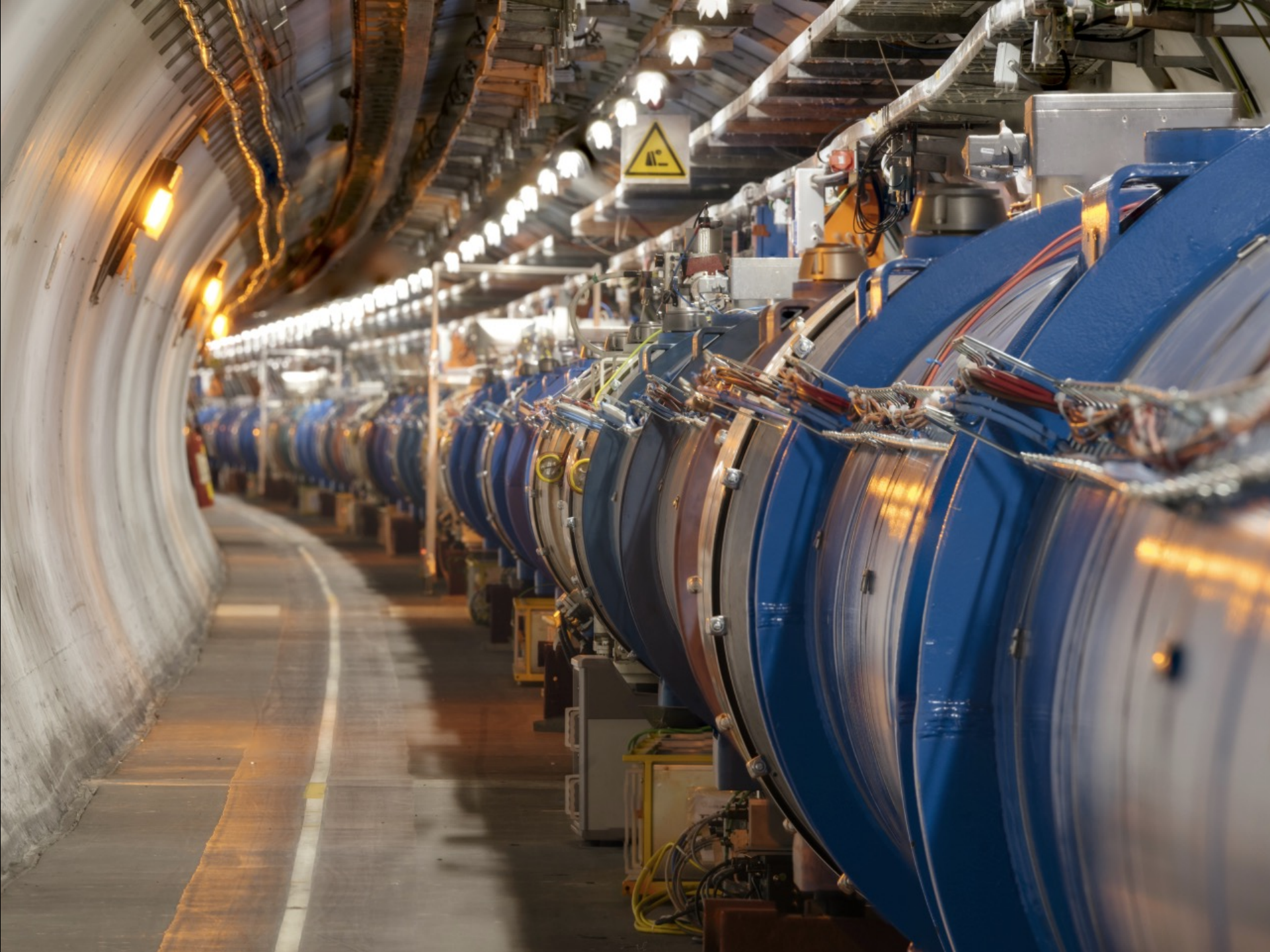
while [measurements](#) require that the vacuum angle  $\theta$  is tiny!

The non-zero angle  $\theta$  implies non-zero neutron electric dipole moment (EDM)

The angle is stringently constrained by neutron EDM measurements

Most sensitive measurement on the neutron EDM to date achieved by the **PSI** experiment **nEDM**:

$$d_n = (0.0 \pm 1.1_{\text{stat}} \pm 0.2_{\text{sys}}) \times 10^{-26} \text{ e}\cdot\text{cm}.$$

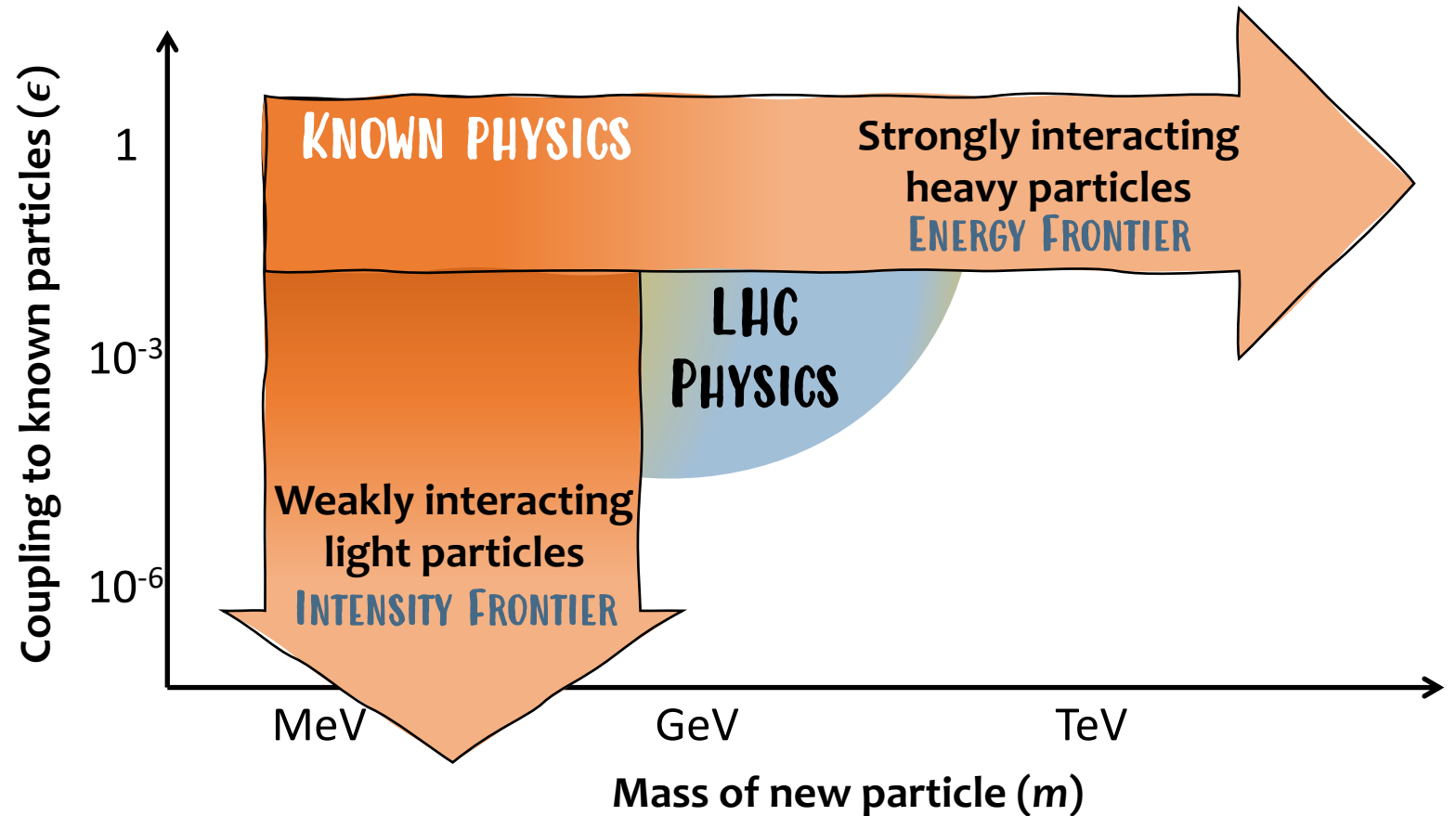




# THE LANDSCAPE OF NEW PARTICLES @ COLLIDERS

- Collider physics: a plethora of measurements and searches
- The Standard Model is complete and confirmed; Burning questions still remain!

	2.4 MeV	1.3 GeV	170 GeV	0
	u	c	t	$\gamma$
	4.8 MeV	104 MeV	4.2 GeV	0
	d	s	b	g
	<2 eV	<2 eV	<2 eV	91 GeV
	$\nu_L$	$\nu_M$	$\nu_H$	Z
	0.5 MeV	16 MeV	1.8 GeV	80 GeV
	e	$\mu$	$\tau$	W
				126 GeV
				H



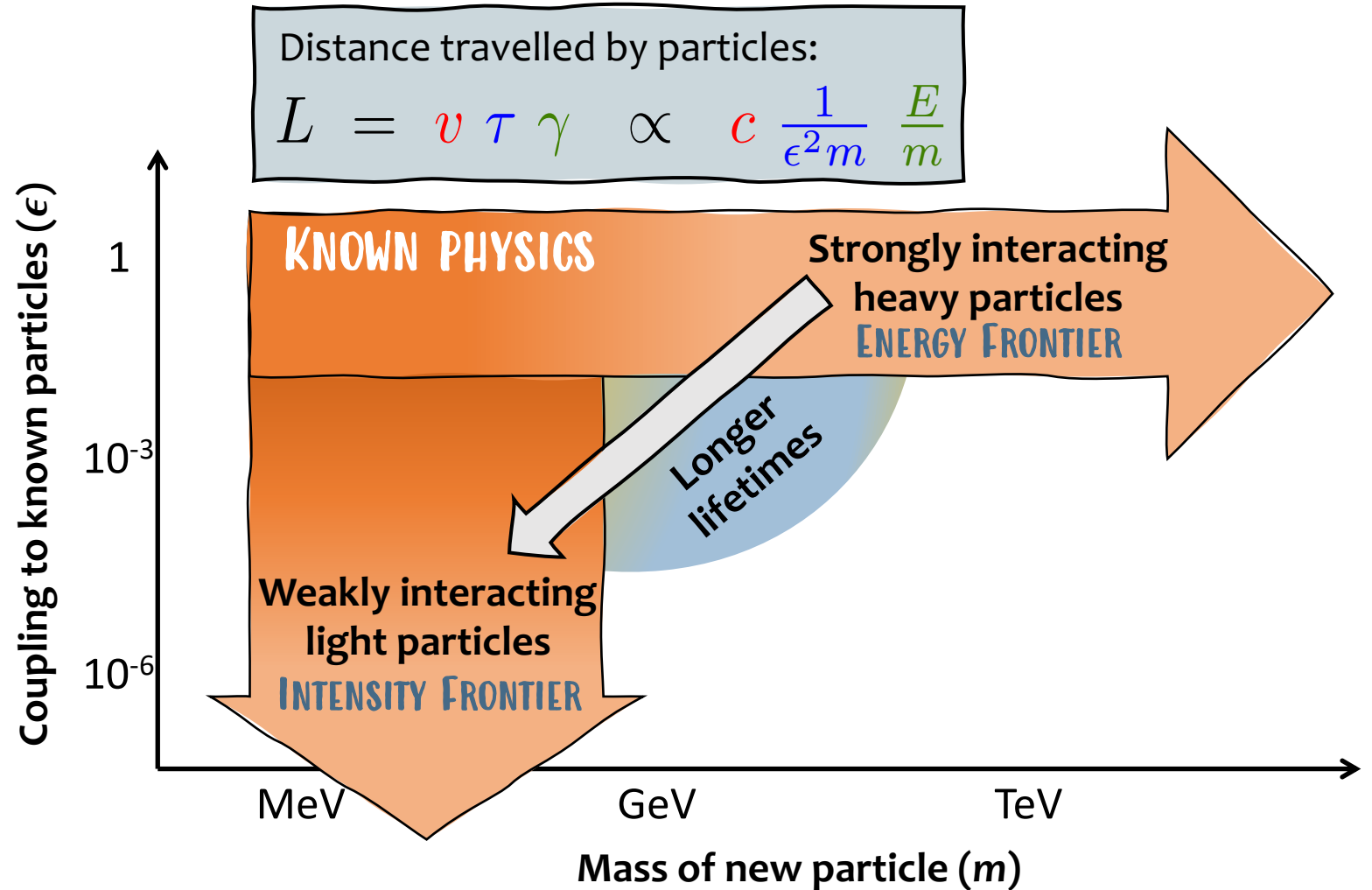
# THE LANDSCAPE OF NEW PARTICLES @ COLLIDERS

## Lifetime

a characteristic of weakly interacting light particles

Distinct signatures

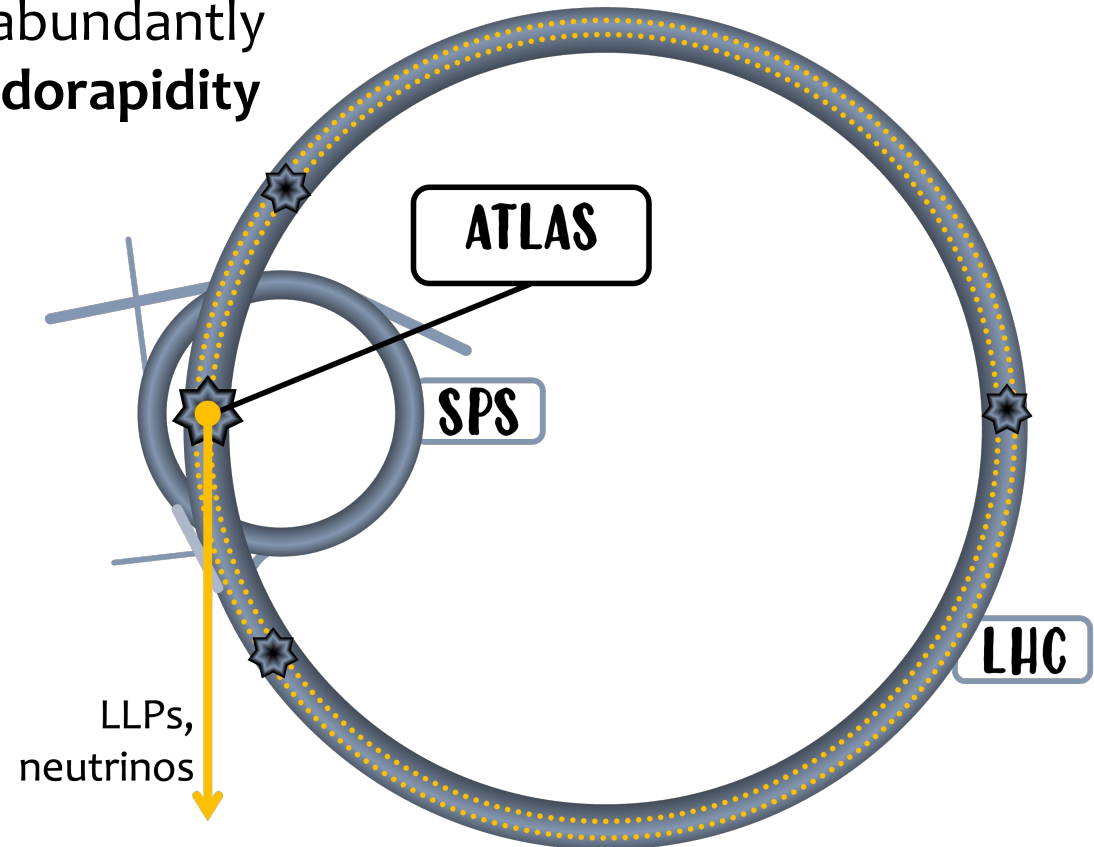
Opportunity for exploration!



# FORWARD SEARCH EXPERIMENT AT THE LHC

Searches for new weakly interacting light particles, coupling to SM via mixing with SM “portal” operator

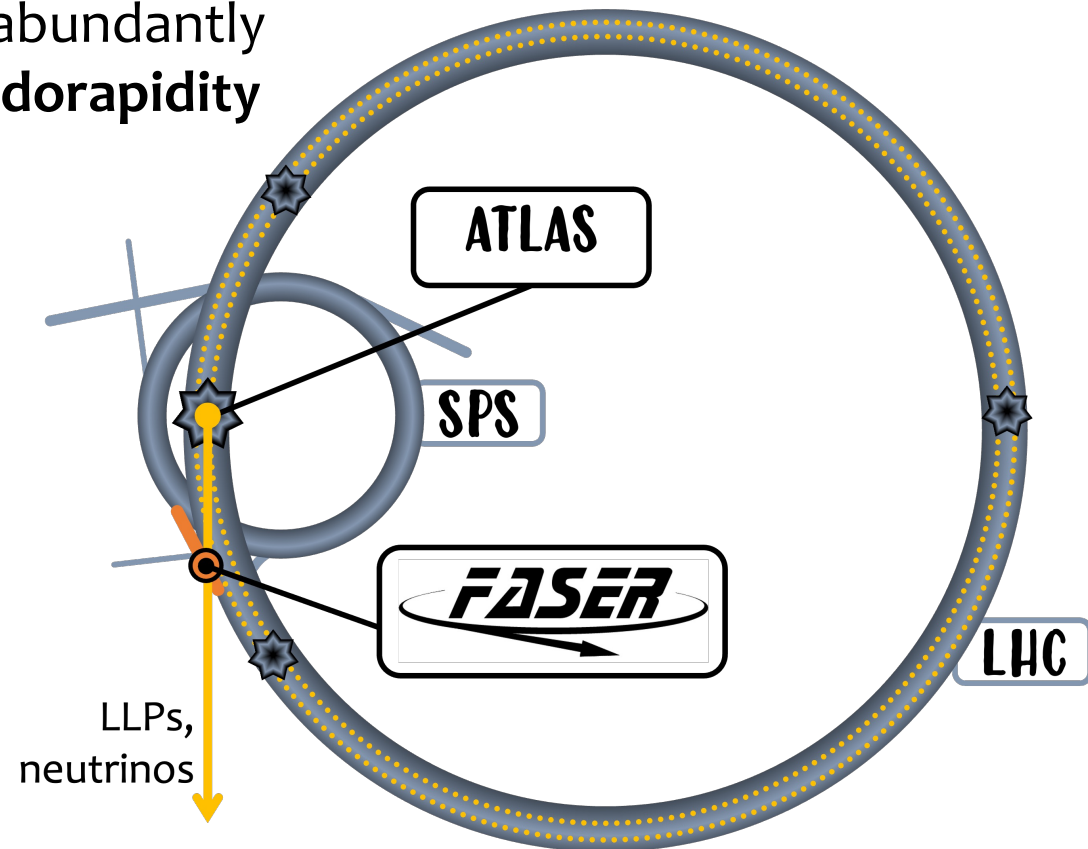
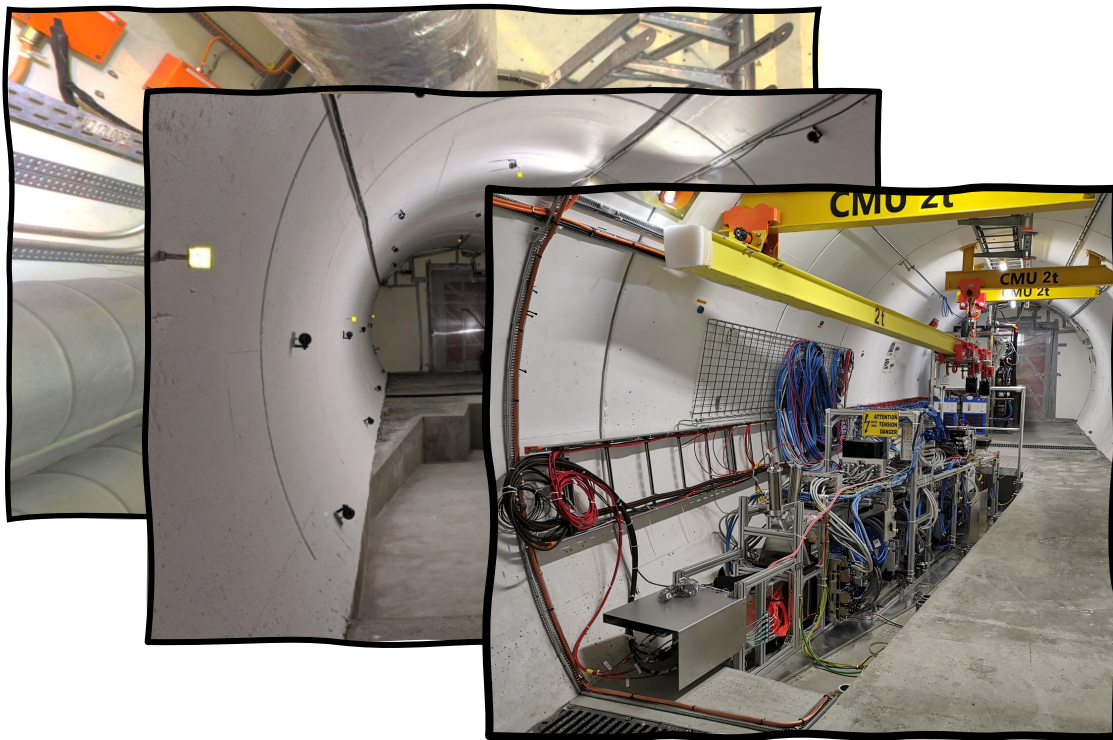
- Produced in decays of light mesons (e.g.  $\pi$ ,  $K$ ), abundantly present in p-p collisions, **primarily in large pseudorapidity**



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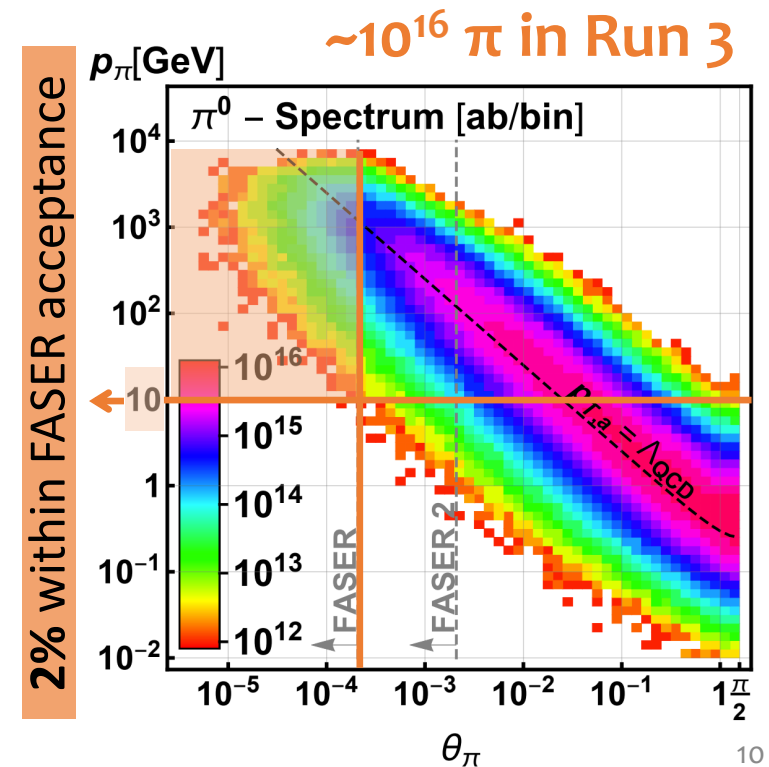
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- Produced in decays of light mesons (e.g.  $\pi$ , K), abundantly present in p-p collisions, **primarily in large pseudorapidity**
- **FASER acceptance:**
  - 20 cm diameter, 480 m from ATLAS IP
  - (ONLY  $10^{-6}\%$  solid angle)
  - but still  $O(10^{16}) \pi$  in Run 3!





# FORWARD SEARCH EXPERIMENT AT THE LHC

Searches for new weakly interacting light particles, coupling to SM via mixing with SM “portal” operator

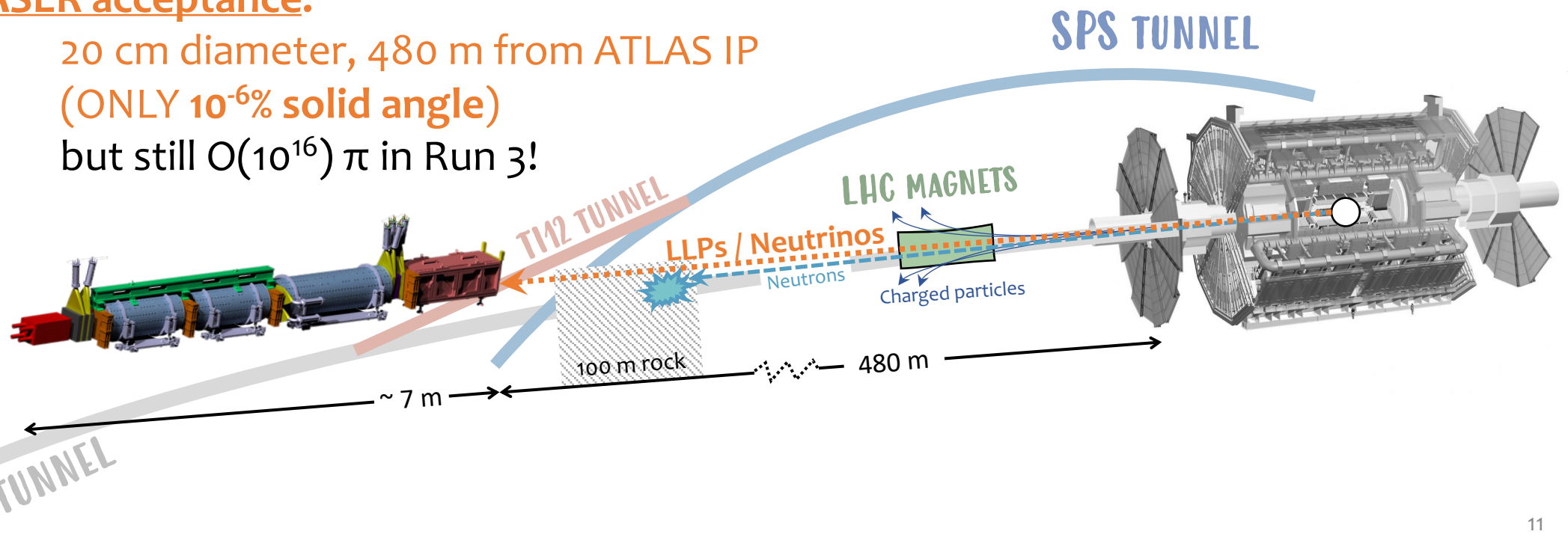
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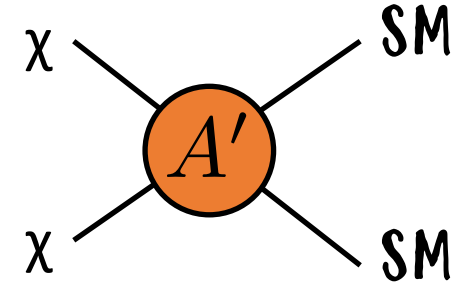
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(ONLY  $10^{-6}\%$  solid angle)

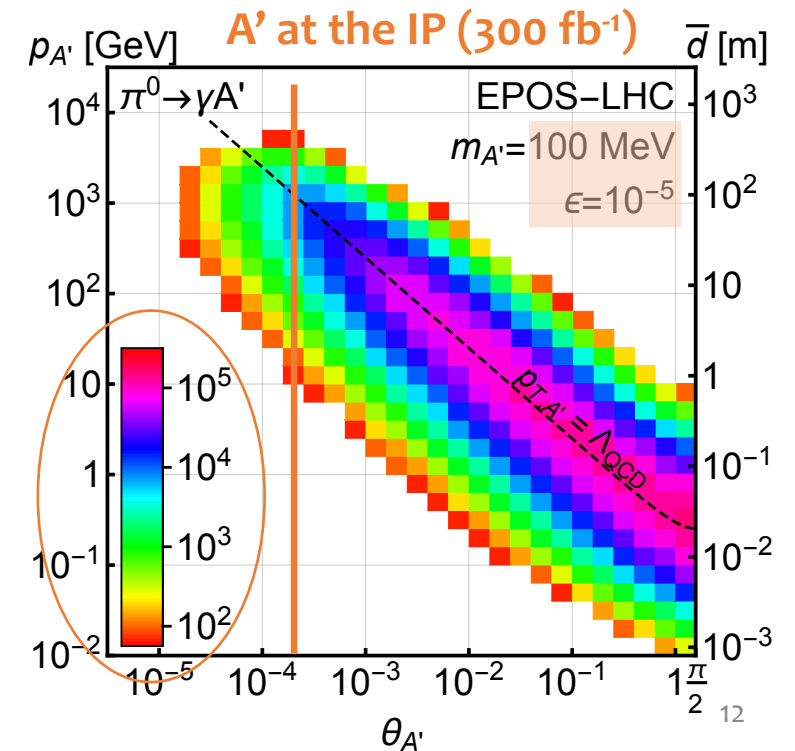
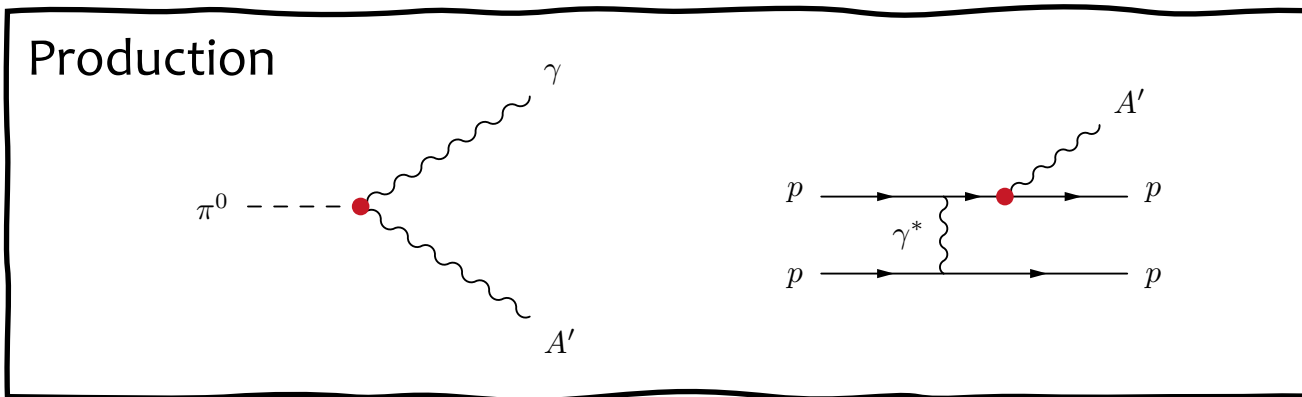
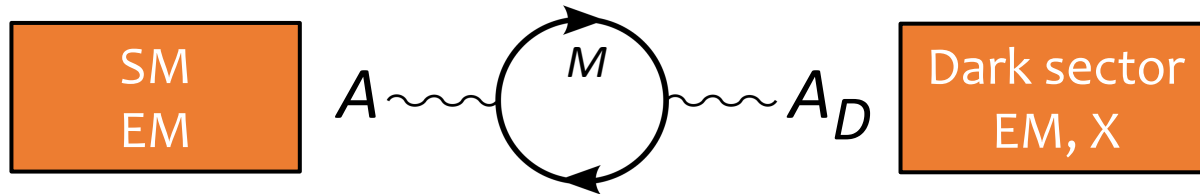
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# AN EXAMPLE PHYSICS CASE: DARK PHOTON



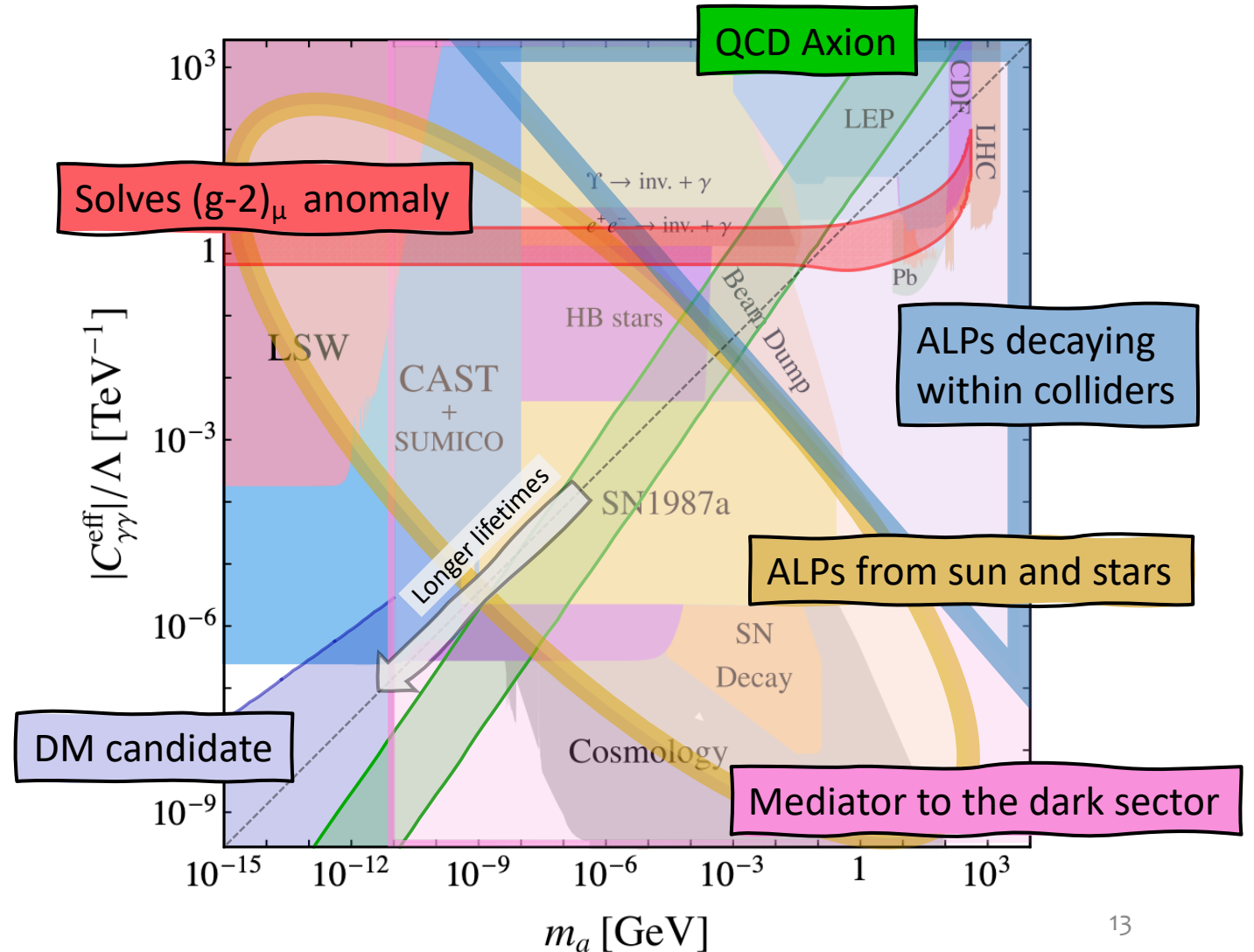
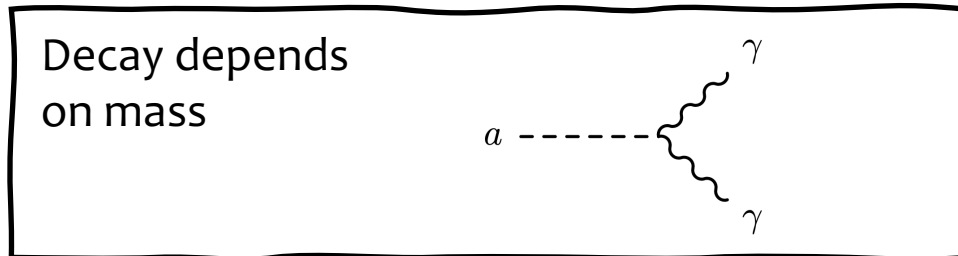
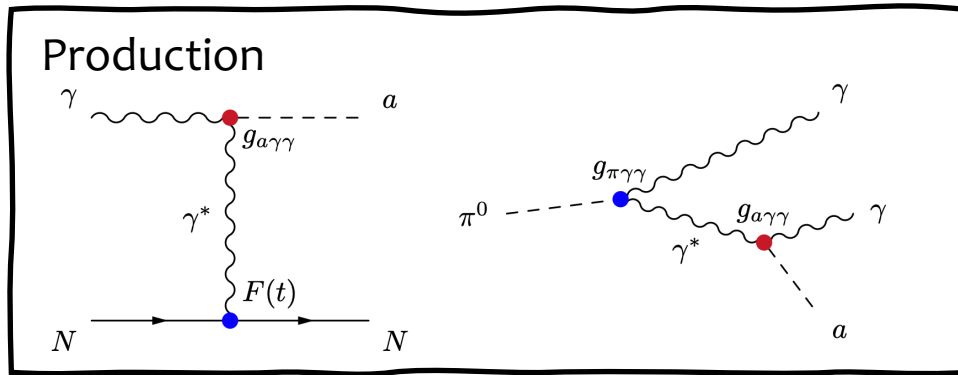
- New **massive** gauge boson in a dark sector with dark matter candidate  $X$
- Spin 1, **couples weakly to SM fermions** ( $\epsilon Q_f$  coupling, small  $\epsilon$ ) through mixing with the photon
  - Will be searched for via its **decay to an electron-positron pair**
- For  $m_{A'}=100$  MeV,  $\epsilon \sim 10^{-5}$  and  $E \sim \text{TeV}$ , can travel long distance before decay



# ANOTHER EXAMPLE: AXION-LIKE PARTICLES (ALPs)

Qualitatively different: “High-energy photon beam dump experiment”

- Pseudoscalar SM-singlets; can appear in theories with broken global symmetries
- Low mass particles with suppressed couplings to SM

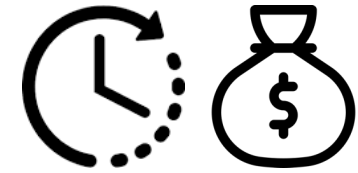


Thanks to Andrea Thamm for the figure!

THE ***FASER*** EXPERIMENT

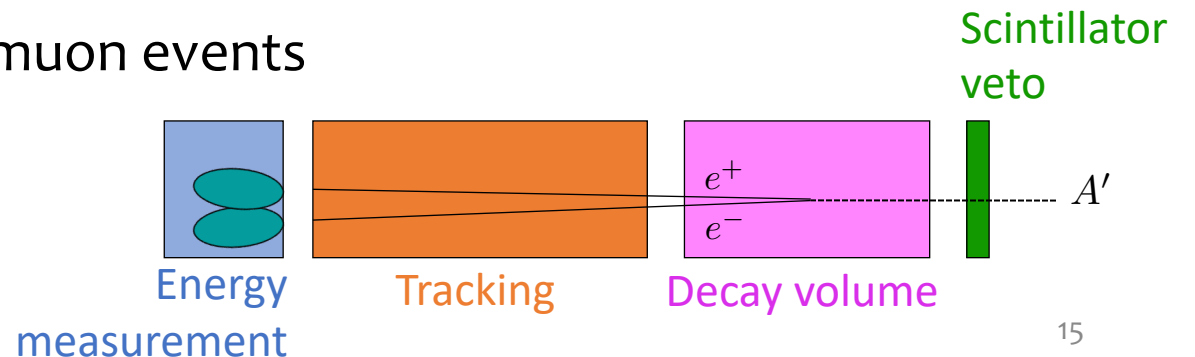
# DETECTOR CONCEPT

- **Drivers for choices:** Tight timeline between experiment approval and installation & the limited budget.
  - Detector that can be constructed and installed *quickly & cheaply*
  - Have tried to re-use existing detector components where possible
  - Aimed for a simple, robust detector (access difficult)
  - Tried to minimize the services to simplify the installation and operations

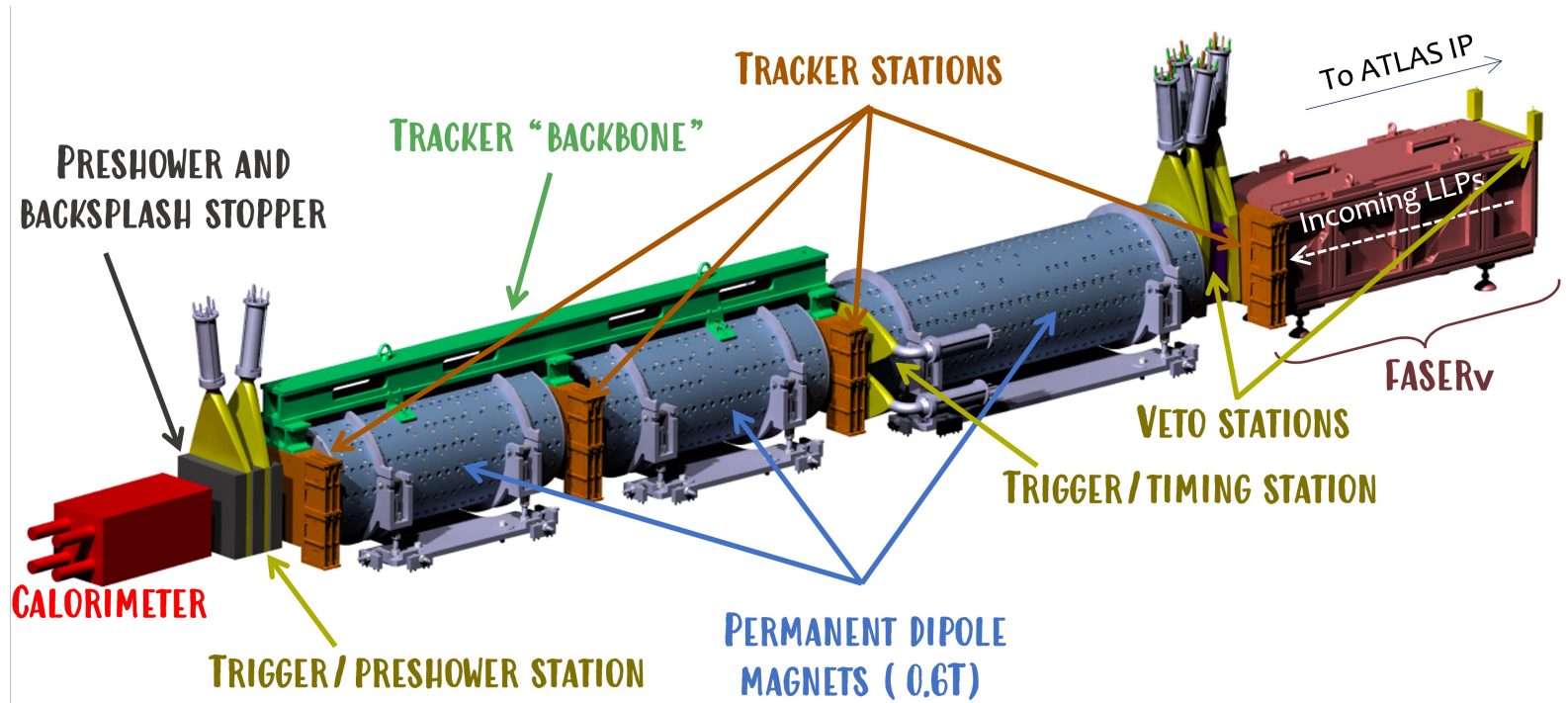


- **Many challenges of the large LHC experiments not there for FASER:**

- trigger rate  $O(500\text{Hz})$  – mostly single muon events
- low radiation
- low occupancy / event size



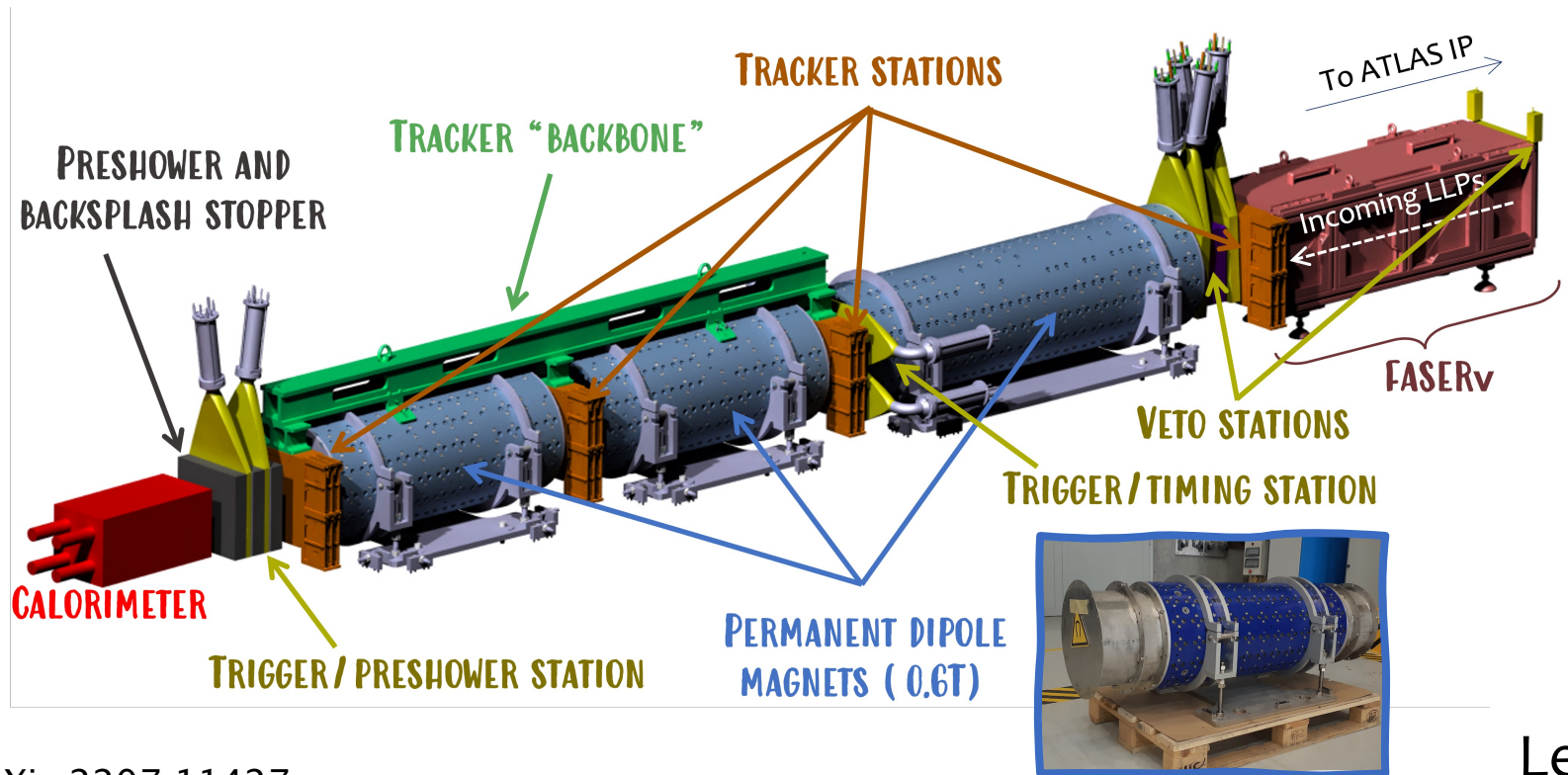
# **FASER** DETECTOR



Length: 7 m  
Aperture: 20 cm

Length of decay volume: 1.5 m

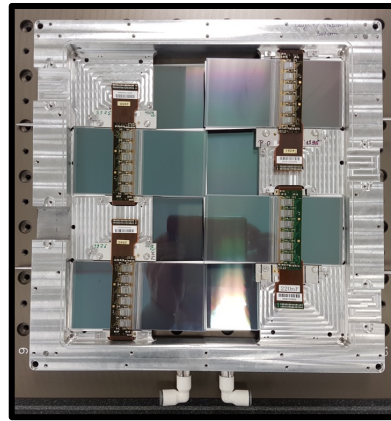
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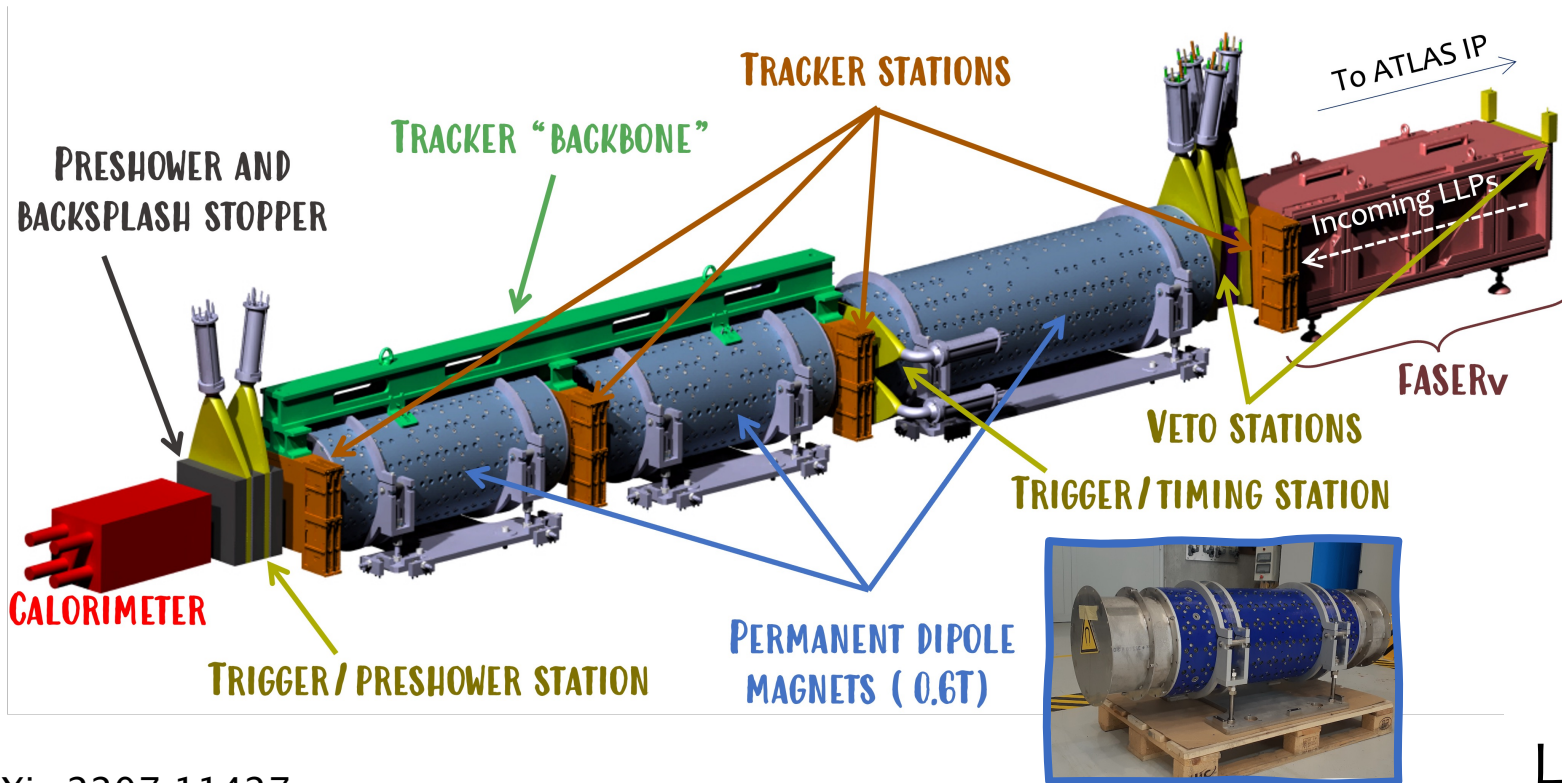
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# FASER DETECTOR



A tracker layer made of ATLAS SCT modules

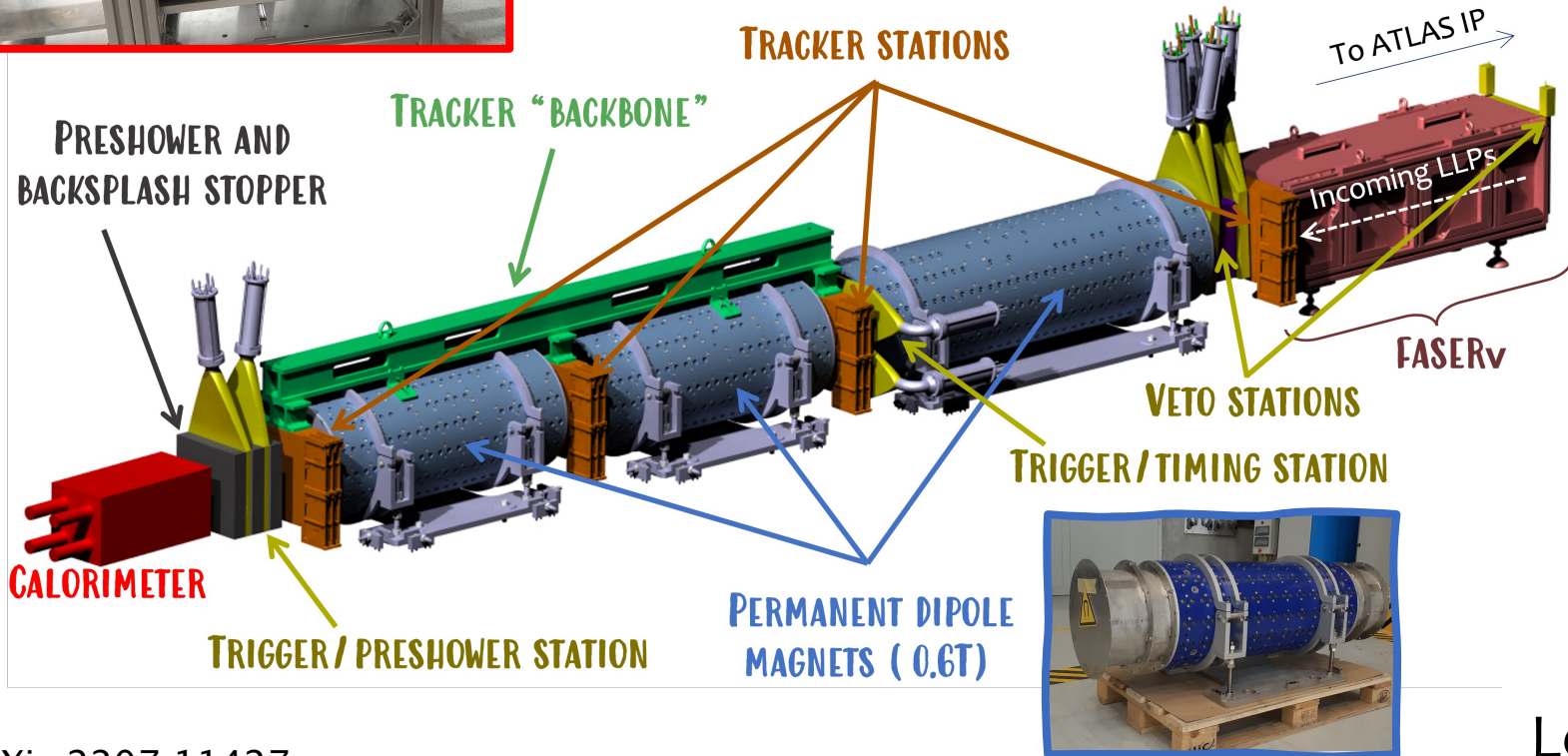
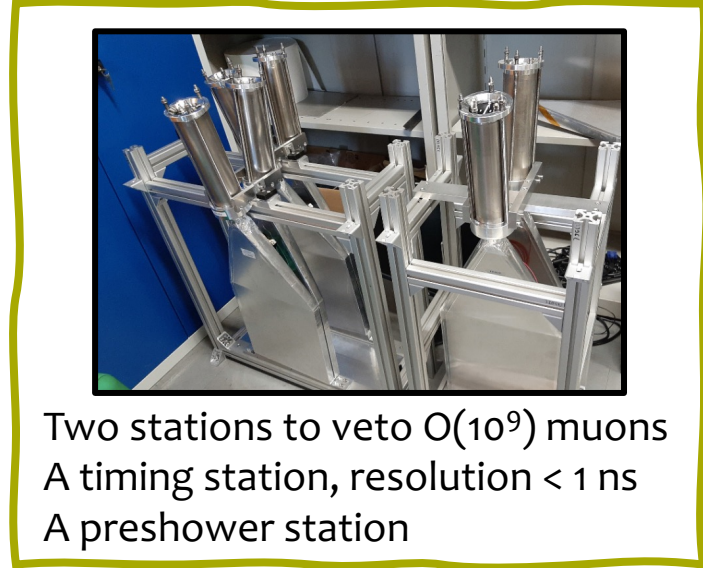
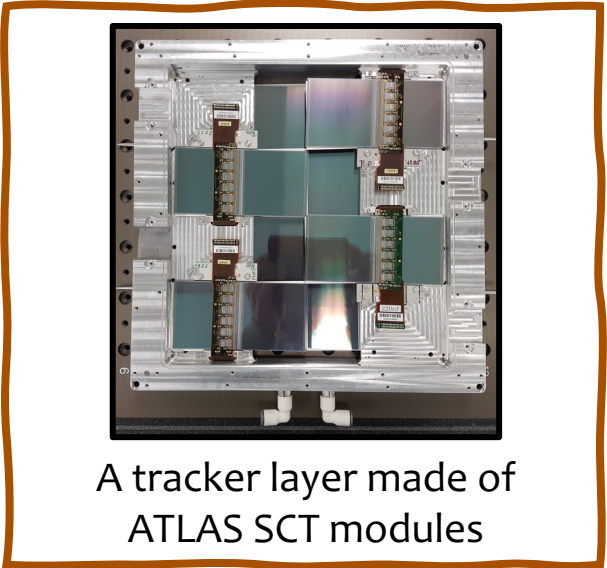
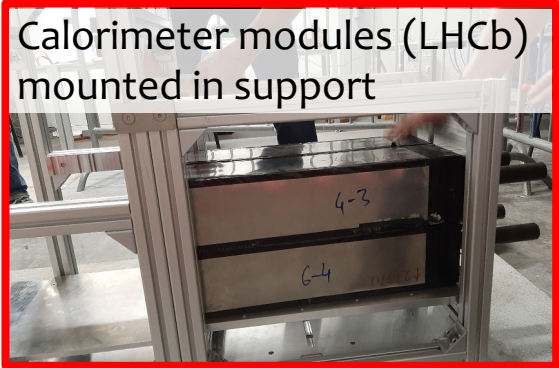


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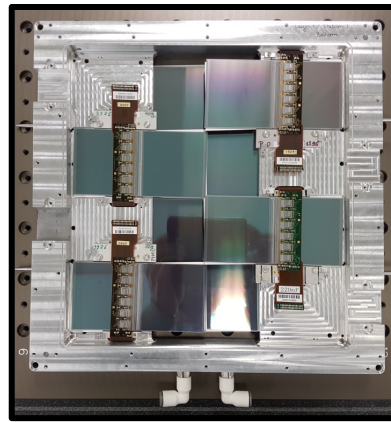
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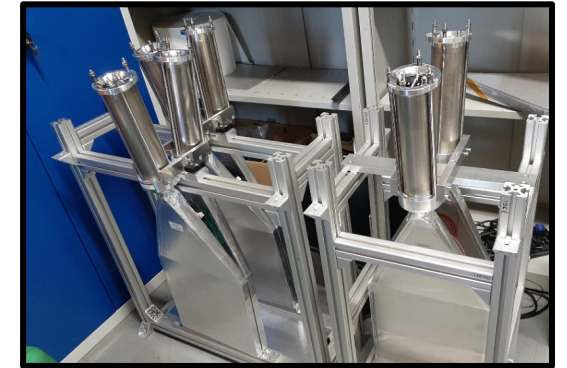
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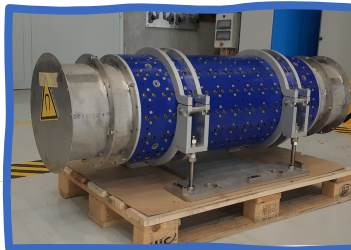
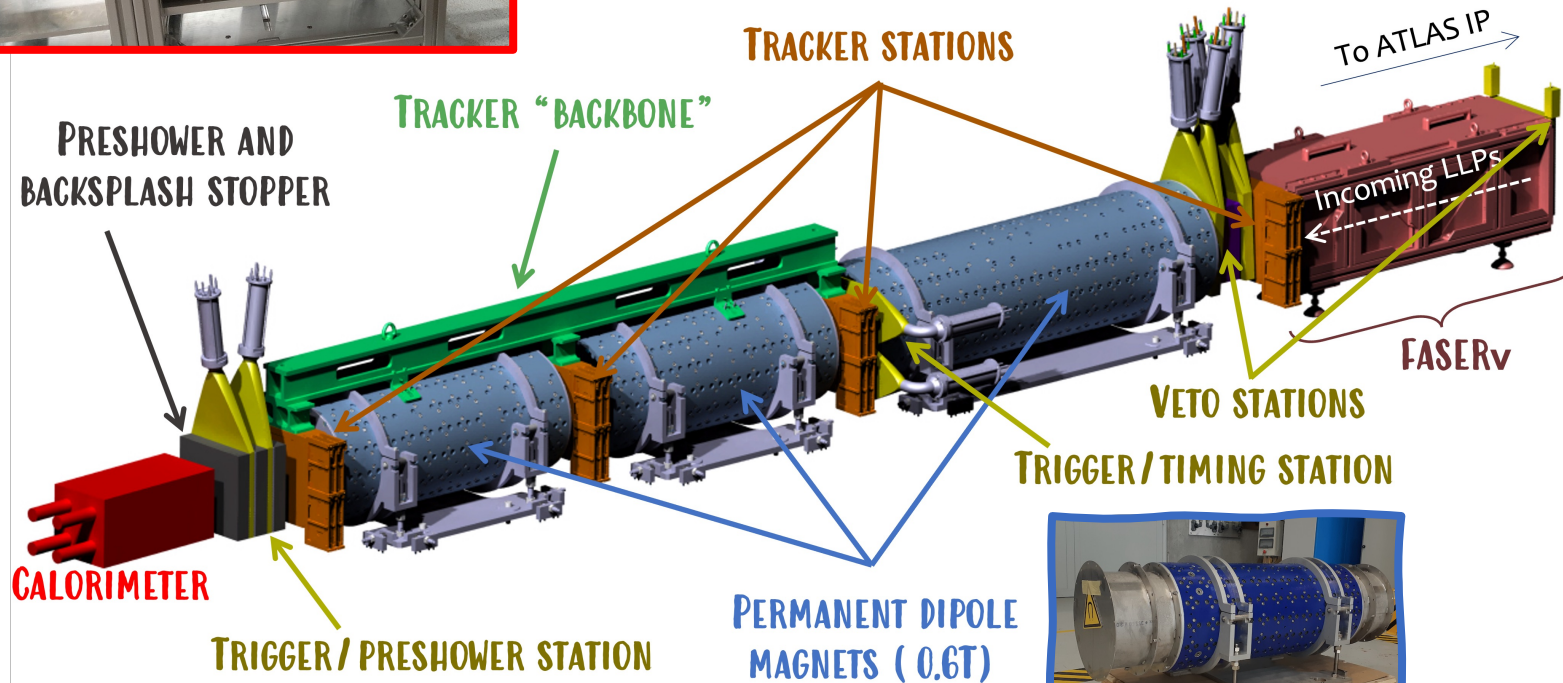
Calorimeter modules (LHCb) mounted in support



A tracker layer made of ATLAS SCT modules



Two stations to veto  $O(10^9)$  muons  
A timing station, resolution  $< 1$  ns  
A preshower station

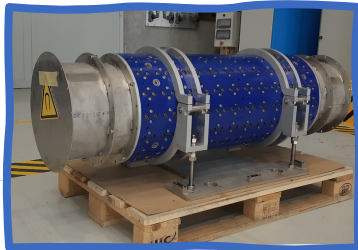
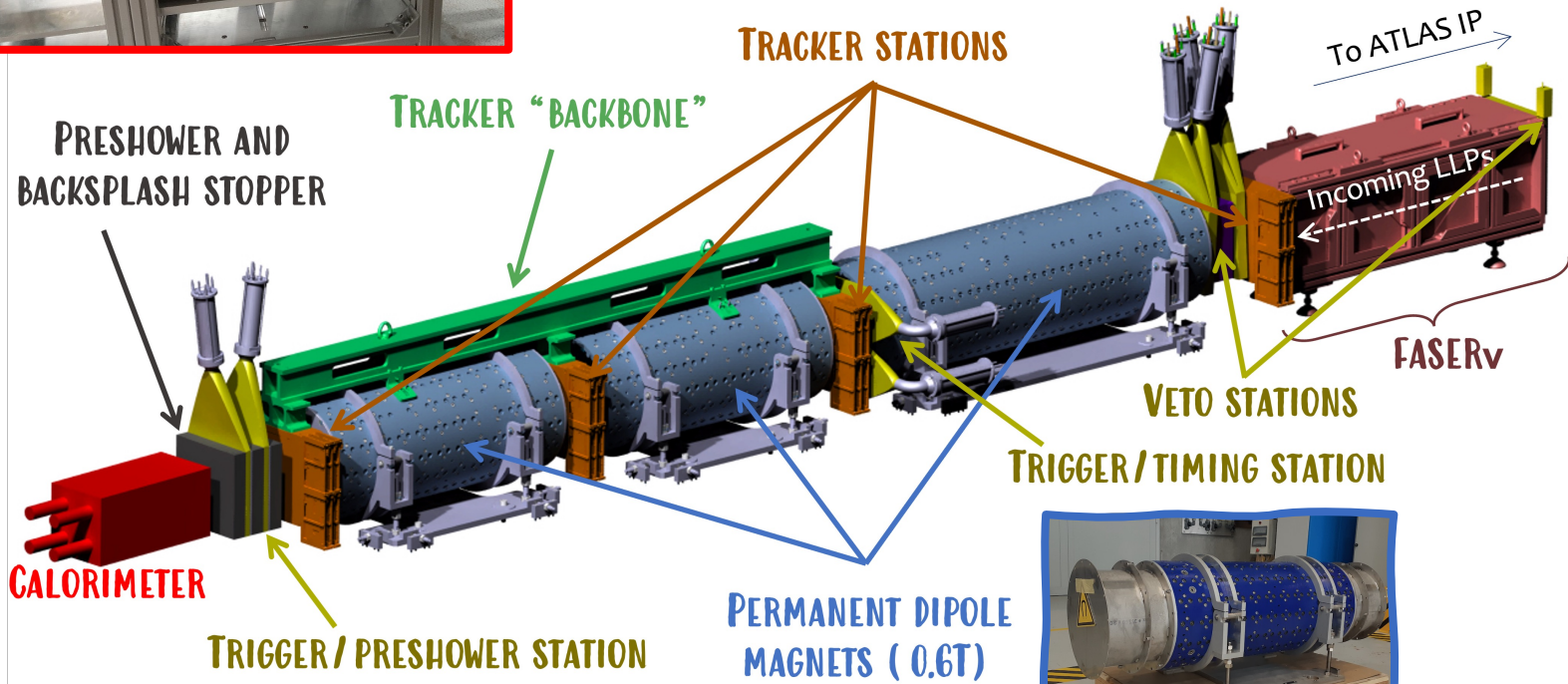
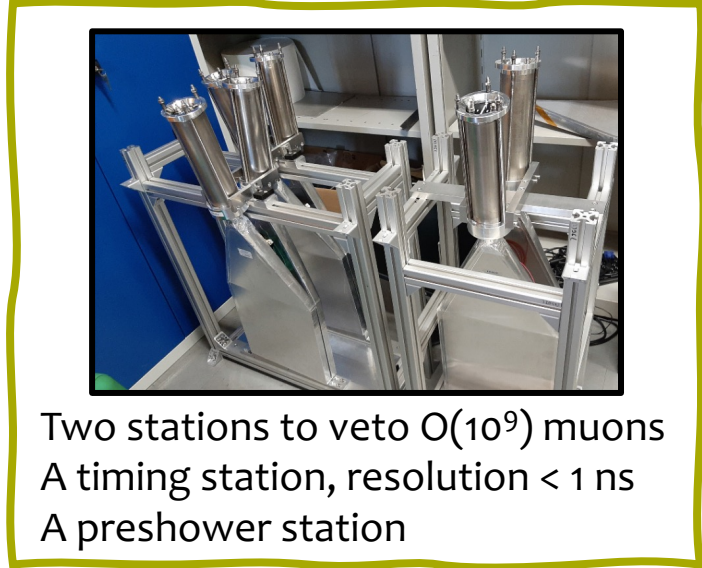
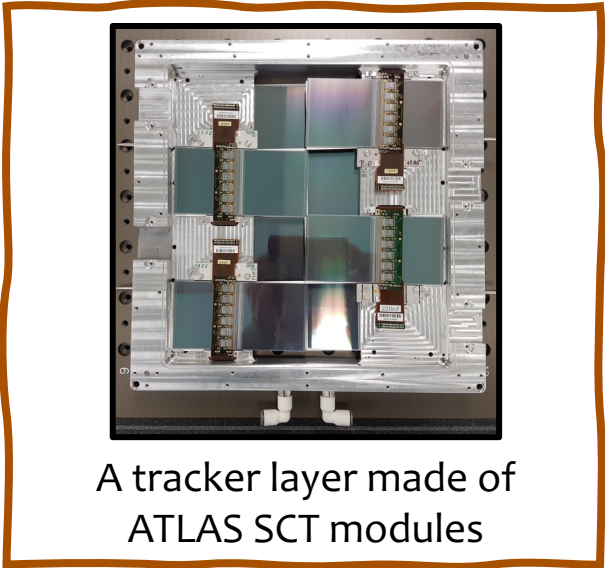
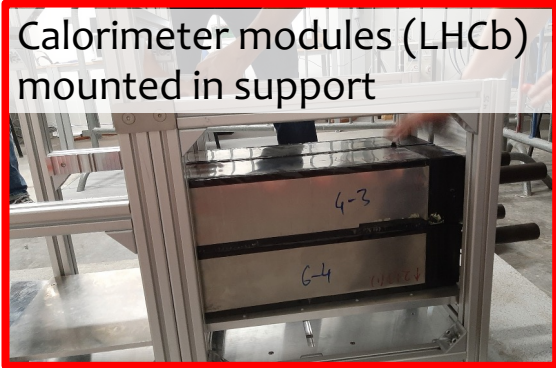


Triggers from scintillators and calorimeter  
Expected trigger rate  $< 1$  kHz, dominated by muons from the IP

Length: 7 m  
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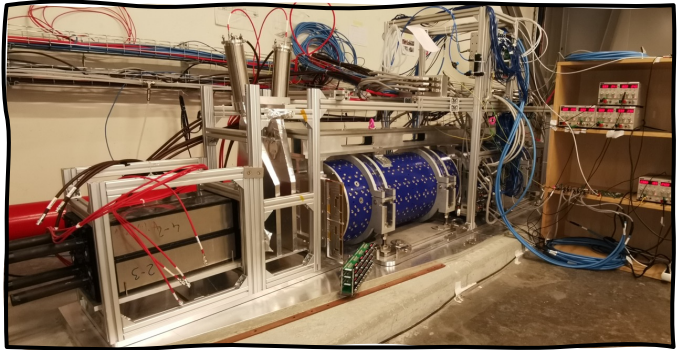


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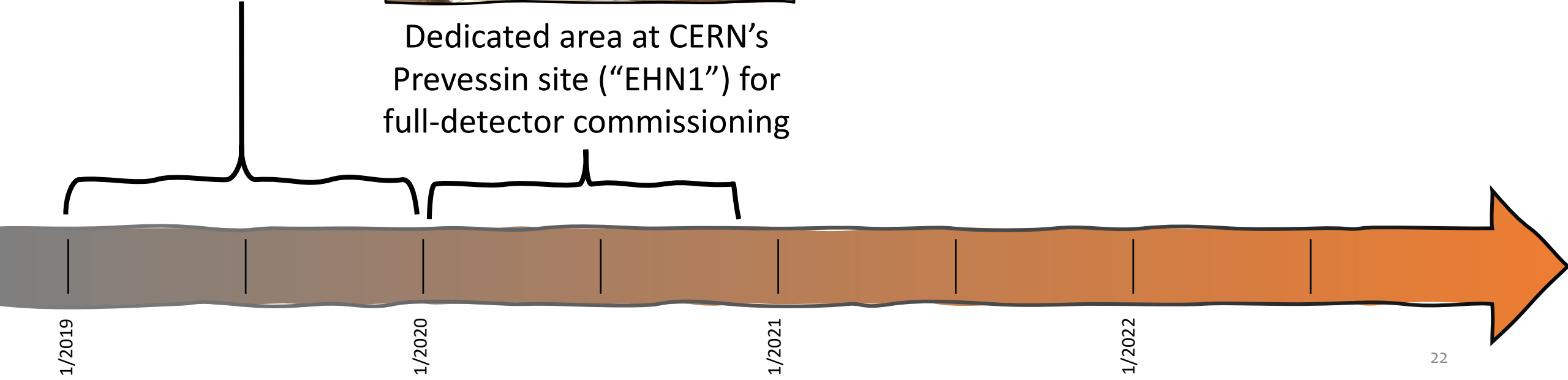
# COMMISSIONING

## TIMELINE

Dedicated labs at CERN and UniGe for individual component testing



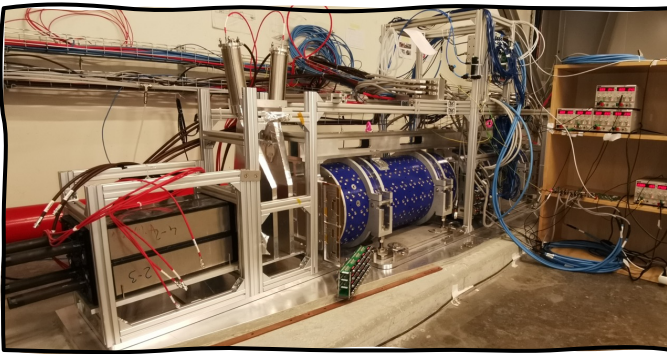
Dedicated area at CERN's Prevezin site ("EHN1") for full-detector commissioning



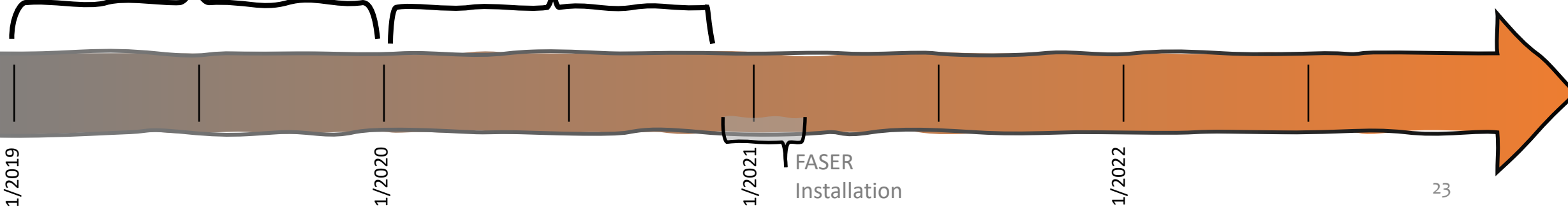


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CMU 2t

CMU 2t

ATTENTION  
HAUTE TENSION  
DANGER



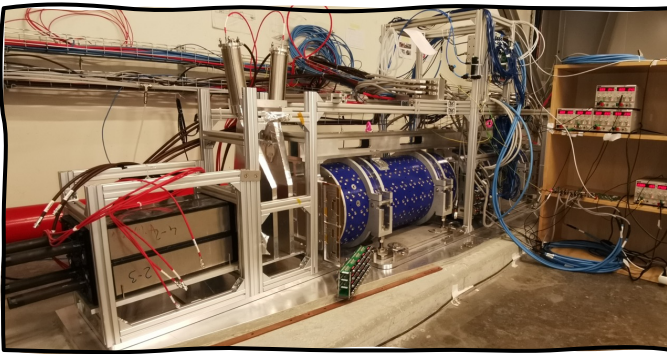
FASER

FASER

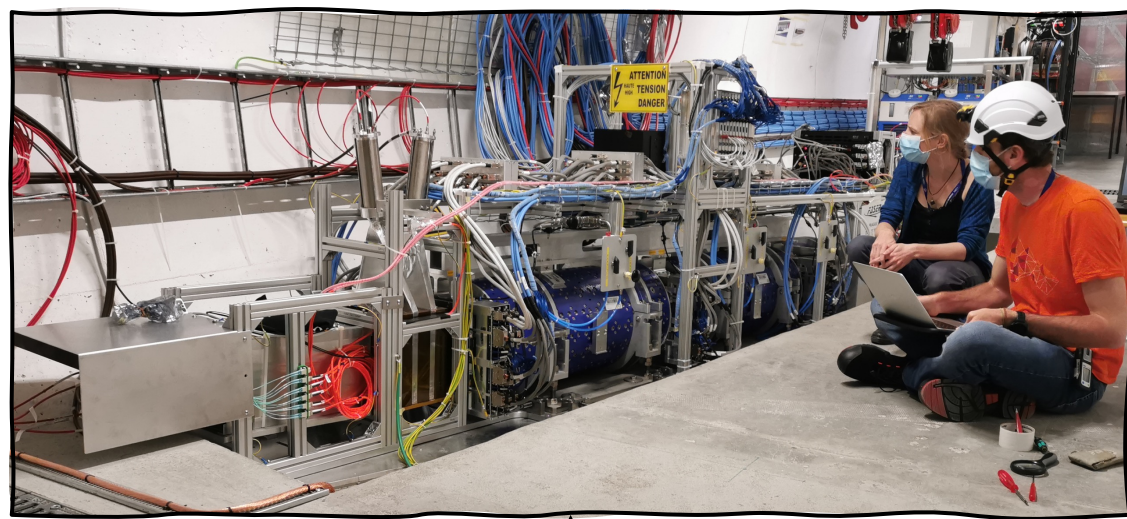


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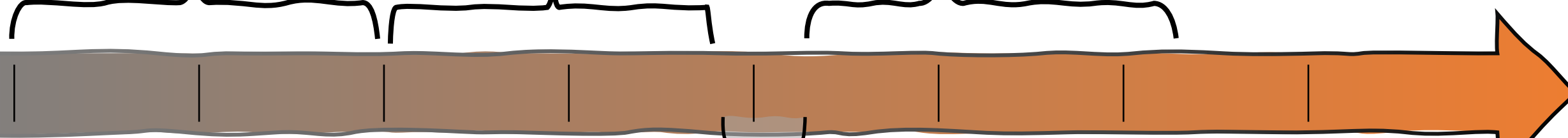
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Extensive in-situ  
commissioning



1/2019

1/2020

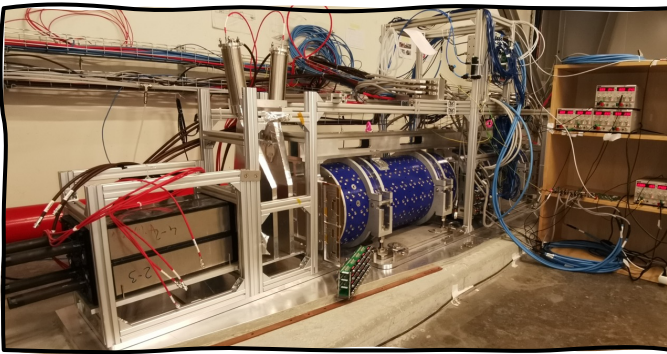
1/2021

FASER  
Installation

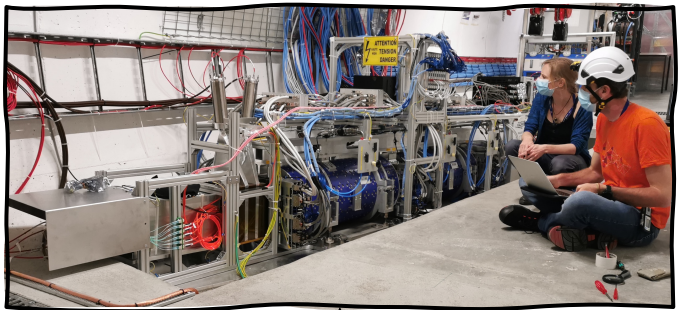
1/2022

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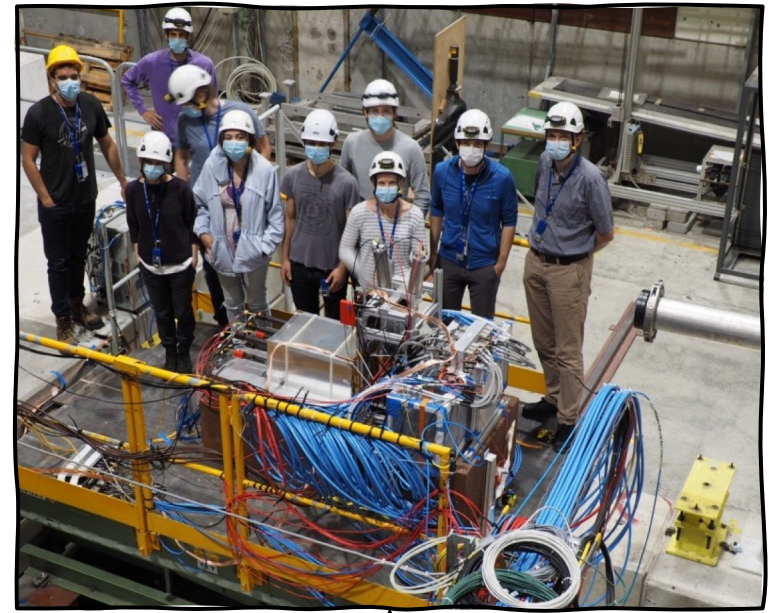
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Extensive in-situ  
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Testbeam

1/2019

1/2020

1/2021

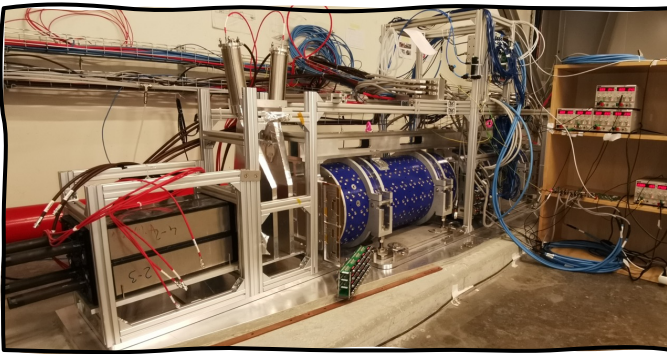
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1/2022

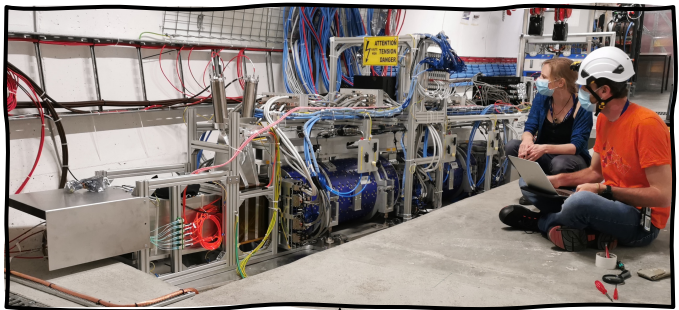


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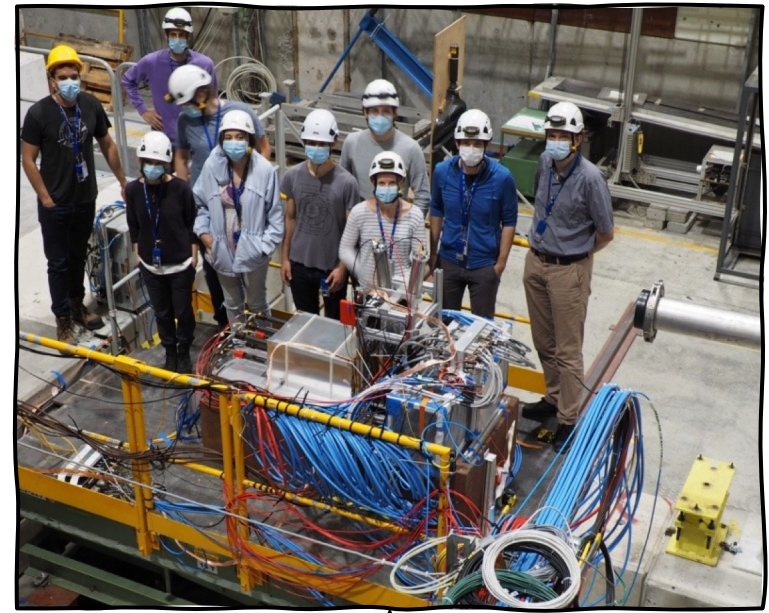
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1/2019

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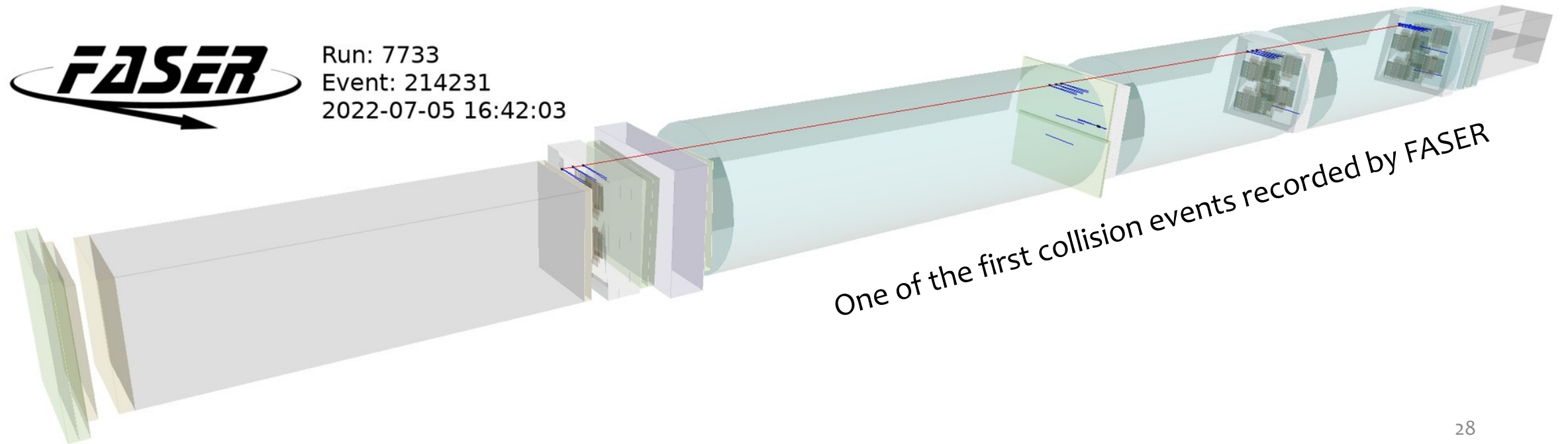
1/2022

# RUN3 DATA TAKING

- Have successfully collected 13.6 TeV data
- Initial detector performance (being) assessed
- Offline and analysis software up and running
- **First analysis results produced for spring 2023 conferences**



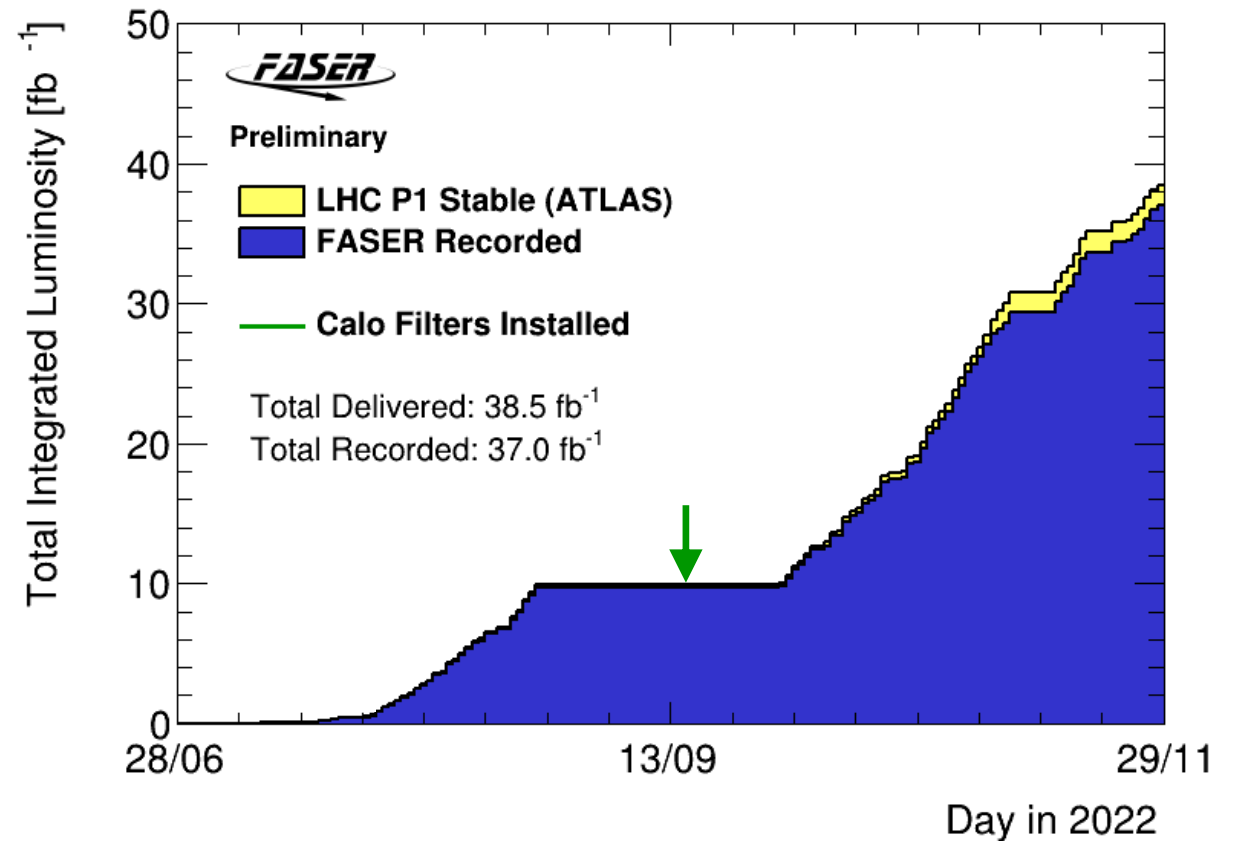
Run: 7733  
Event: 214231  
2022-07-05 16:42:03



One of the first collision events recorded by FASER

# FASER OPERATIONS

- Recorded 96.1% of delivered lumi
- Dead-time of  $< 2\%$
- About 2% data loss due to two specific operational issues
- Calorimeter gain optimised for
  - low E ( $< 300$  GeV) initially
  - high E later in the run



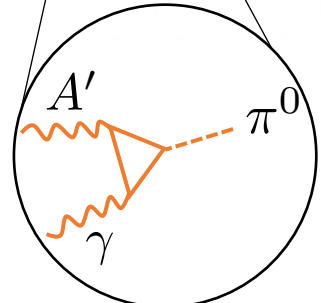
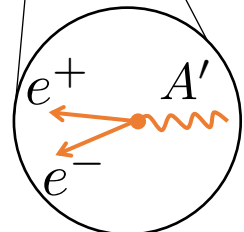
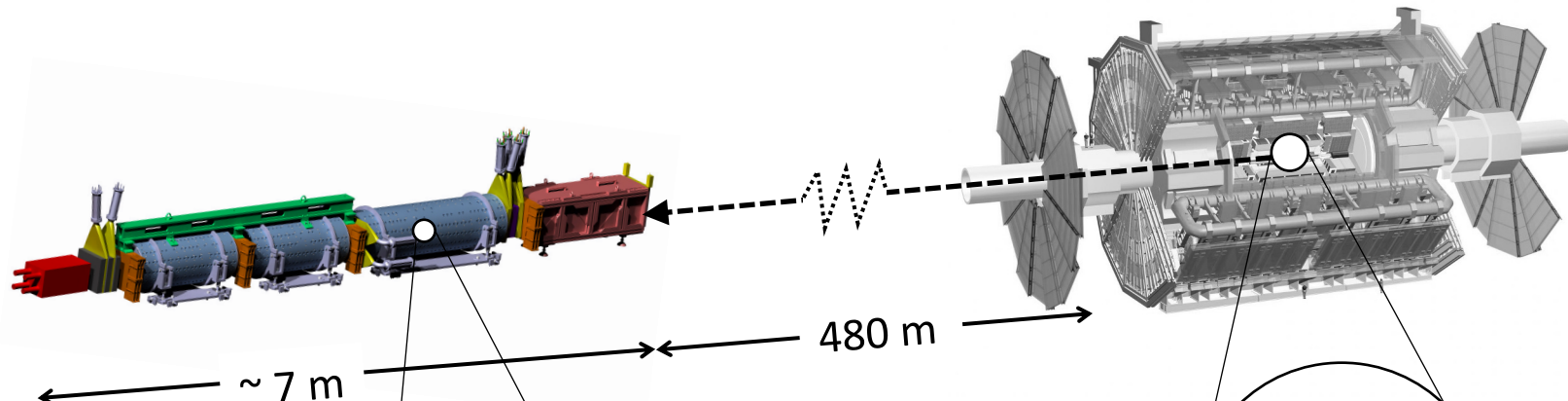
Analyses produced use 27/fb or 35.4/fb

# FIRST DARK PHOTON SEARCH WITH *FASER*

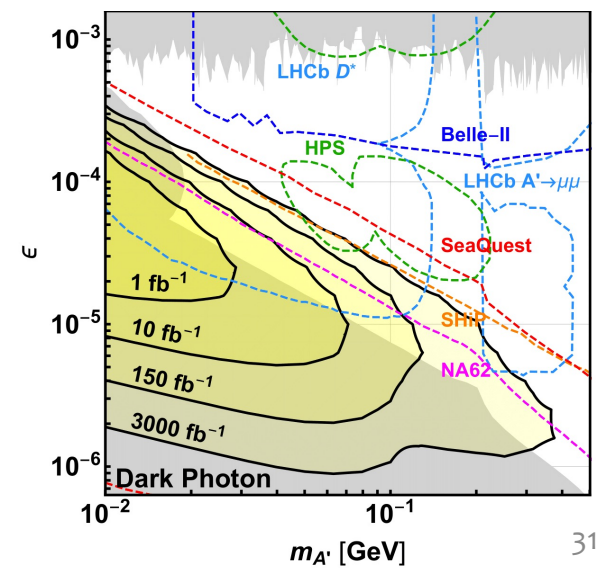
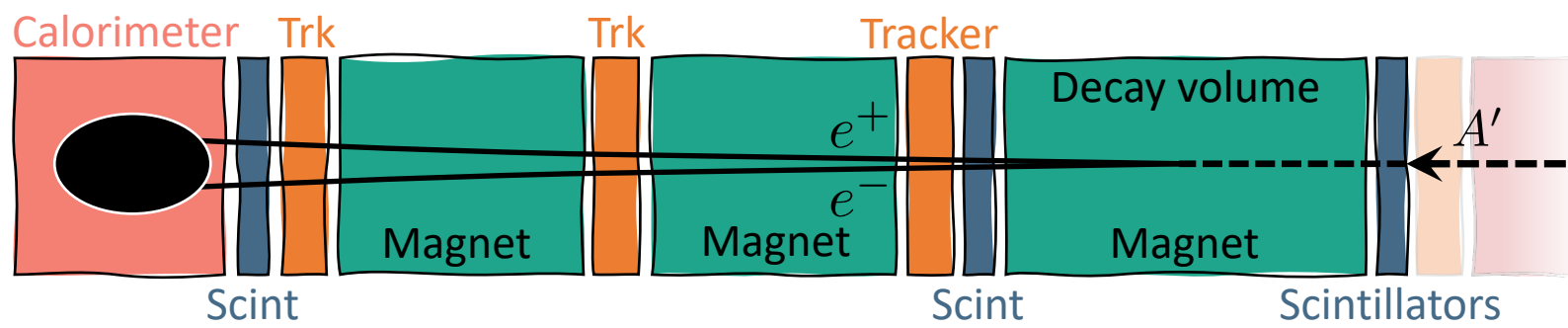


# A KEY SIGNATURE

# Dark photon ( $A'$ )



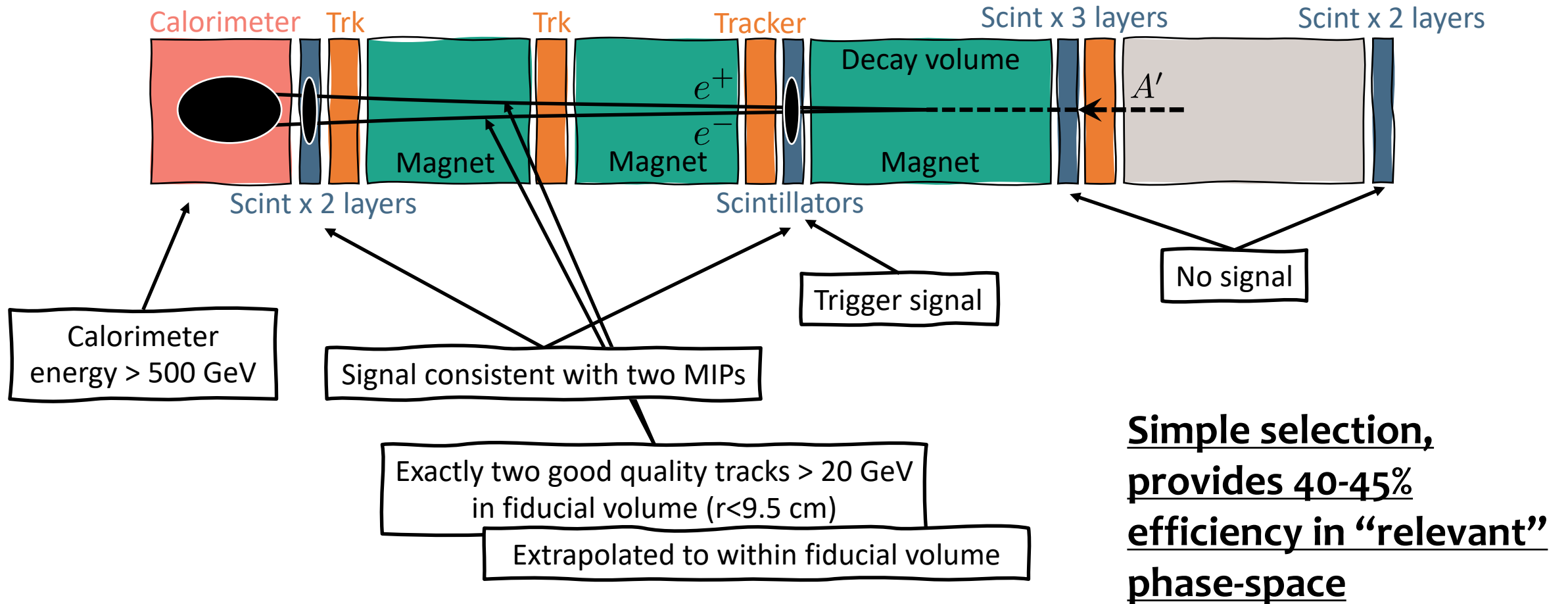
Decay products collimated requirements for magnetic field & high-resolution tracker



Assuming 3 signal events and no backgrounds

# SIGNAL SELECTION

Dark photon ( $A'$ )

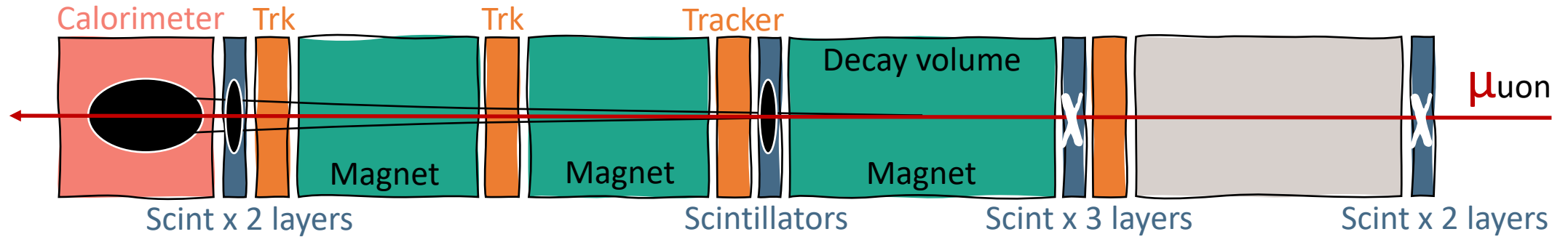




# BACKGROUNDS

Dark photon ( $A'$ )

## 1. Muons that are not vetoed by any of the 5 scintillator layers



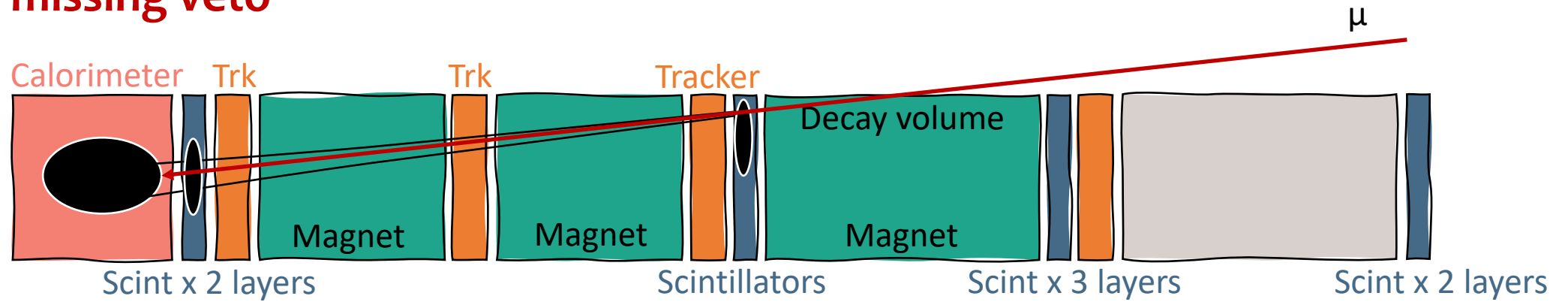
Scintillator	Efficiency
NuVeto-0	0.9999805(5)
NuVeto-1	0.9999810(5)
Veto-0	0.9999985(1)
Veto-1	0.9999984(1)
Veto-2	0.9999986(1)

- Total scintillator inefficiency  $< 10^{20}$
- Even with  $O(10^8)$  muons (2022):
  - **zero** probability that one won't be vetoed

# BACKGROUNDS

Dark photon ( $A'$ )

## 2. Muons missing veto



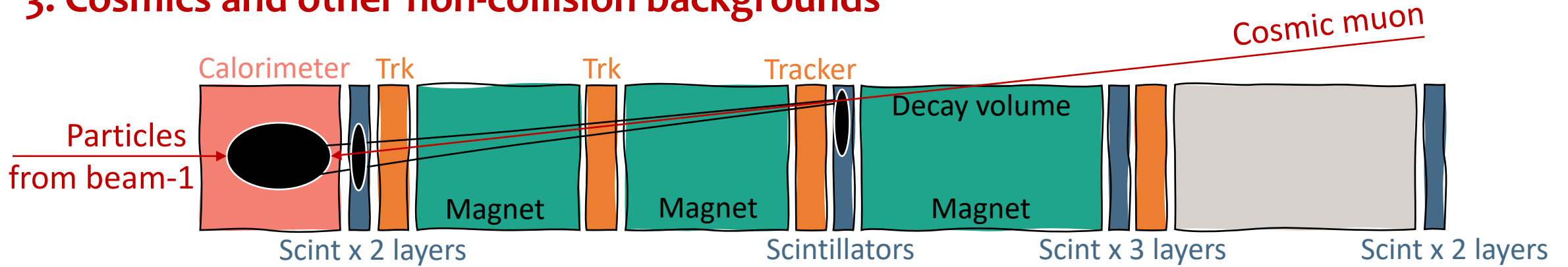
- Evaluated using MC
- **Zero** background



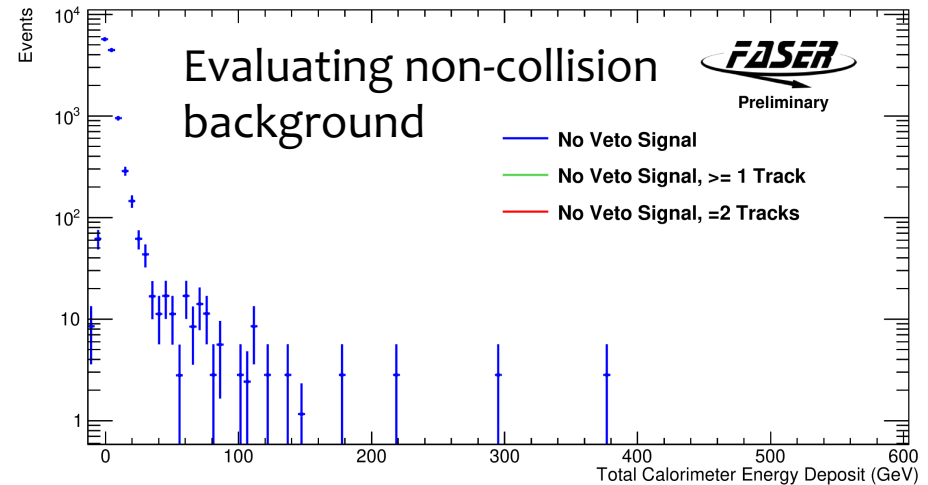
# BACKGROUNDS

## Dark photon ( $A'$ )

### 3. Cosmics and other non-collision backgrounds



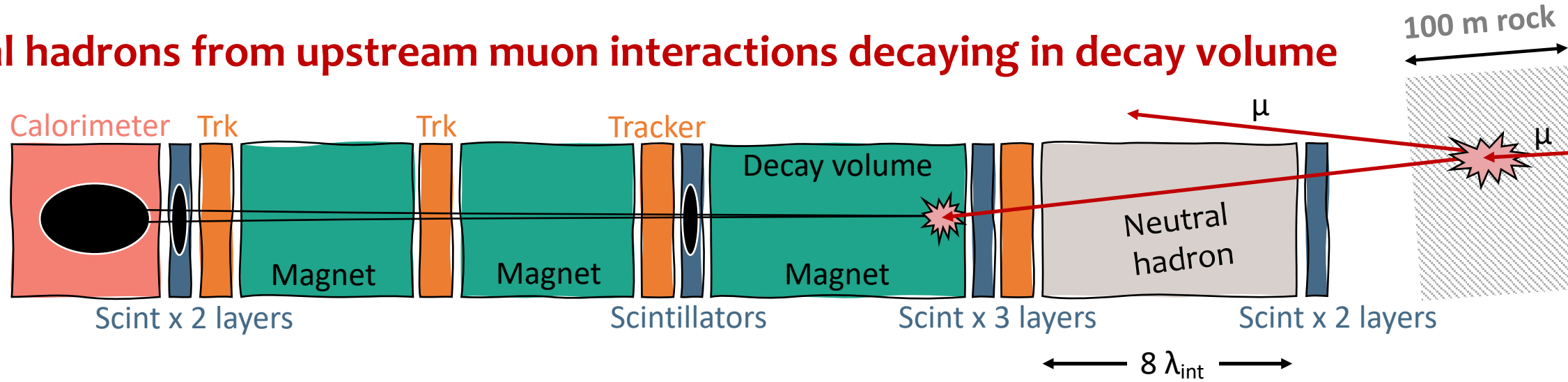
- Evaluated in non-colliding bunches and runs without beam
- **Zero** background



# BACKGROUNDS

Dark photon ( $A'$ )

## 4. Neutral hadrons from upstream muon interactions decaying in decay volume



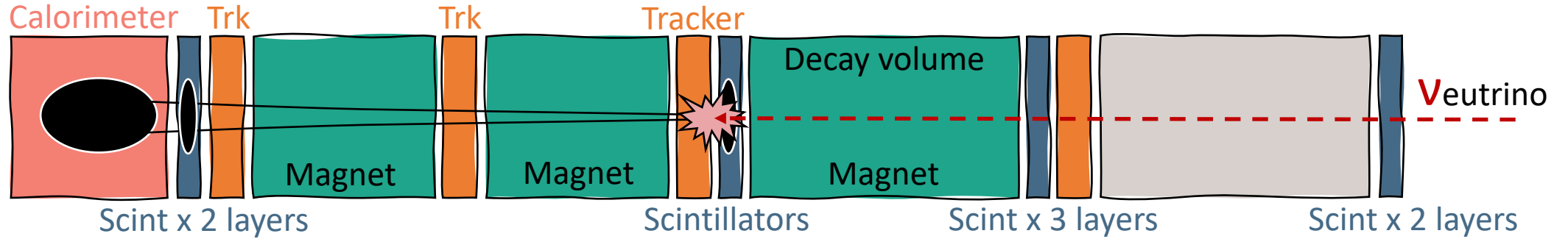
- Evaluated using three-track events, extrapolating from low energy to high energy and from loose to nominal veto requirements

- Estimated background:  $(2.2 \pm 3.1) 10^{-4}$  events

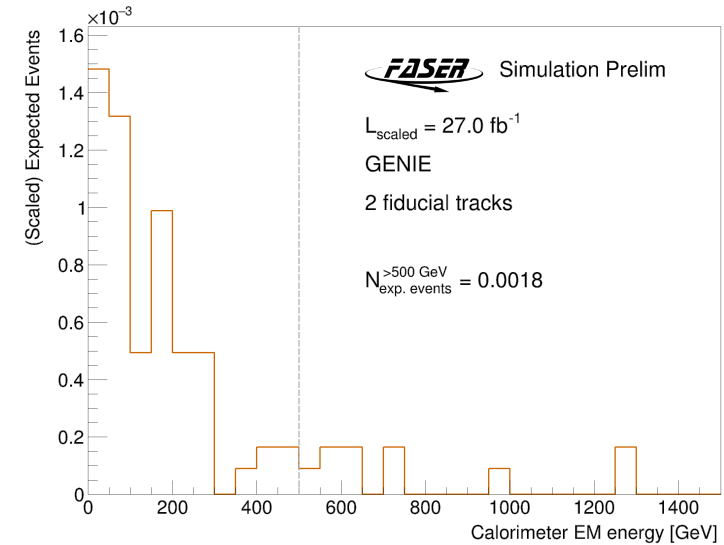
# BACKGROUNDS

Dark photon ( $A'$ )

## 5. Neutrino interactions

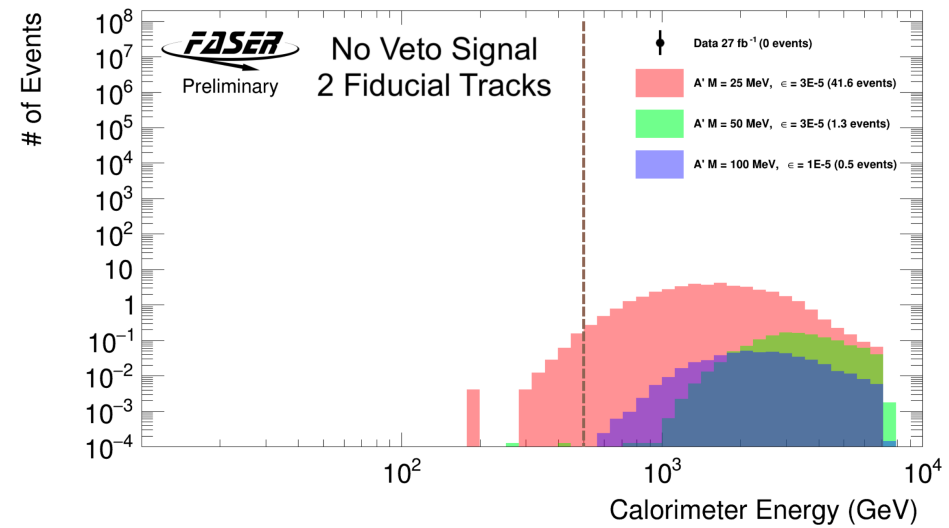
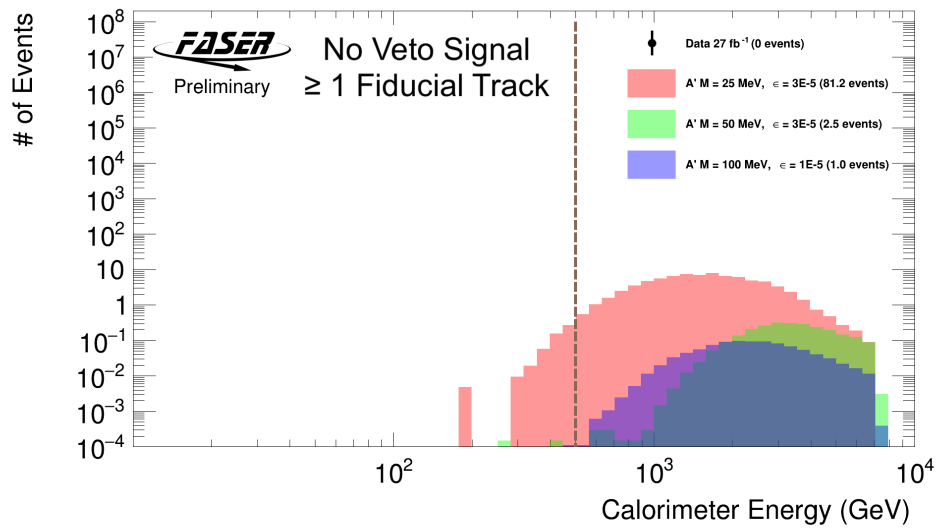
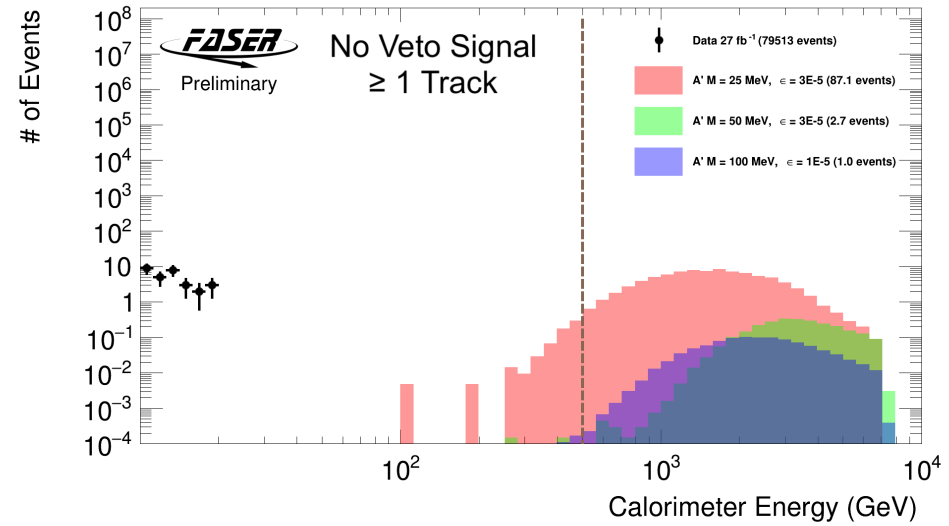
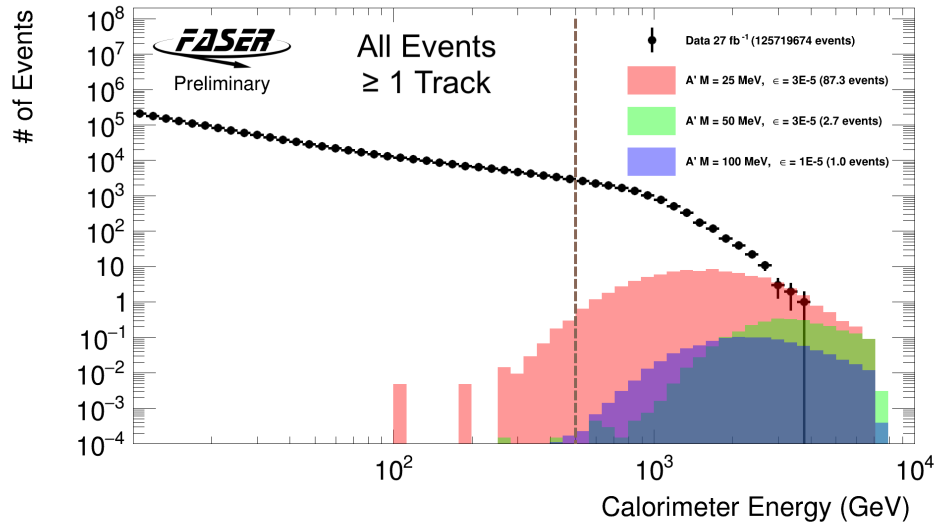


- Evaluated using MC
  - GENIE generator
  - 300/ab
- Estimated background:  
 **$(1.8 \pm 2.4)10^{-3}$  events**



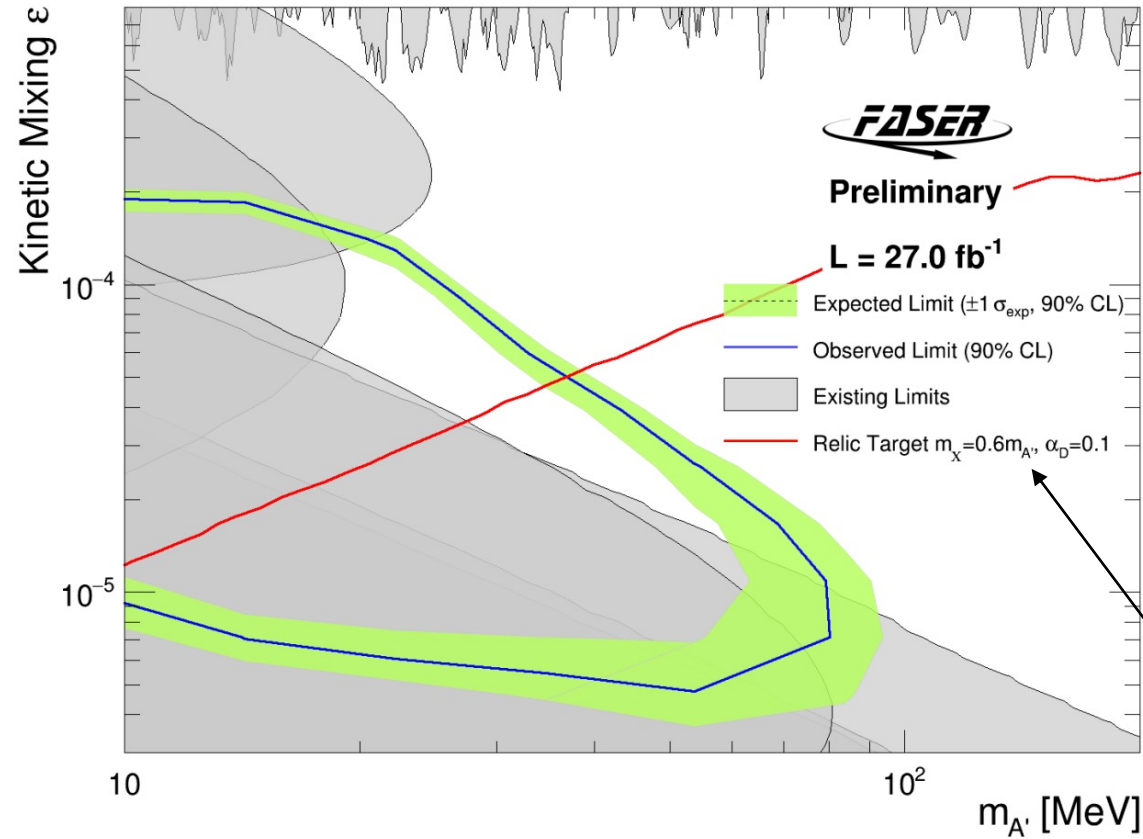
# LOOKING IN DATA

Total background:  $(0.0020 \pm 0.0024)$  events  
**No events seen in unblinded signal region**



# EXCLUSION REACH

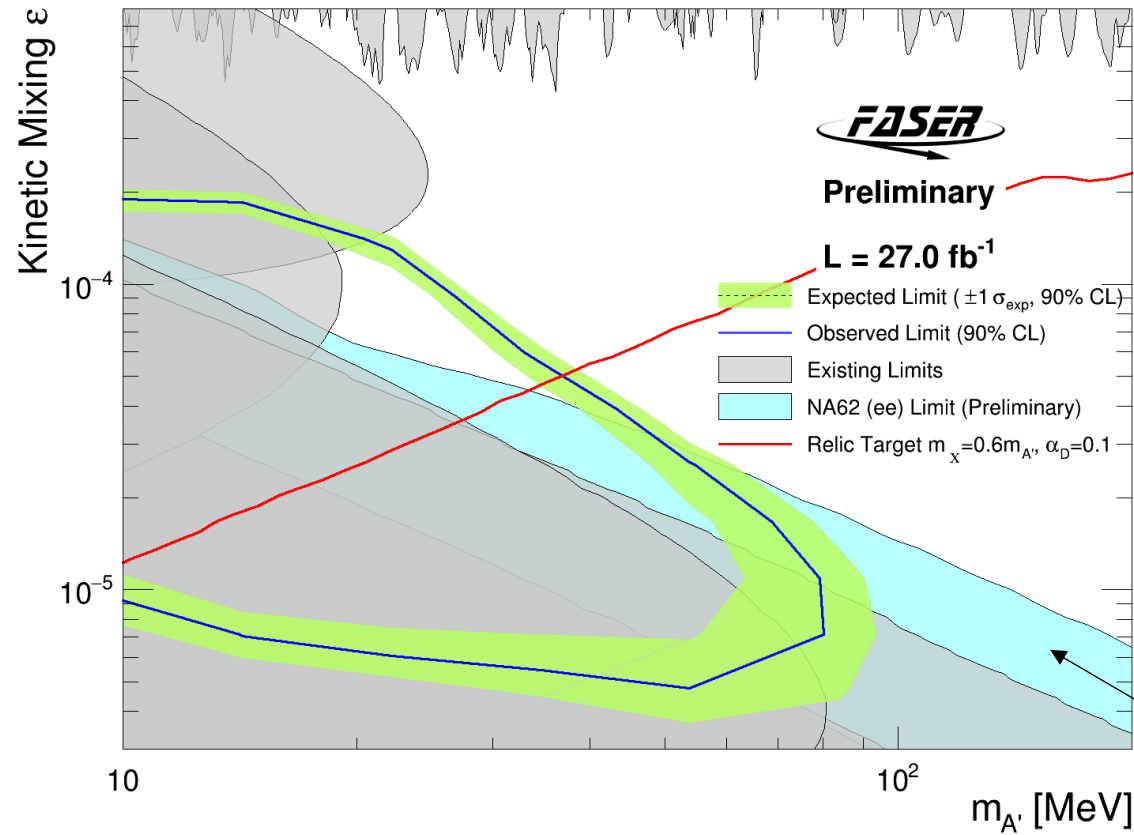
On previously unexplored phase-space



- Parameters for which DM  $\chi$  annihilates via  $\chi\chi \rightarrow A' \rightarrow ff$
- Model dependent *line*, but region favoured by DM relic density

# EXCLUSION REACH

On previously unexplored phase-space



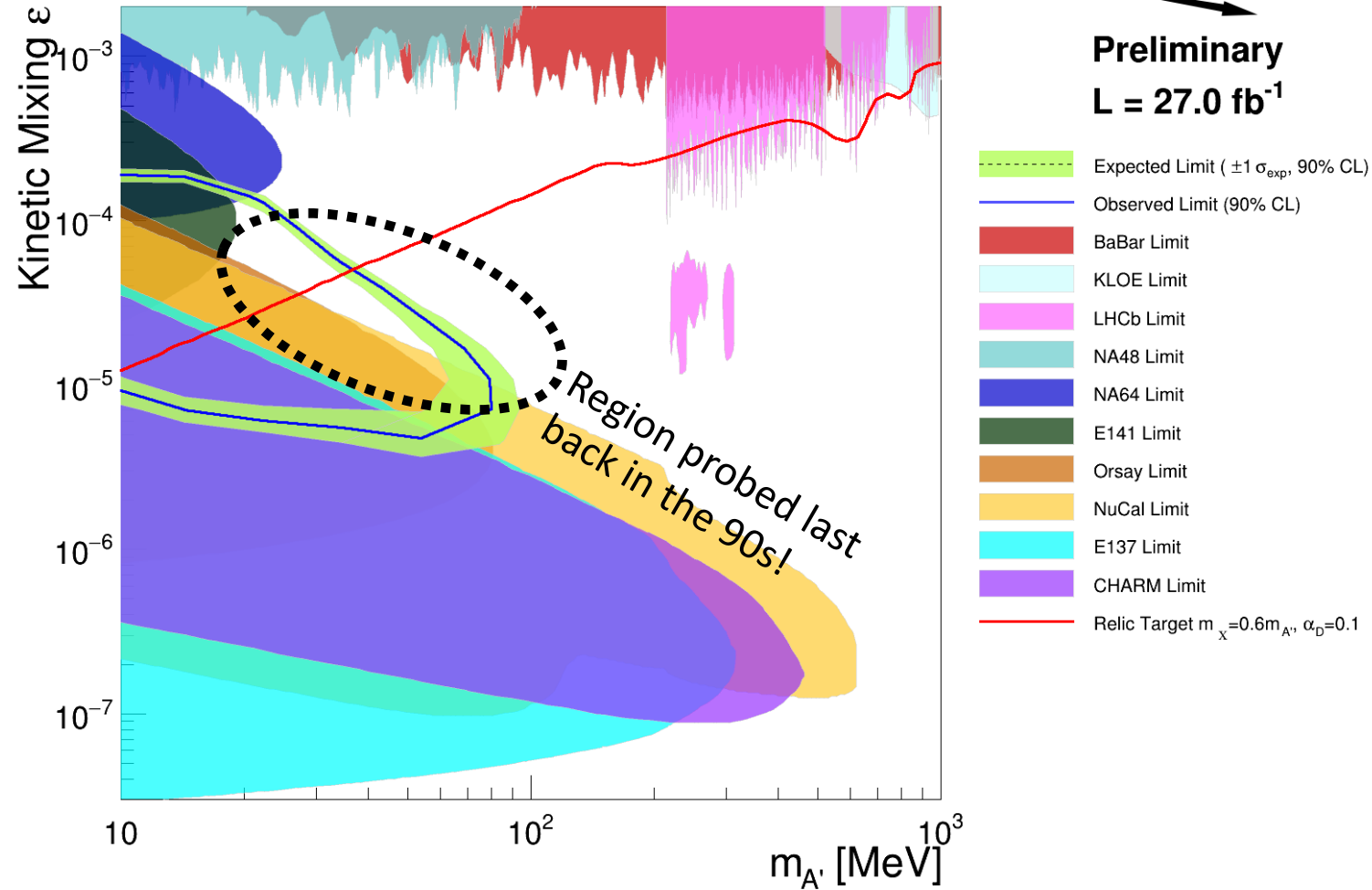
For La Thuile 23, NA62 too announced a preliminary result, complementary to the FASER one

# EXCLUSION REACH

On previously unexplored phase-space



**Preliminary**  
**L = 27.0 fb<sup>-1</sup>**





**FÄSER**



The logo consists of the word "FÄSER" in a bold, italicized, sans-serif font. The letters are black and set against a thick, black, horizontally-oriented oval background. Below the oval, there is a decorative pattern of small, black, stylized 'v' characters arranged in a curved, wave-like shape that follows the bottom edge of the oval.

# HUGE FLUX OF HIGH-ENERGY NEUTRINOS

- Why not exploit FASER to also measure properties of neutrinos at the highest man-made energies ever recorded!

## A bit of history



Experiments to study collider neutrinos have been proposed since the 80s, e.g.:

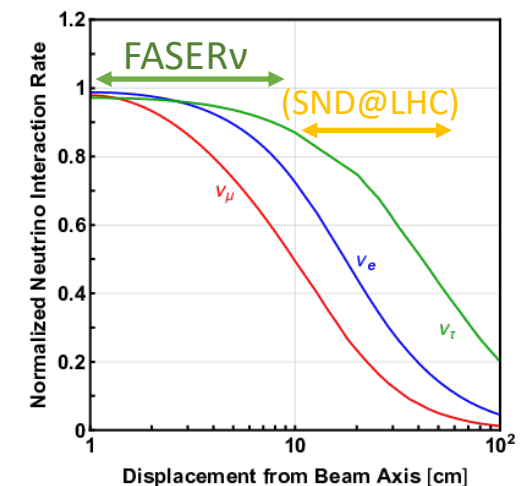
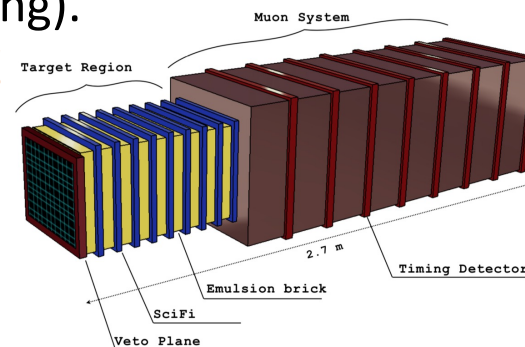
- A. De Rujula and R. Ruckl, “Neutrino and muon physics in the collider mode of future accelerators” ECFA-CERN Workshop on large hadron collider in the LEP tunnel, pp. 571–596, **1984**.
- Klaus Winter, “Observing tau neutrinos at the LHC”, LHC workshop, **1990**.

Other recent concrete experiment proposals include XSEN and SND@LHC.

SND@LHC approved for Run 3 data taking (and is running).

**FASERv was the first such experiment to be approved!**

More on FASERv in what follows...

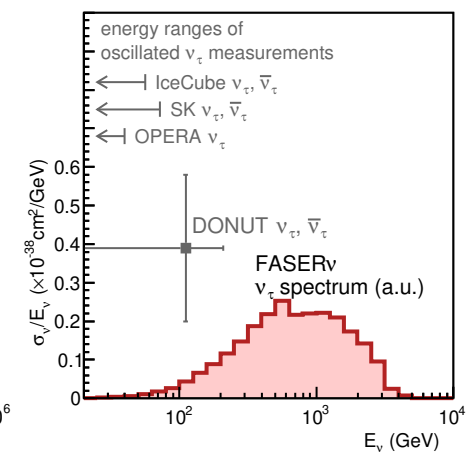
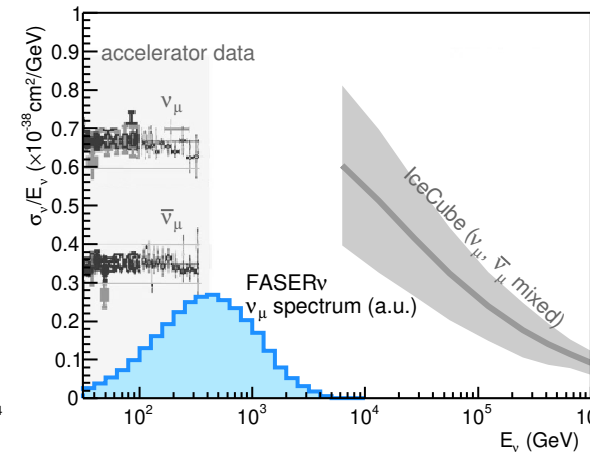
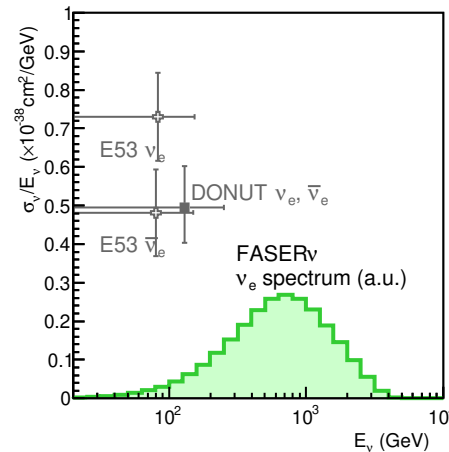


# HUGE FLUX OF HIGH-ENERGY NEUTRINOS

- Why not exploit FASER to also measure properties of neutrinos at the highest man-made energies ever recorded!

- **Expected spectra:** complementary to existing experiments
- **Expected cross section reach:** extends current measurements already with 150/fb
- **Rich QCD physics explorations**
- **Refine simulations that currently vary greatly** (EPOS-LHC, QGSJET, DPMJET, SIBYLL, PYTHIA, ...)

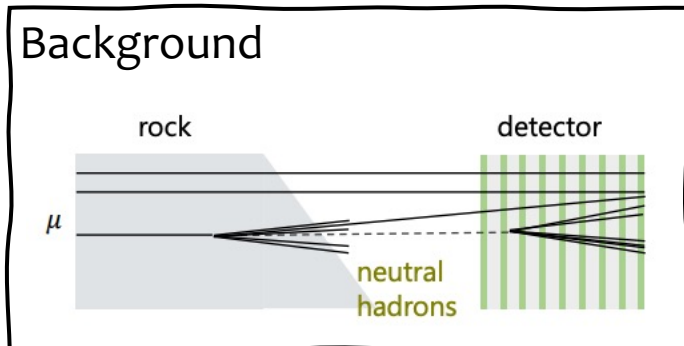
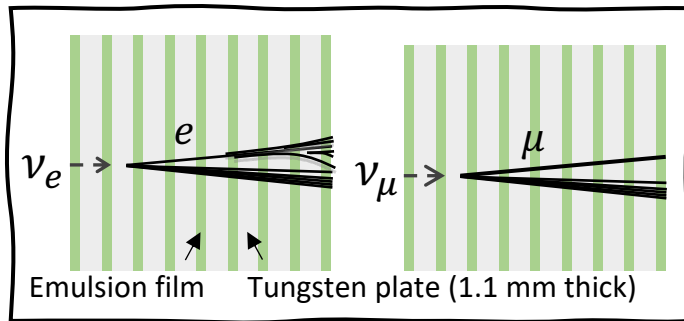
150/fb @14TeV	$\nu_e$	$\nu_\mu$	$\nu_\tau$
Main production source	kaon decay	pion decay	charm decay
# traversing FASERnu 25cm x 25cm	$O(10^{11})$	$O(10^{12})$	$O(10^9)$
# interacting in FASERnu (1 tn Tungsten)	$\sim 1000$	$\sim 20000$	$\sim 10$



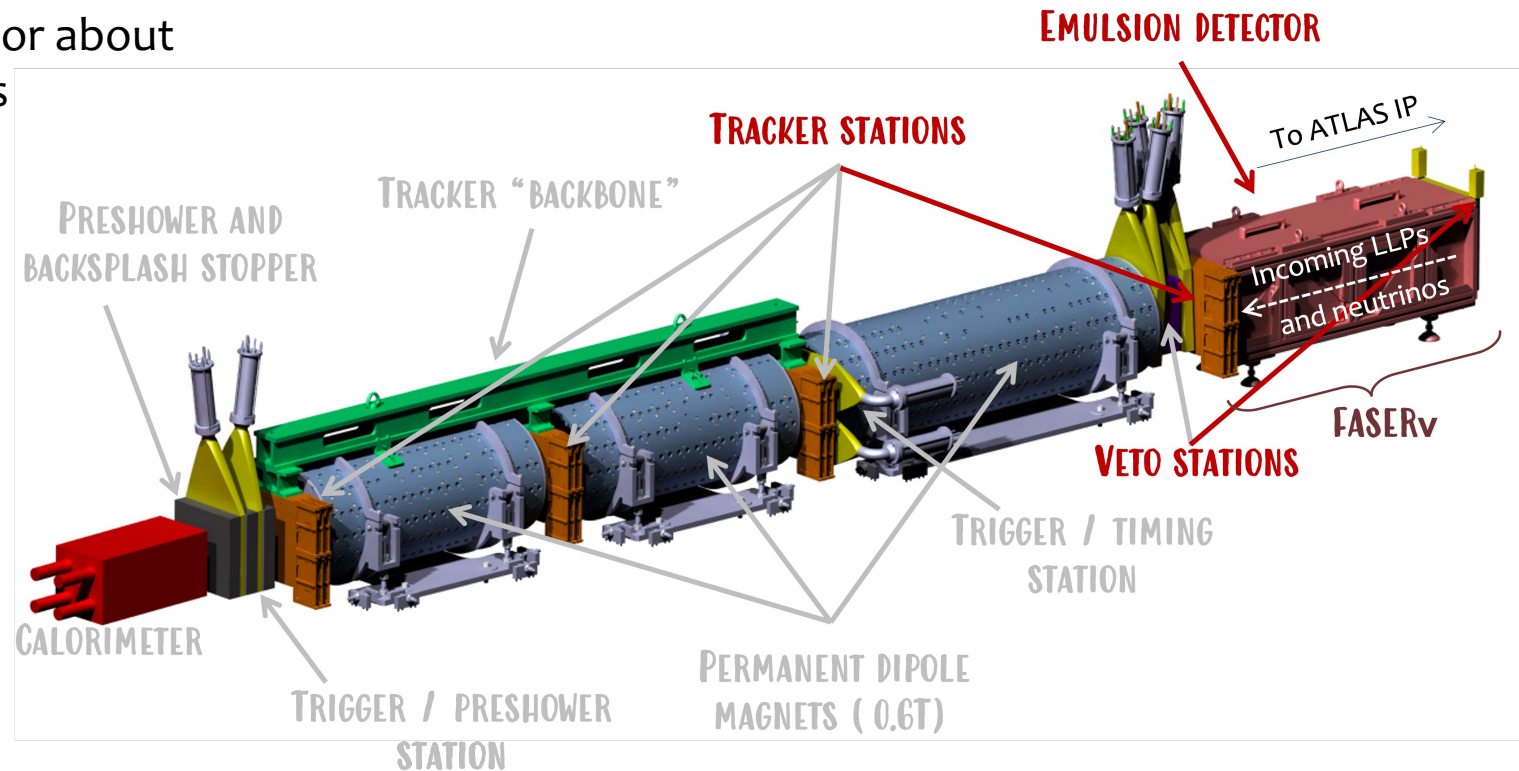
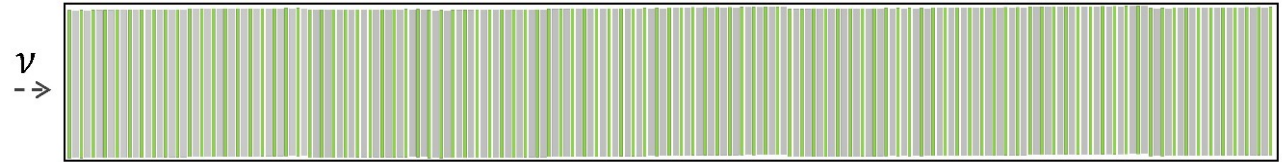
# FASER DETECTOR

Emulsion detector:

- Excellent track resolution
- No timing resolution
- **Challenges:** replace the 1-ton-scale detector about 3 times/year & **develop** the emulsion films



Total 730 emulsion films interleaved with 1.1-mm-thick tungsten plates

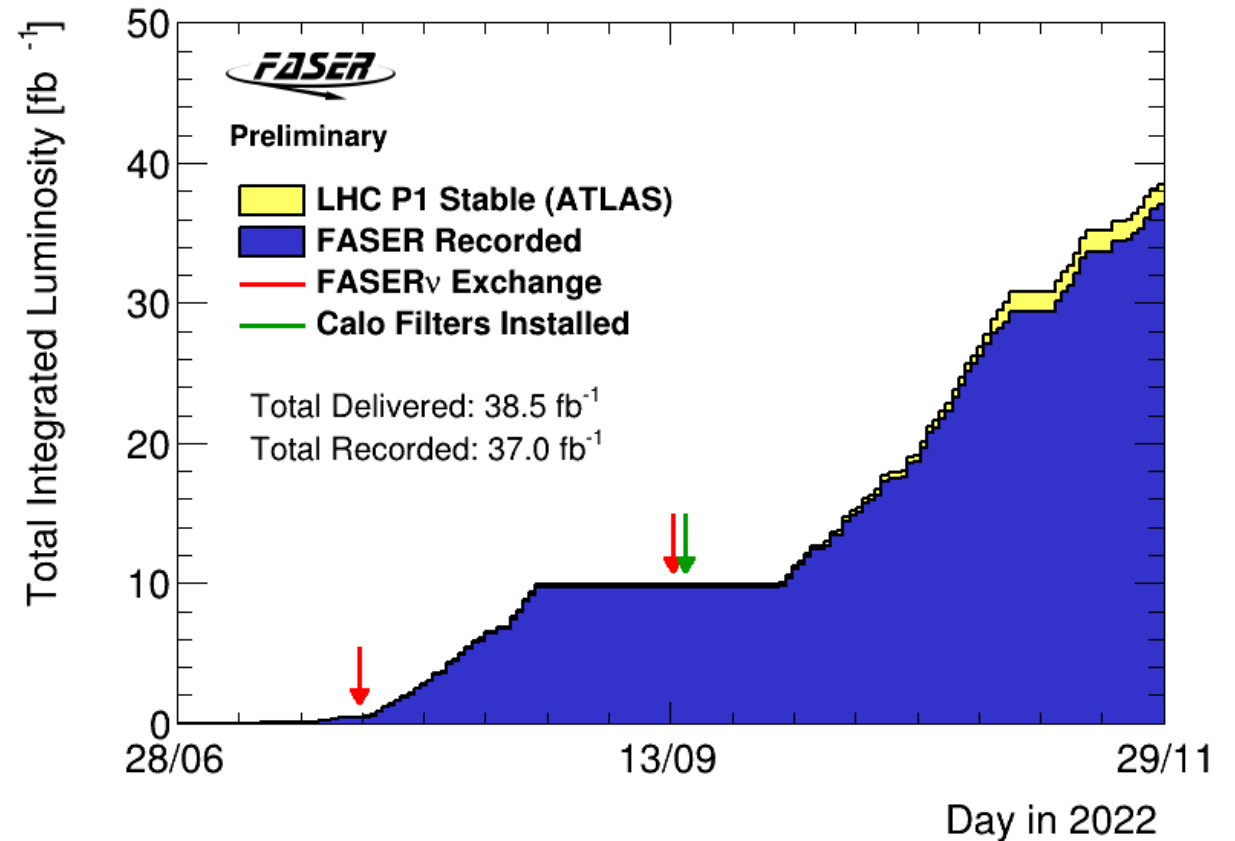


# FASER DETECTOR

Emulsion detector:

- Excellent track resolution
- No timing resolution
- **Challenges:** replace the 1-ton-scale detector about 3 times/year & **develop** the emulsion films

→ The first FASER neutrino detection is based the “electronic” part of the detector.



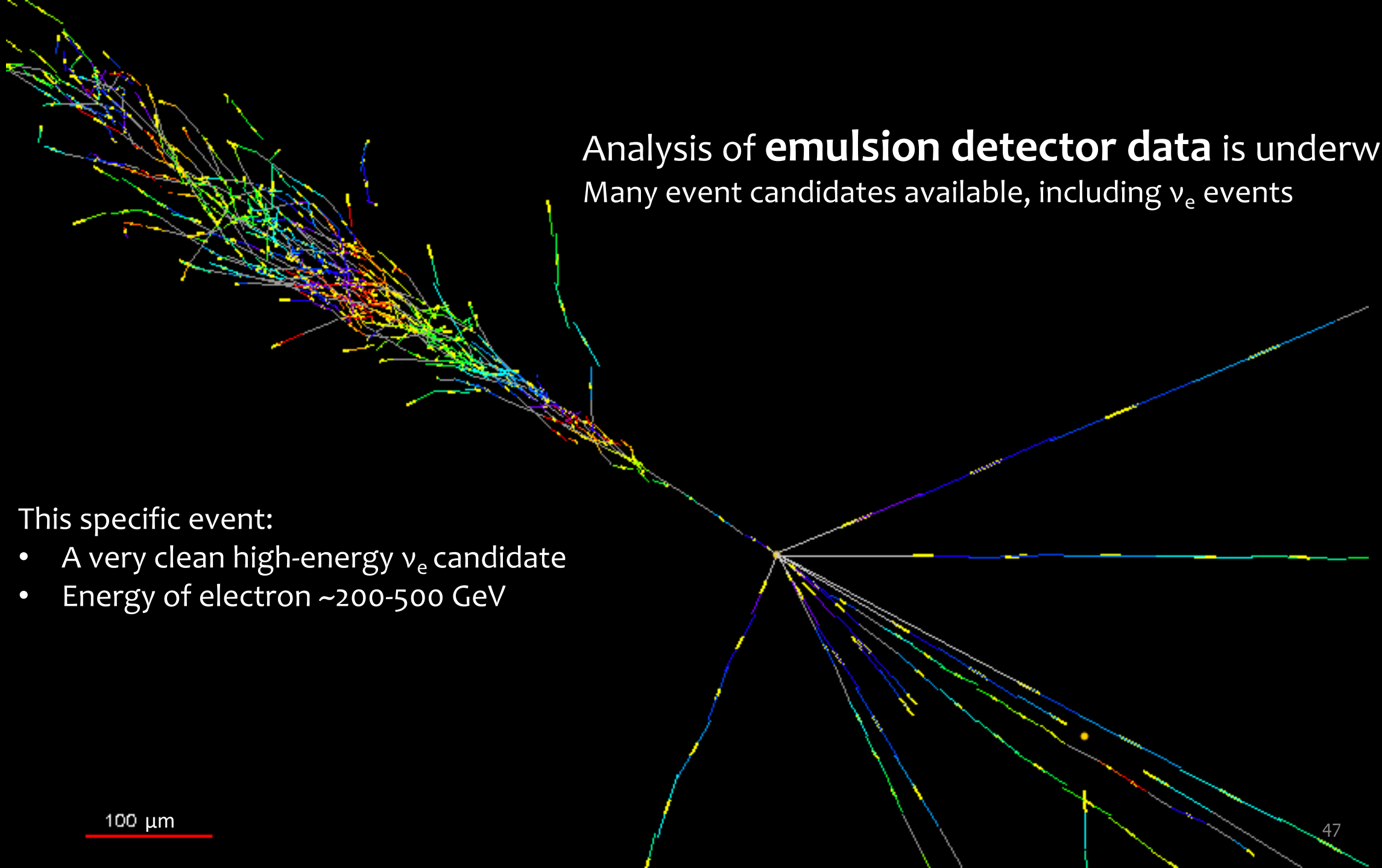


Analysis of **emulsion detector data** is underway  
Many event candidates available, including  $\nu_e$  events

This specific event:

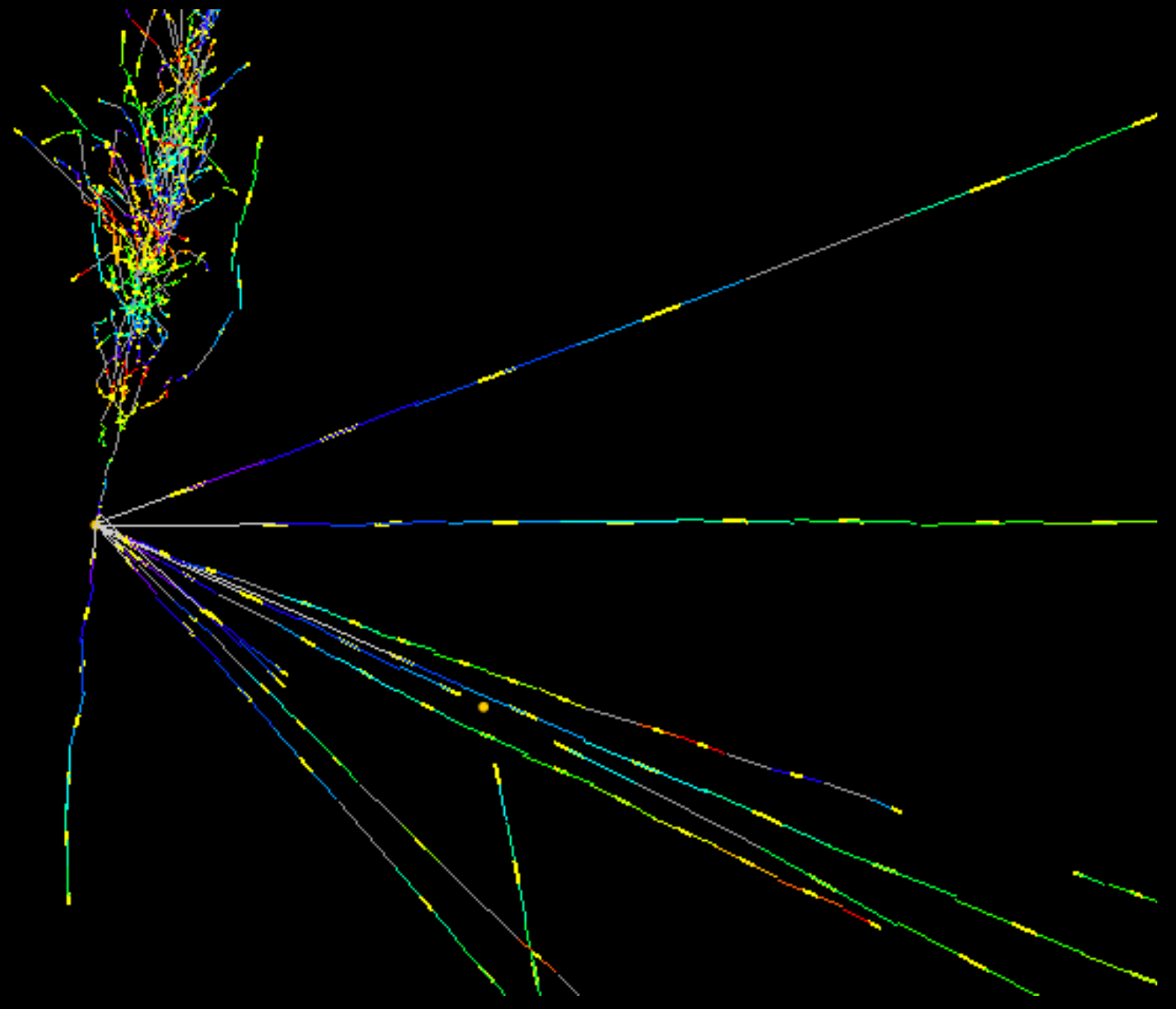
- A very clean high-energy  $\nu_e$  candidate
- Energy of electron  $\sim 200\text{-}500$  GeV

100  $\mu\text{m}$



100

---



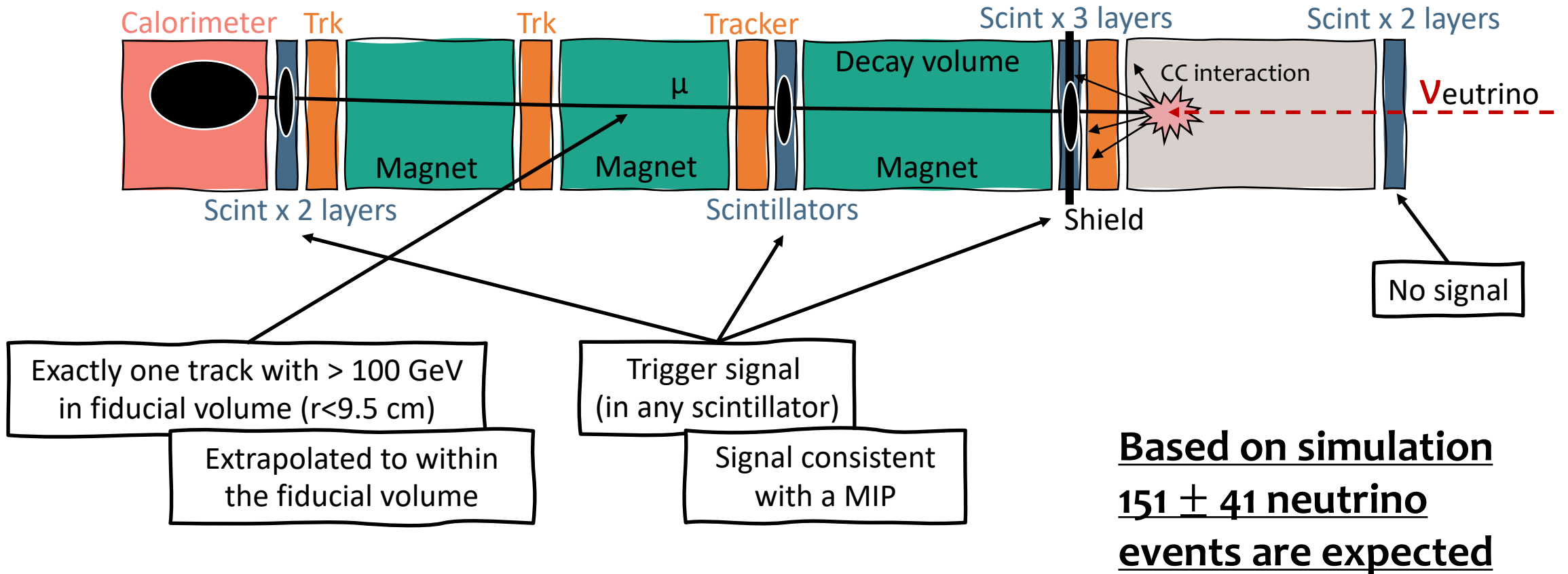
FIRST DIRECT NEUTRINO DETECTION IN A HADRON COLLIDER

WITH



# SIGNAL SELECTION

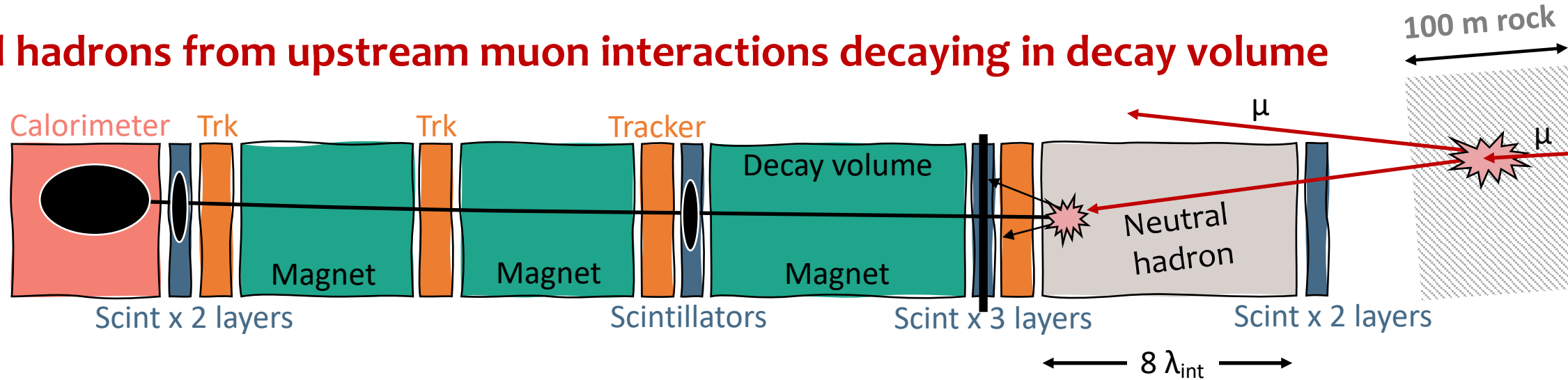
## Neutrinos in FASER



# BACKGROUNDS

## Neutrinos in FASER

### 1. Neutral hadrons from upstream muon interactions decaying in decay volume



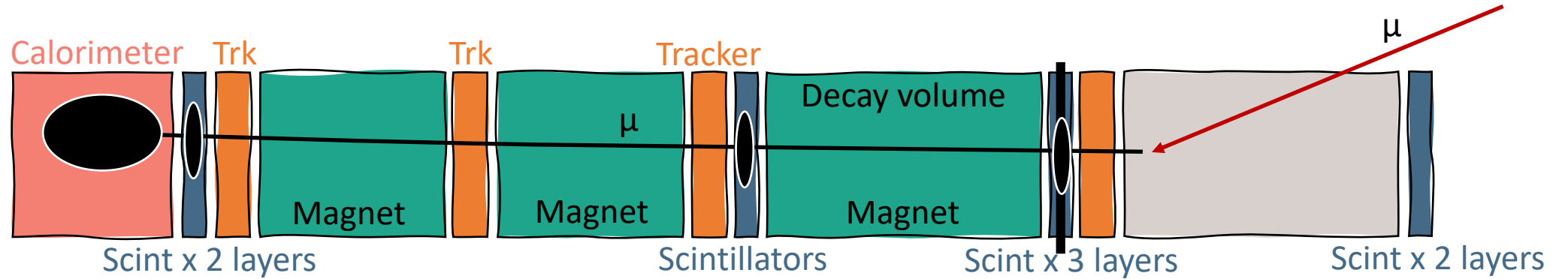
- Evaluated using two-step simulation
  - $O(300)$  neutral hadrons with  $E > 100$  GeV would reach FASER
  - Most of them absorbed in tungsten
- Estimated background:  **$(0.11 \pm 0.06)$  events**



# BACKGROUNDS

## Neutrinos in FASER

### 2. Scattered muons

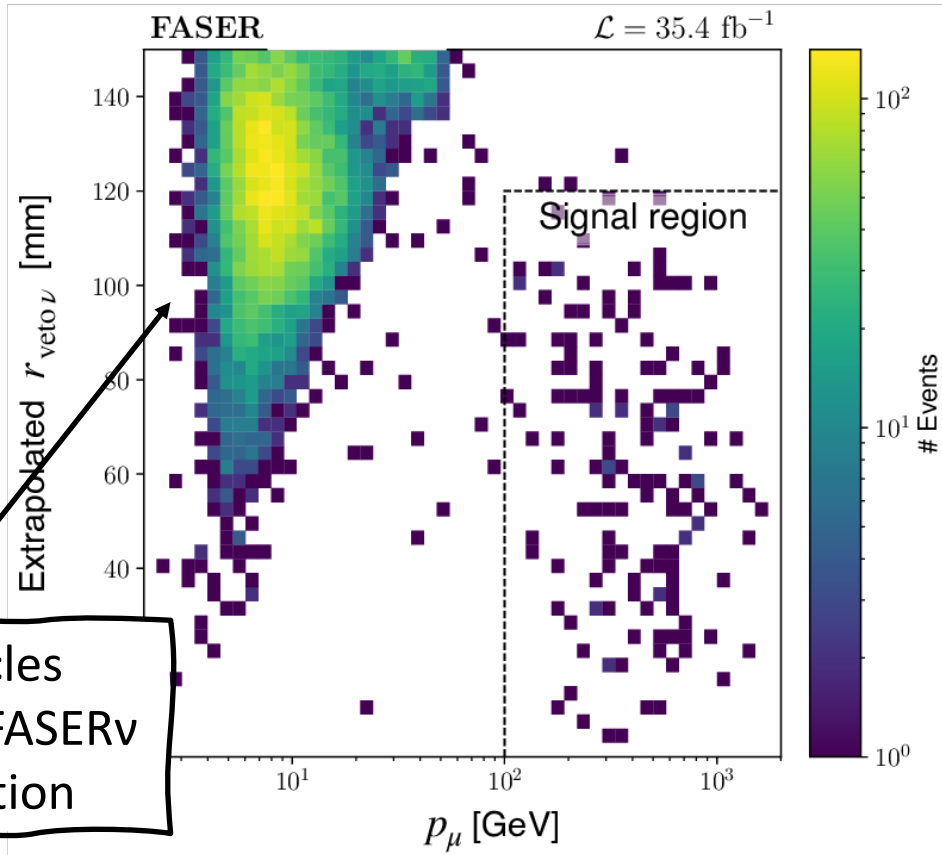


- Estimated from events with single track just outside the fiducial region
- Estimated background:  **$(0.08 \pm 1.83)$  events**

# LOOKING IN DATA

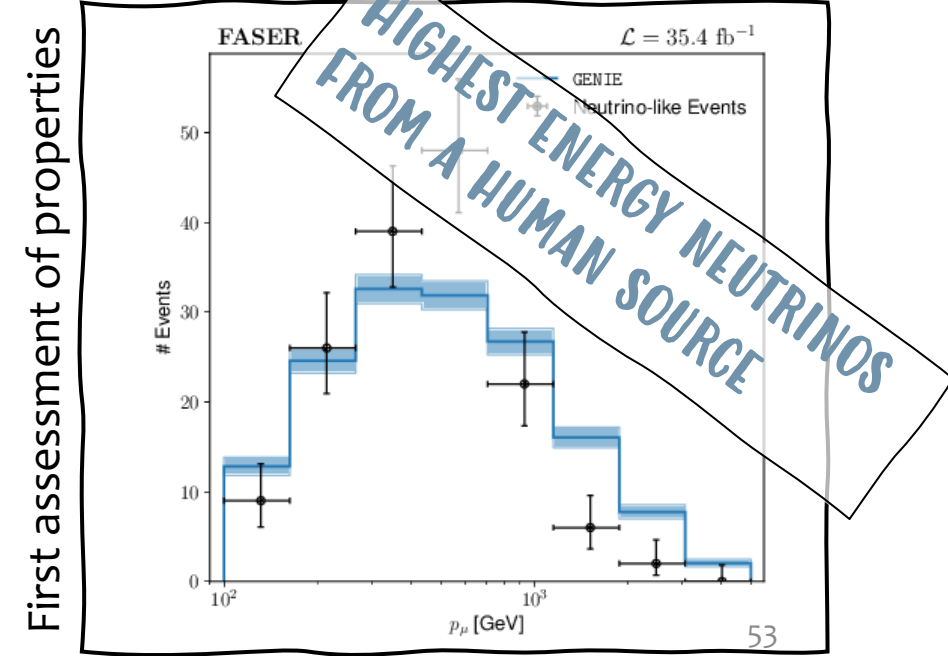
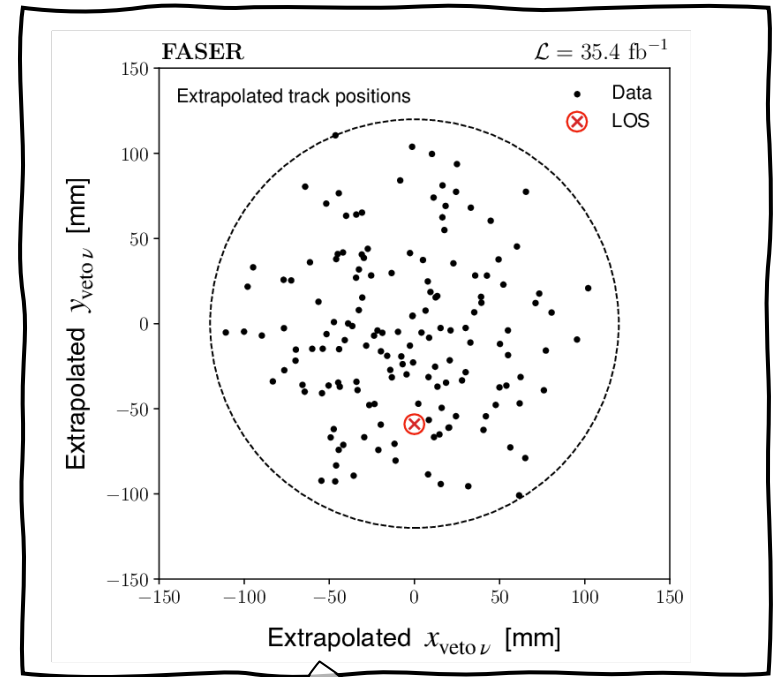
$153_{-13}^{+12}$  events  $\Rightarrow \gg 5\sigma$  significance

FIRST DIRECT OBSERVATION OF COLLIDER NEUTRINOS!



Charged particles that miss the FASERv scintillator station

N.B.: SND@LHC also had a collider neutrino observation with 8 signal events and  $\sim 0.1$  background events (arXiv:2305.09383)



**FASER**

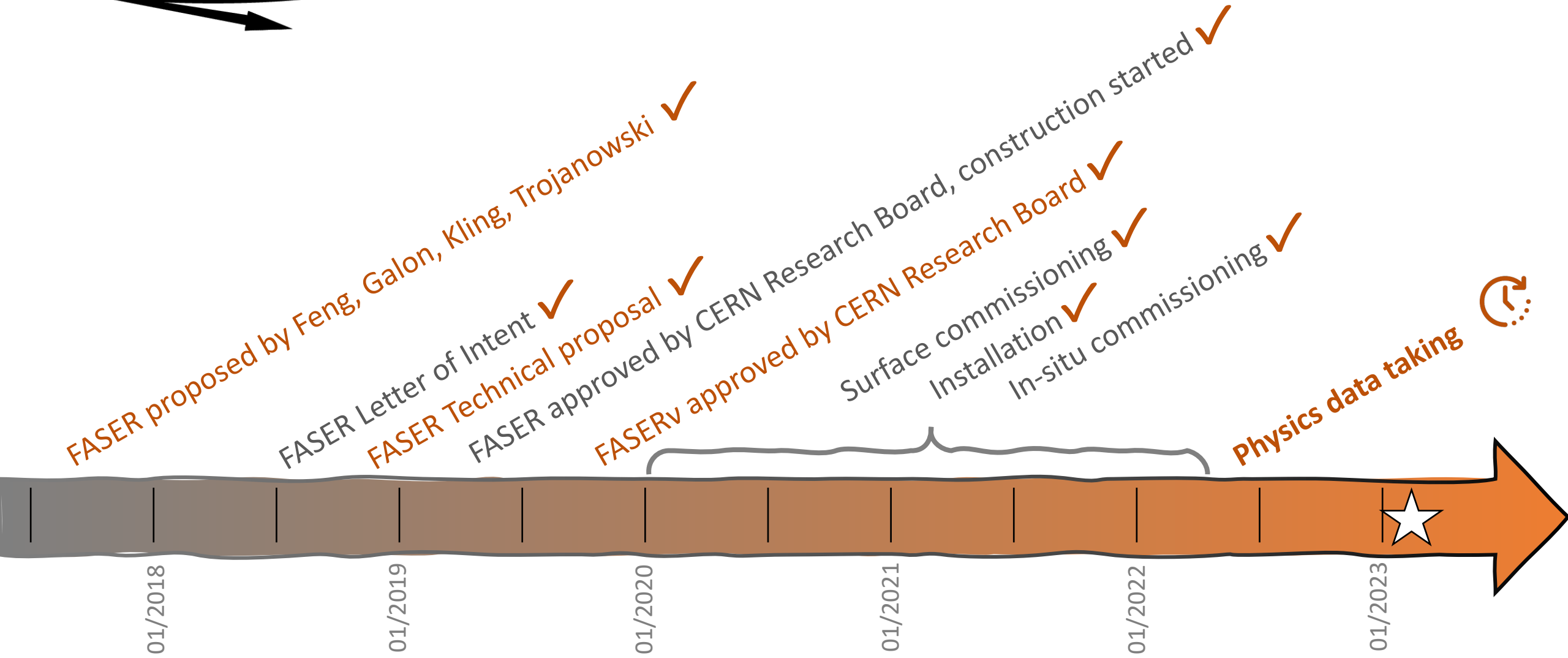
LHC Fill 8315, Event 47032829,  
2022-10-27 08:52:45

↑↑  
57 clusters

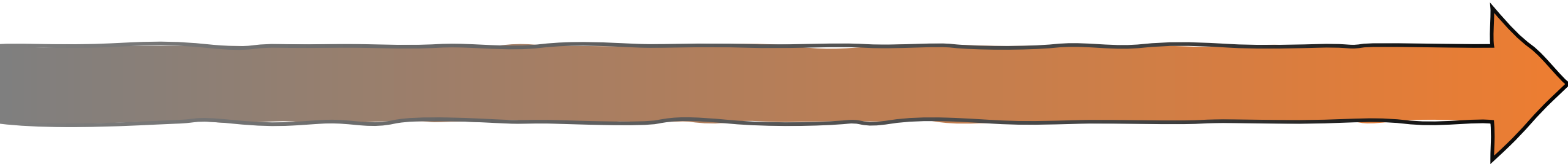
$p_{\mu} \approx 850$  GeV, negative charge



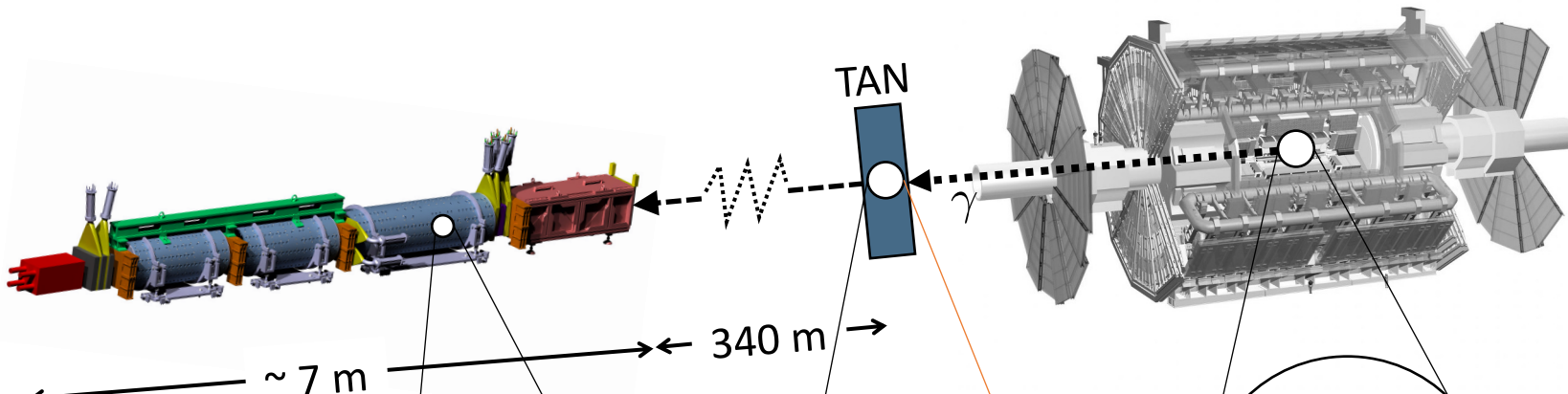
# GLOBAL TIMELINE



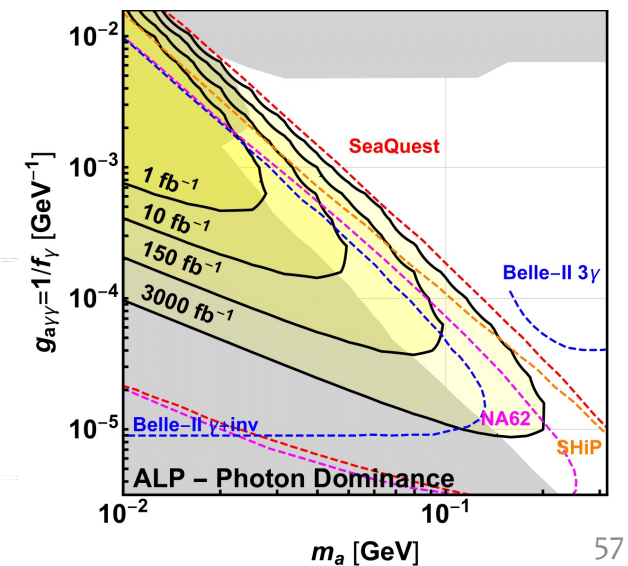
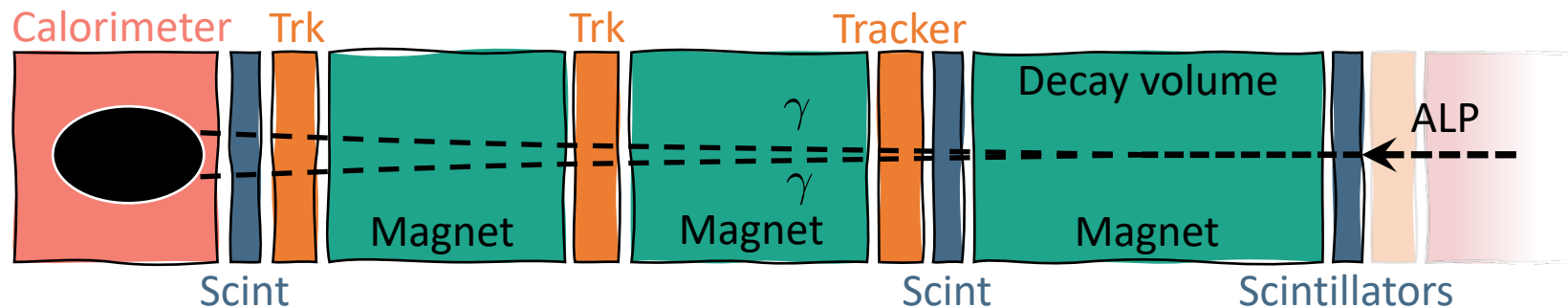
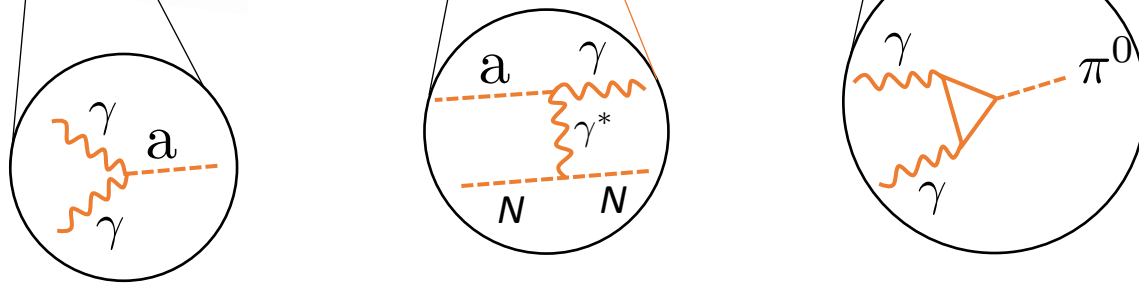
**WHAT'S BEYOND 2023 ?**



# ANOTHER PHYSICS EXAMPLE Axion-like particle (ALP, $a$ )



Decay products collimated  
 2- $\gamma$  signature can't be resolved  
 with present detector: **upgrade**



Assuming 3 signal events  
 and no backgrounds

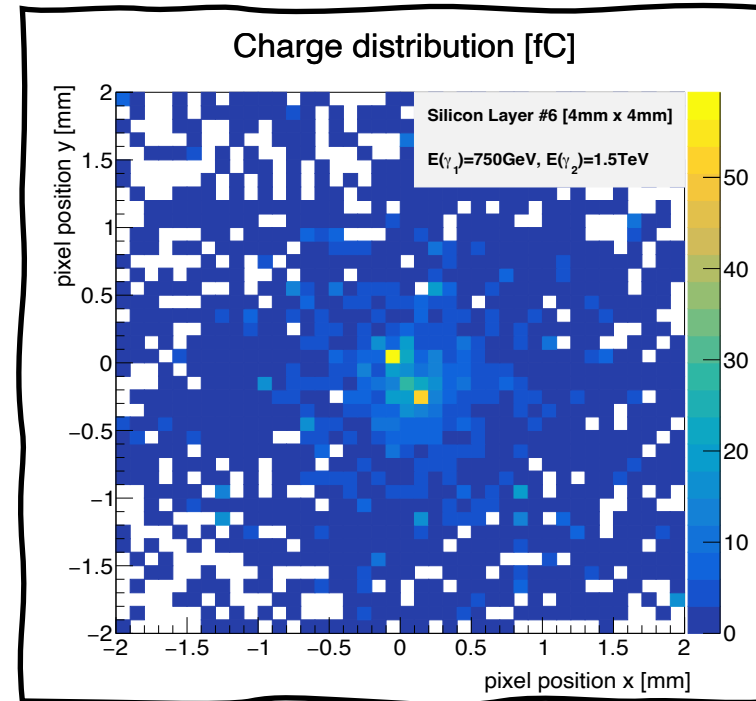
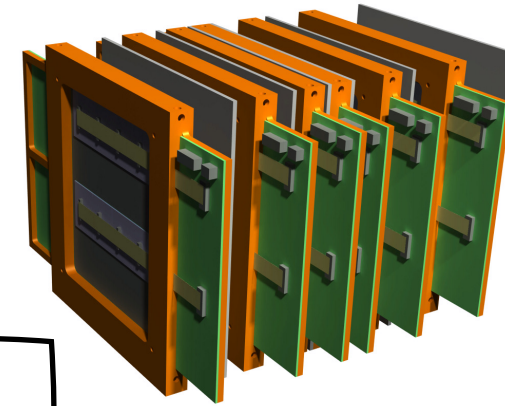


# PRESHOWER UPGRADE

## TO ENABLE $2-\gamma$ PHYSICS

- Existing pre-shower to be replaced with a high-resolution silicon pre-shower detector using monolithic pixel ASICs

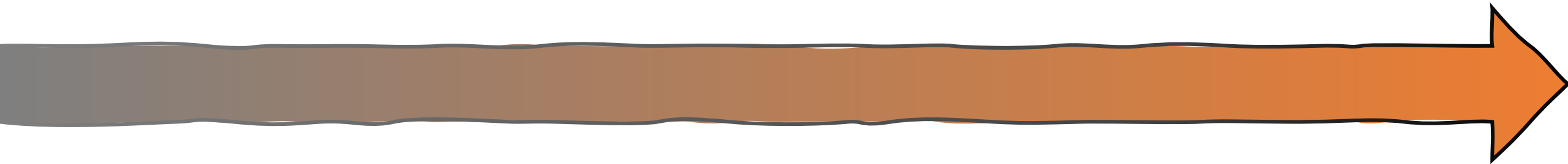
Preproduction ASICs in testbeam, Sept 2022



- Distance between two photons:  $200\ \mu\text{m}$
- Distinguishable!

Detector to be used for 2025 data taking

**WHAT'S BEYOND RUN 3 ?**



# BEYOND FASER?

A TEASER FOR THE PROPOSED

# FORWARD PHYSICS FACILITY

The rich physics program in the far-forward region strongly motivates creating a dedicated Forward Physics Facility to house far-forward experiments for the HL-LHC era from 2028-2040s

Well aligned with the recommendations of recent community studies

*The full physics potential of the LHC and the HL-LHC [...] should be exploited.*

1st recommendation of the 2020 European Strategy Update

*Our highest immediate priority accelerator and project is the HL-LHC, [...] including the construction of auxiliary experiments that extend the reach of HL-LHC in kinematic regions uncovered by the detector upgrades.*

Snowmass 2021 Energy Frontier Report

More: [Submitted to P5](#) just in April 2023

[LoI for SNOWMASS-2021](#)

[arXiv:2203.05090](#)

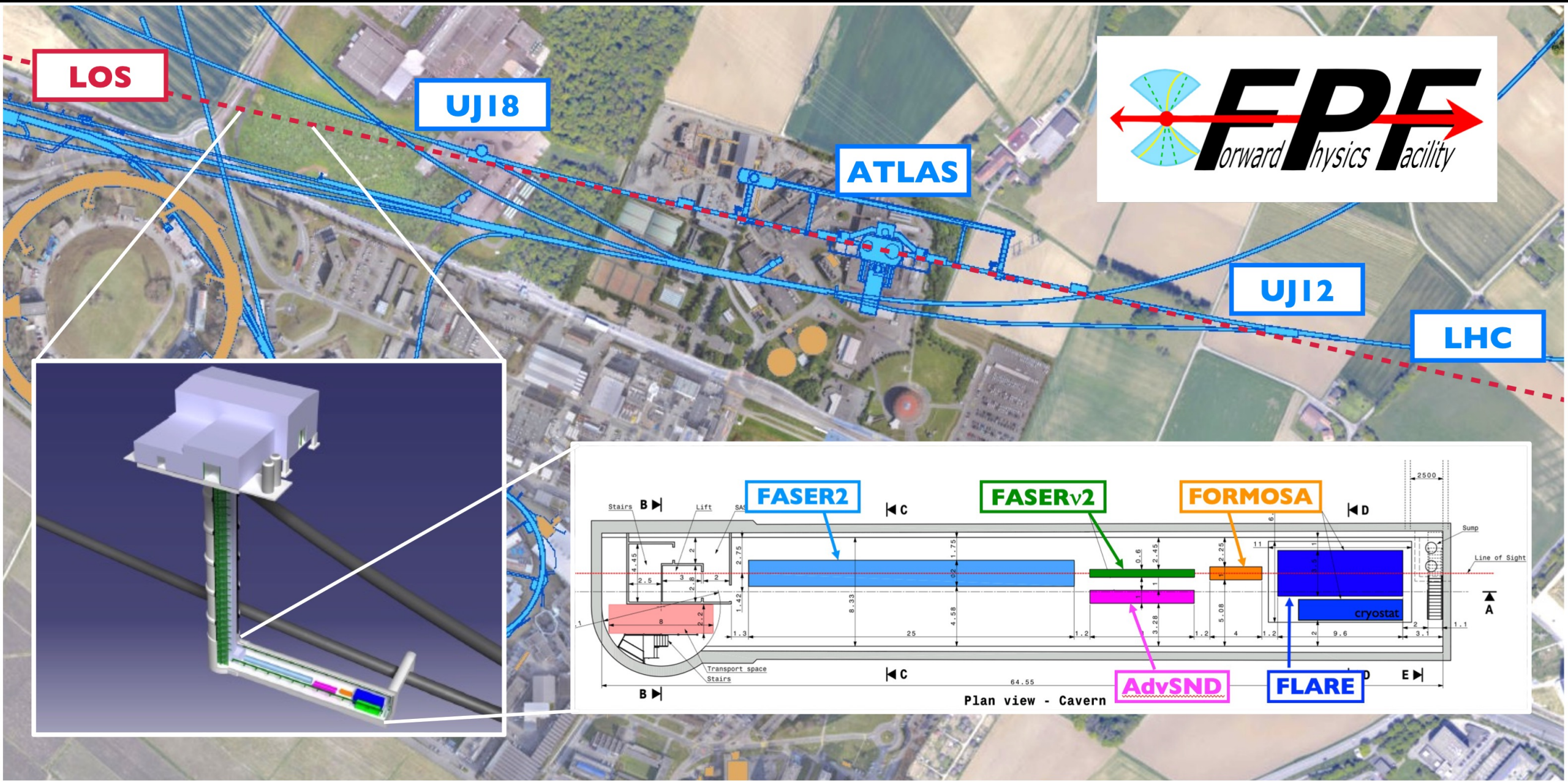
[FPF – Kickoff workshop](#)

[FPF – 5<sup>th</sup> \(latest\) workshop](#)

[FPF – 6<sup>th</sup> workshop coming up next week!](#)

Please join if interested!





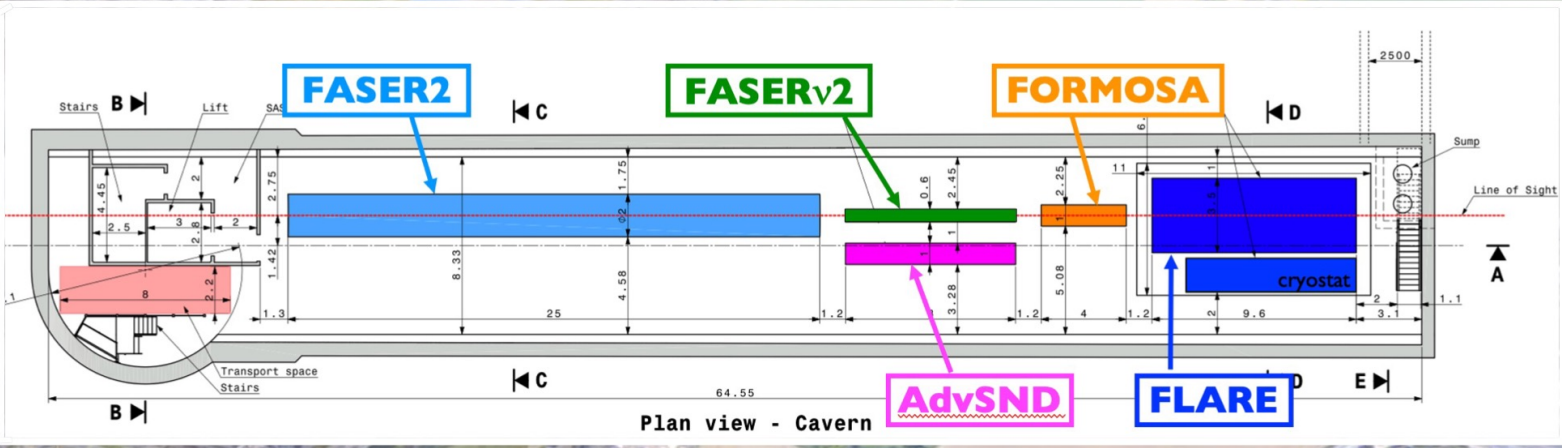
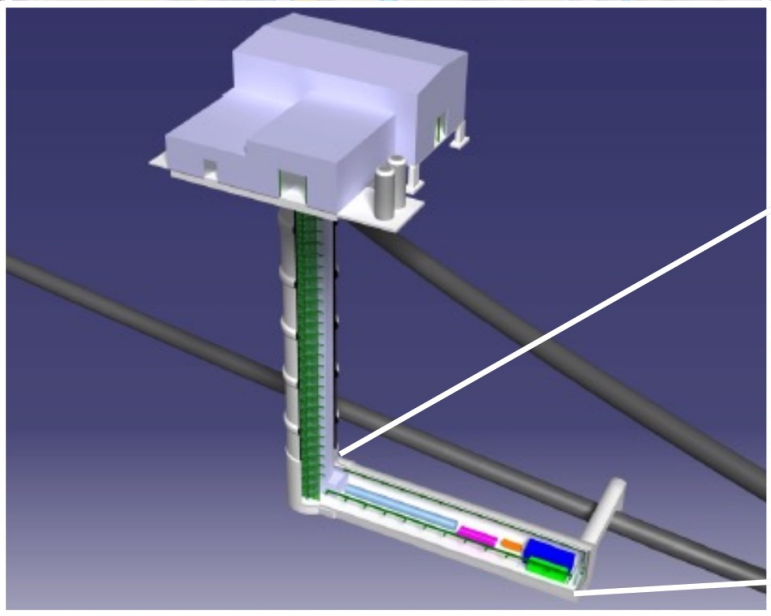
LOS

UJI18

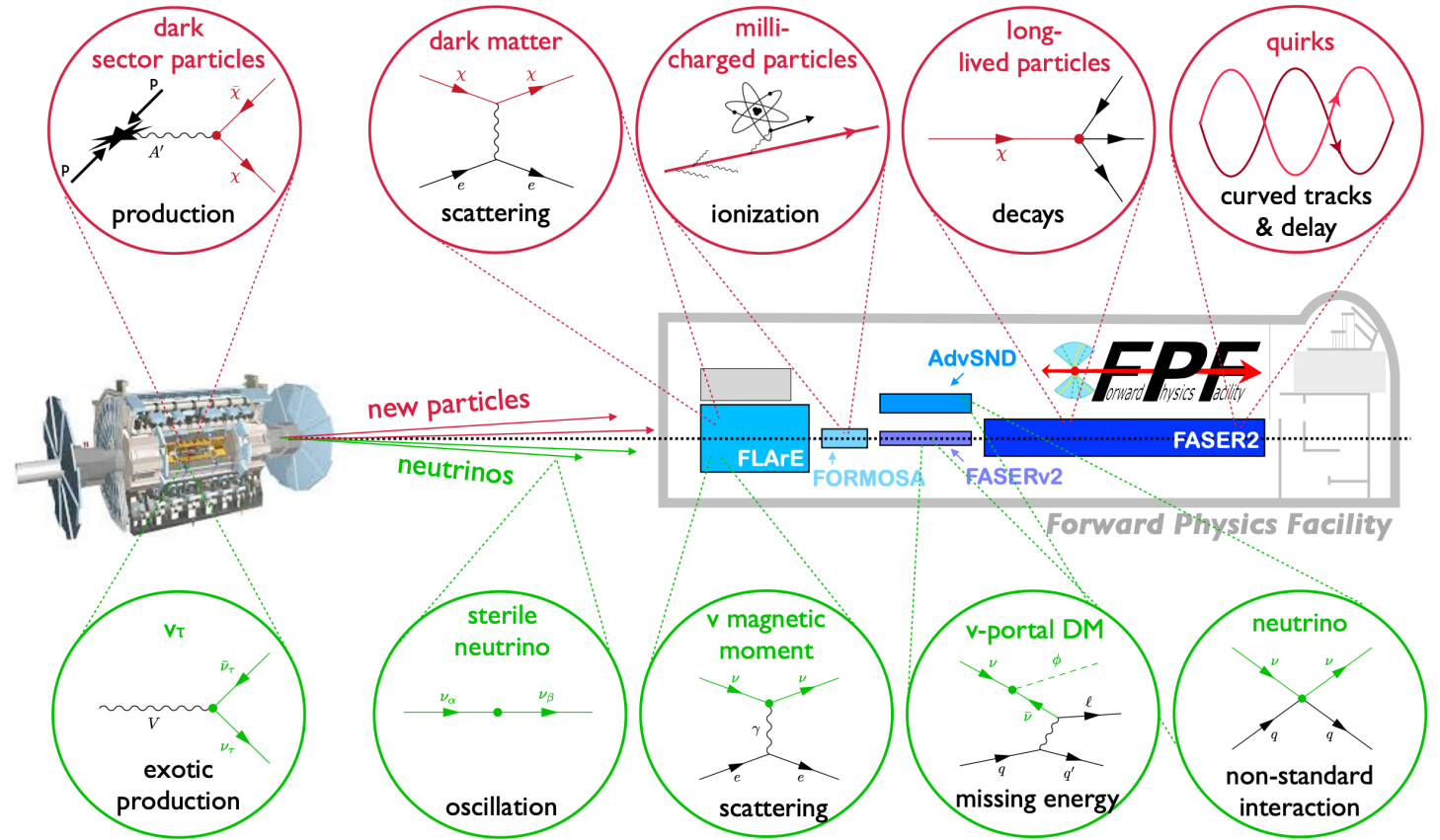
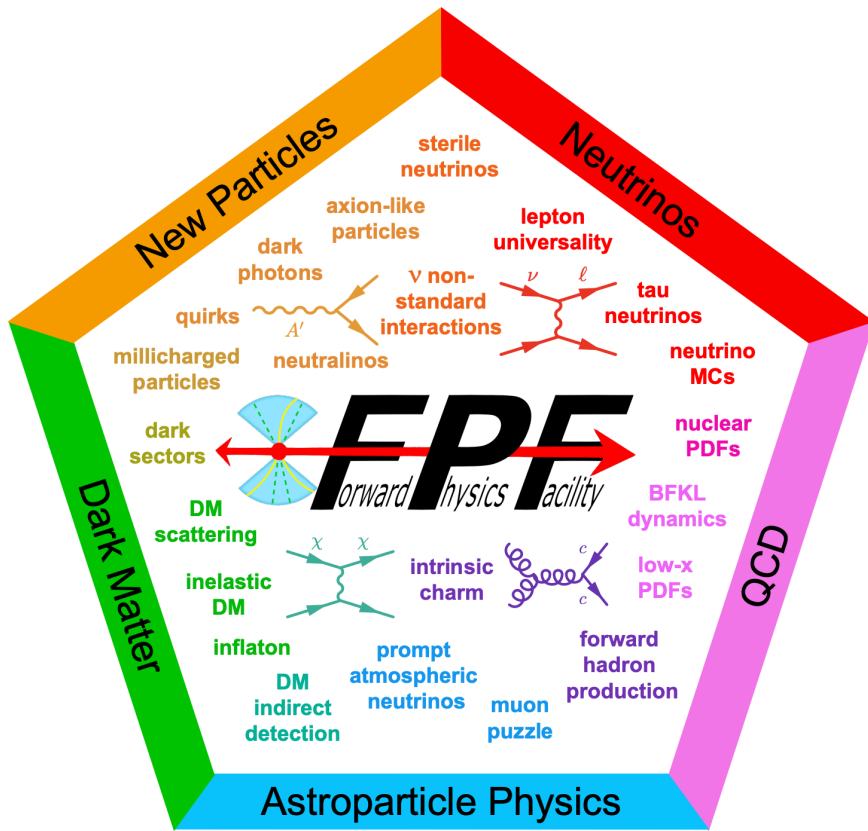
ATLAS

UJI2

LHC



# THE PHYSICS PROGRAMME OF FPF

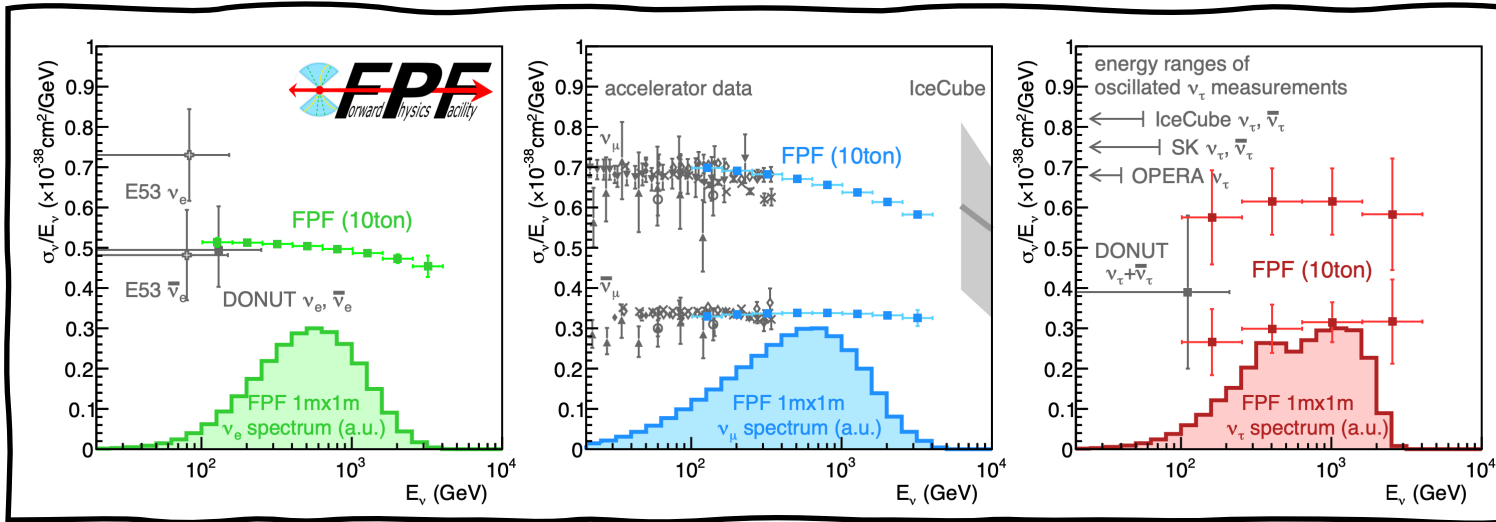




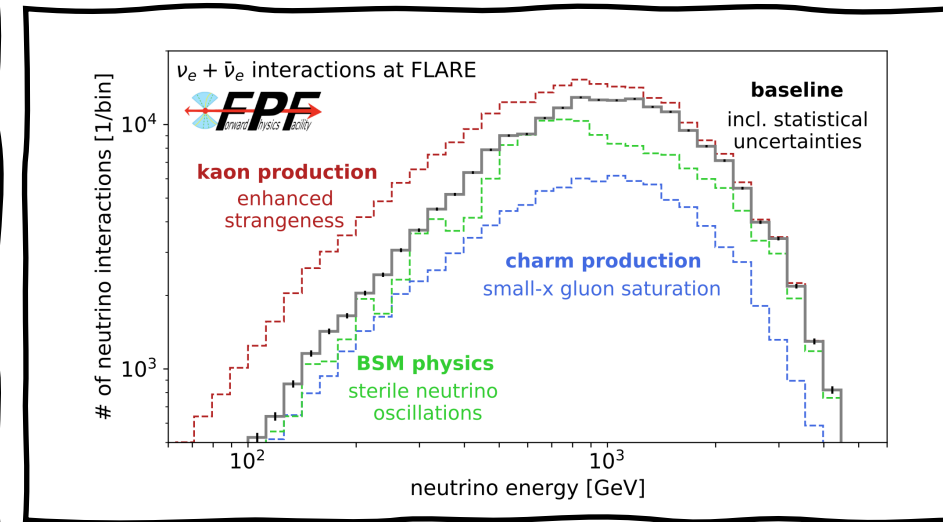
# A RICH NEUTRINO PROGRAMME

	Available lumi	Mass of $\nu$ detector	$\nu_e$	$\nu_\mu$	$\nu_\tau$
Main production source			kaon decay	pion decay	charm decay
# interacting in FASERv	150 / fb	1 tn Tungsten	$\sim 1000$	$\sim 20000$	$\sim 10$
# interacting in FASERv2	3000 / fb	10 tn Tungsten	$\sim 10^5$	$\sim 10^6$	$\sim 10^4$

Unprecedented numbers of detectable neutrinos, at energy ranges where there is currently no available data!



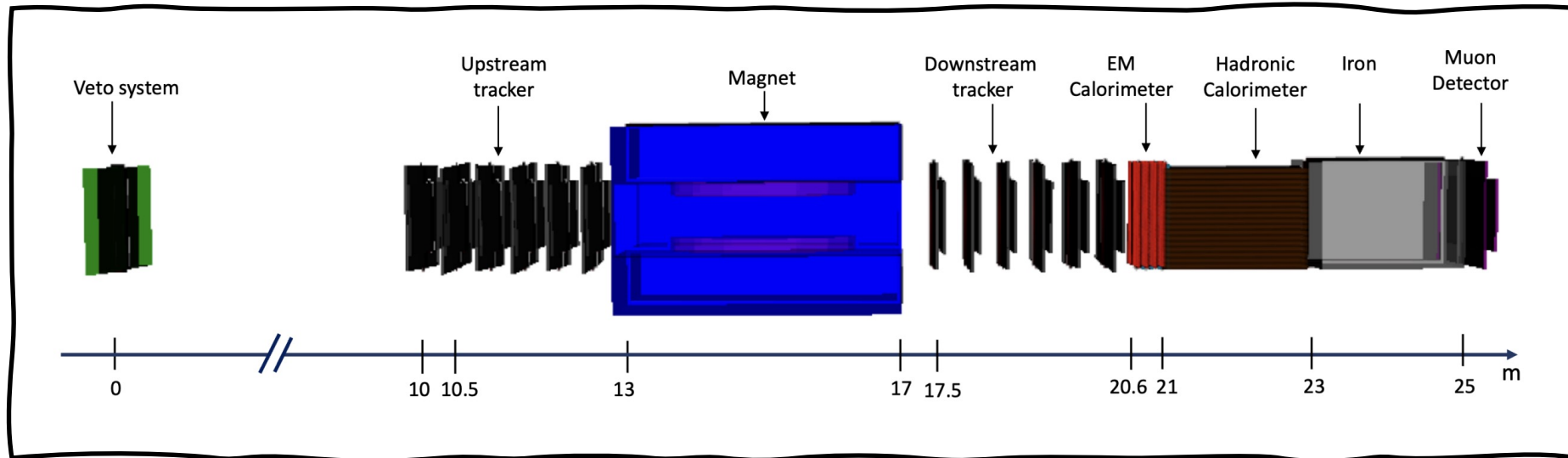
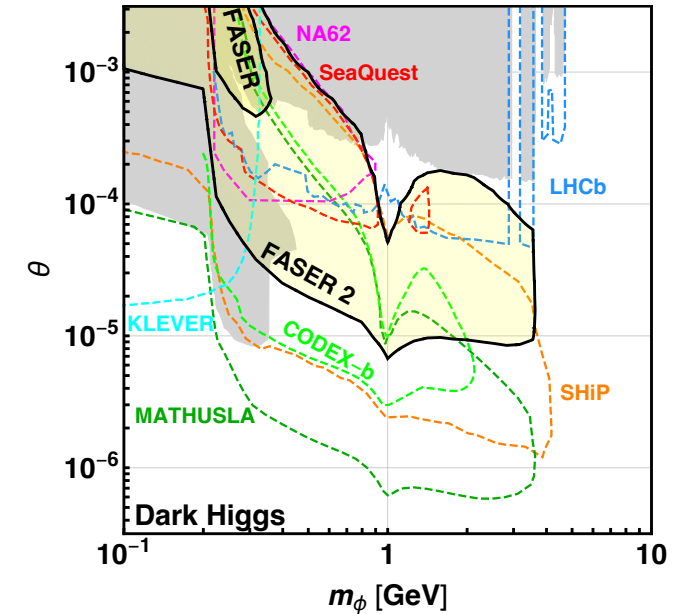
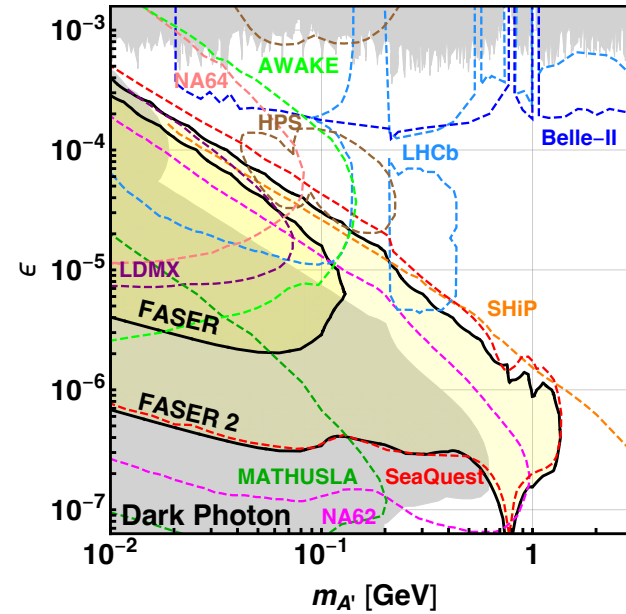
Expected precision of FPF measurements of the neutrino interaction cross section with nucleons



Coloured lines: three examples of physics that can change the expected flux, all probed at FPF

# BSM & FASER2

Increased detector radius to 1 m allows sensitivity to particles produced in heavy meson (B, D) decays increasing physics case **beyond** just increased luminosity



Possible FASER2 layout



# OUTLOOK

- The FASER experiment introduces a **novel approach** to exploit LHC collisions, to:
  - either **make a new discovery or constrain parts of phase-space which no current experiment has access to**; and
  - make the first **collider-originated neutrino measurements**
- Collaboration (& CERN technical teams) worked feverishly to construct, install & commission the detector over the Long Shutdown, & successfully collect Run 3 data
- **First FASER results available !**
  - **Dark photon exclusion in interesting thermal relic region**
  - **First collider neutrino direct detection**
- Have started upgrades and thinking about FASER2 & a future facility to further exploit forward production in LHC collisions!
- **LOTS OF EXCITING PHYSICS AHEAD !**

Stay in touch:



<https://faser.web.cern.ch/>



@FASERexperiment

# FASER THANKS!

For financial support:



**HEISING-SIMONS  
FOUNDATION**

**SIMONS  
FOUNDATION**



**Swiss National  
Science Foundation**



**科研費  
KAKENHI**

FASER Collaboration: 24 institutes, about 87 members

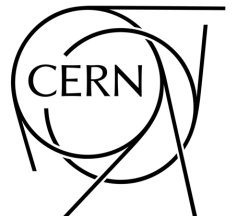


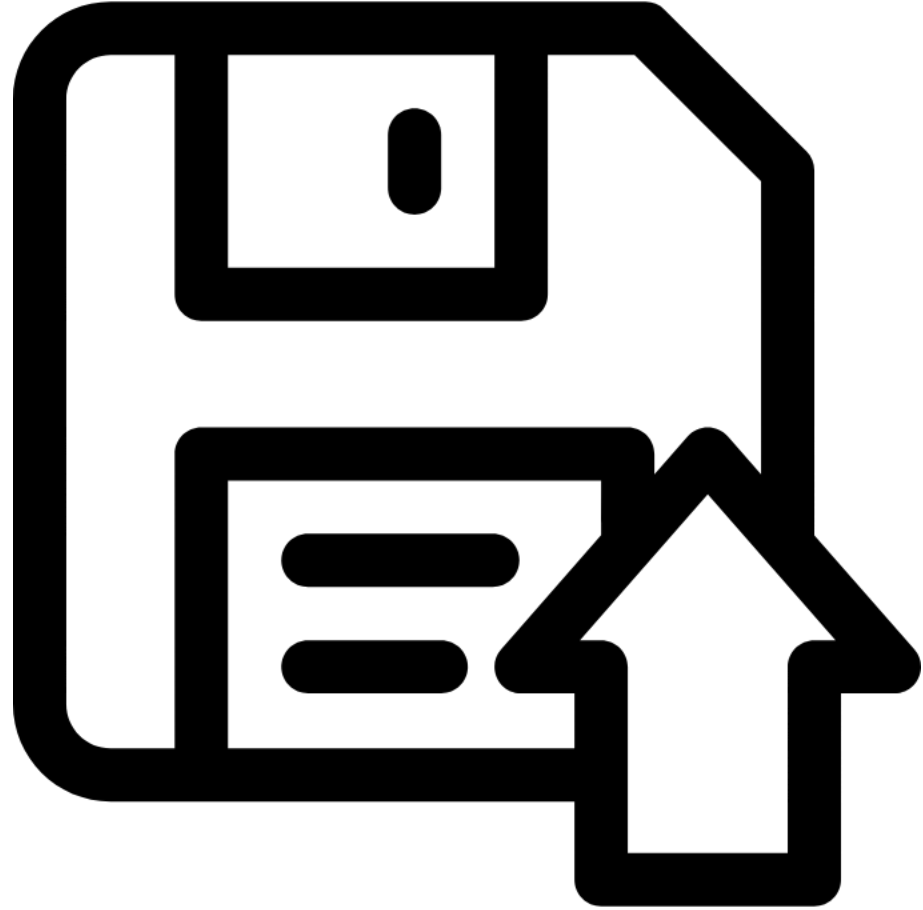
International laboratory  
covered by a cooperation  
agreement with CERN

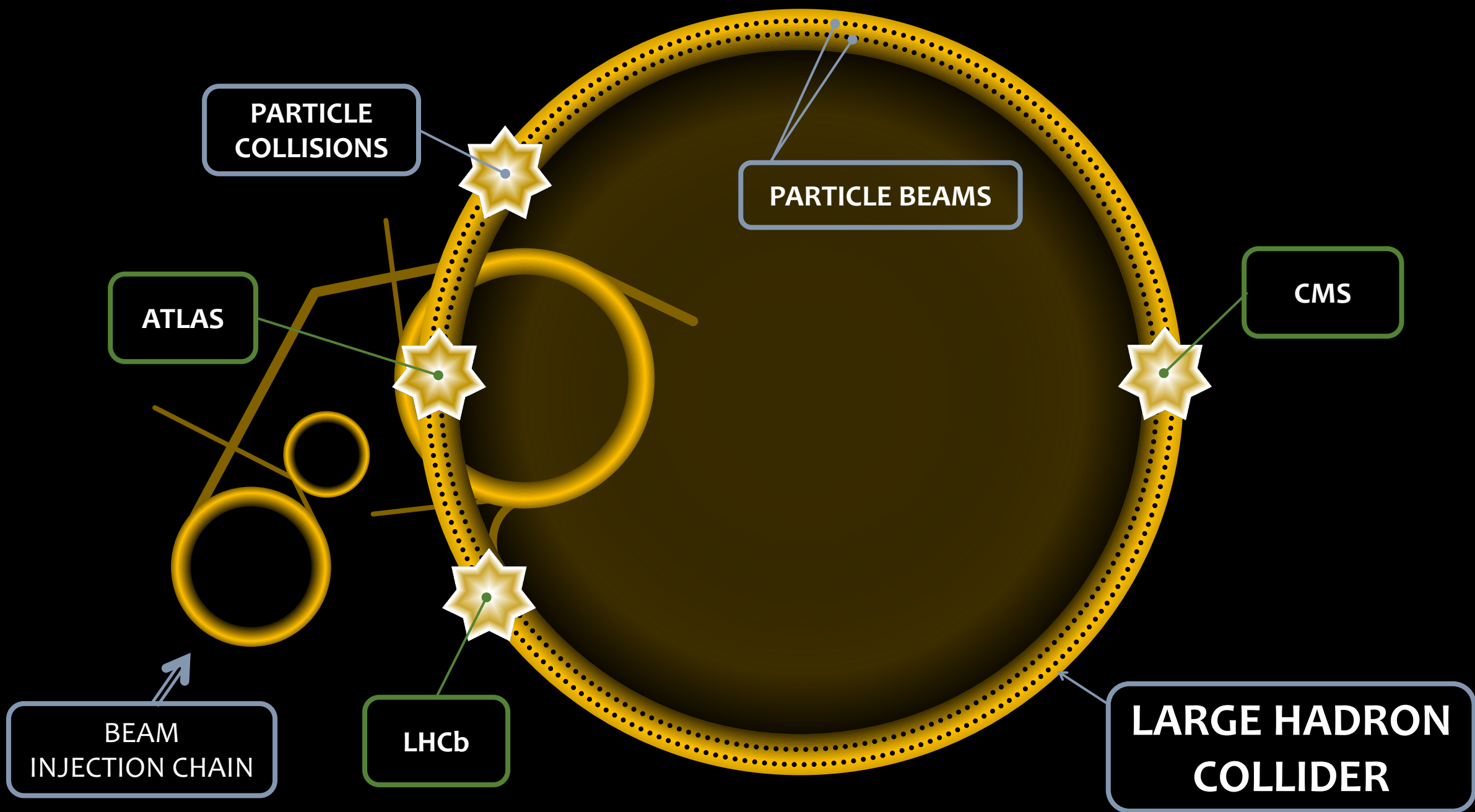


**UNIVERSITÉ  
DE GENÈVE**

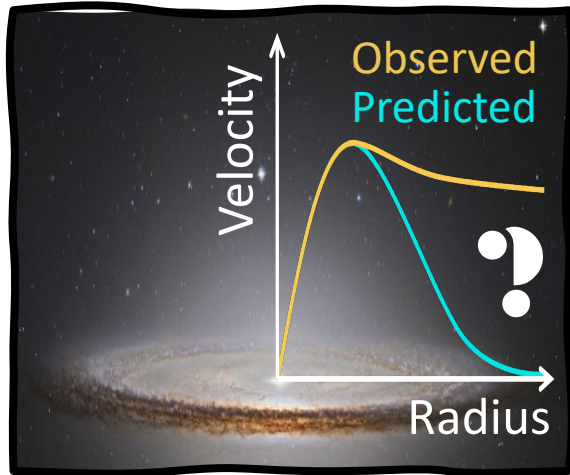
FACULTY OF SCIENCE



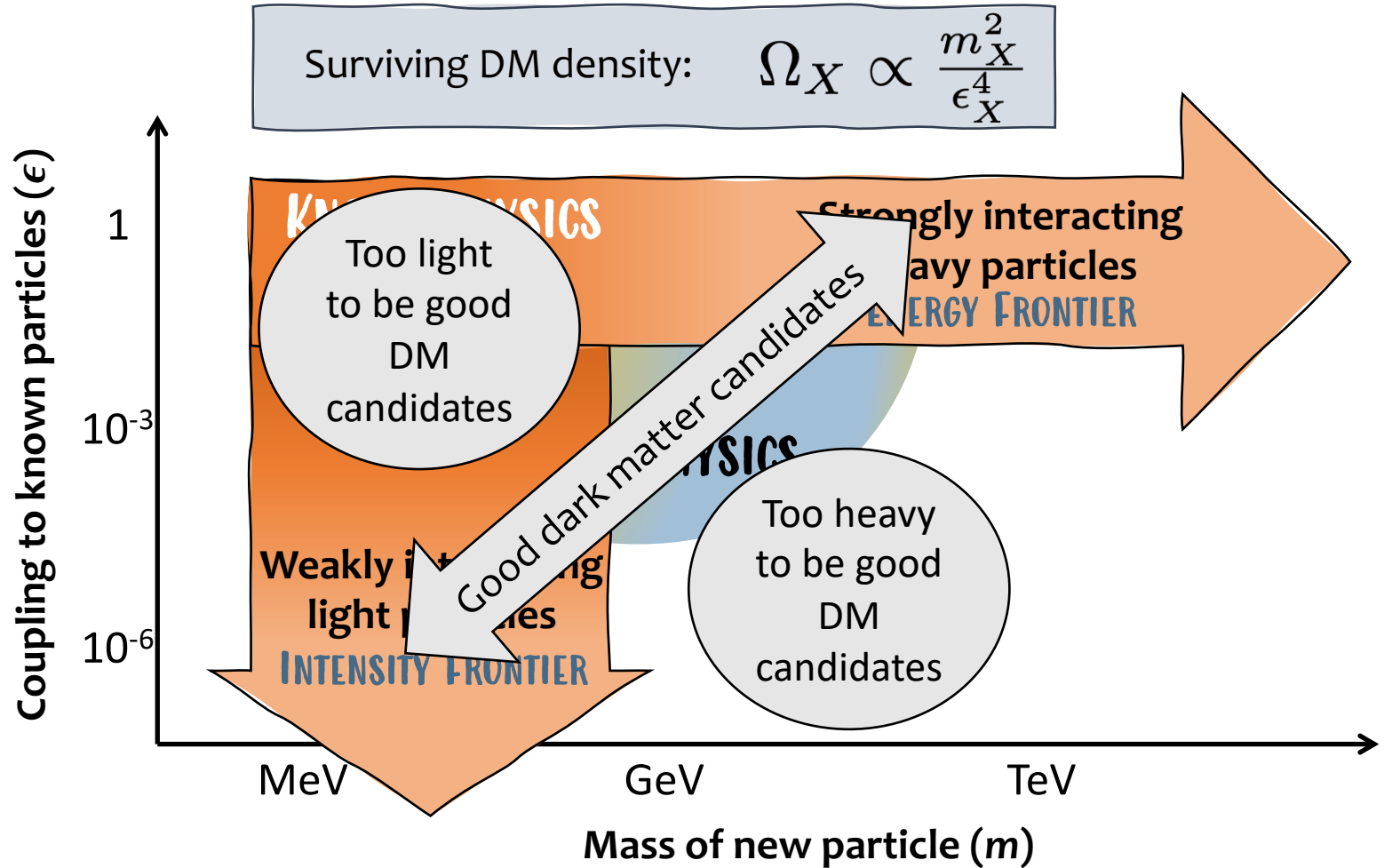
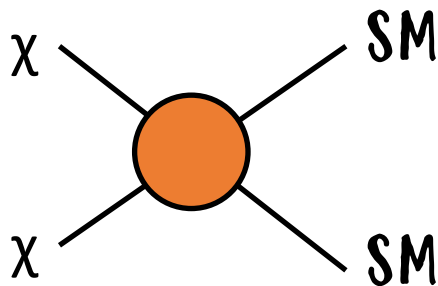




# THE LANDSCAPE OF NEW PARTICLES @ COLLIDERS



- Simple mechanism for DM evolution: “freeze out”







# FEEBLY INTERACTING PARTICLES (FIPs)

- Due to interacting feebly, they are linked to a “hidden sector”
- Couplings between SM and hidden sector result from “portal” operators
- Large number of specific models; can be simplified to the following:

SM  
Higgs  $h$

$h \text{ ----- } (\mu S + \lambda S^2) H^\dagger H \text{ ----- } S$

Dark  
Higgs  $S$

New scalar: **Dark Higgs**; couplings to SM  $\mu, \lambda$

SM  
EM  $A$

$A \rightsquigarrow -\frac{\epsilon}{2 \cos \theta_W} F'_{\mu\nu} F_Y^{\mu\nu} \rightsquigarrow A_D$

Dark  
EM  $A_D$

New vector: **Dark photon**; coupling to SM  $\propto \epsilon Q$

SM  
 $2\gamma$  or  $2f$

$2\gamma \text{ --- } \frac{\alpha}{f_\alpha} F_{\mu\nu} \tilde{F}^{\mu\nu}$   
 $2f \text{ ---- } \frac{\partial_\mu \alpha}{f_\alpha} \bar{\psi} \gamma^\mu \gamma^5 \psi$

ALP  
 $\alpha$

New pseudo-scalar: **ALP**; coupling to SM suppressed  
(Axion Like Particle)

SM  
LH  $\nu$

$\nu \text{ --- } y_N h L \psi_D \text{ --- } N$

HNL  
 $N$

New fermion: **HNL**; coupling to LH SM and  $h \propto y_N$   
(Heavy Neutral Lepton)

- The masses of the new particles can span several orders of magnitude

# DARK PHOTON PHYSICS

Dark photon's properties defined through the Lagrangian terms :

$$\mathcal{L} \supset \frac{1}{2} m_{A'}^2 A'^2 - \epsilon e \sum_f q_f A'_\mu \bar{f} \gamma^\mu f ,$$

For dark photon masses  $2m_e < m_{A'} < 2m_\mu \simeq 211 \text{ MeV}$ , dark photons decay to electrons with  $B(A' \rightarrow e^+e^-) \approx 100\%$  :

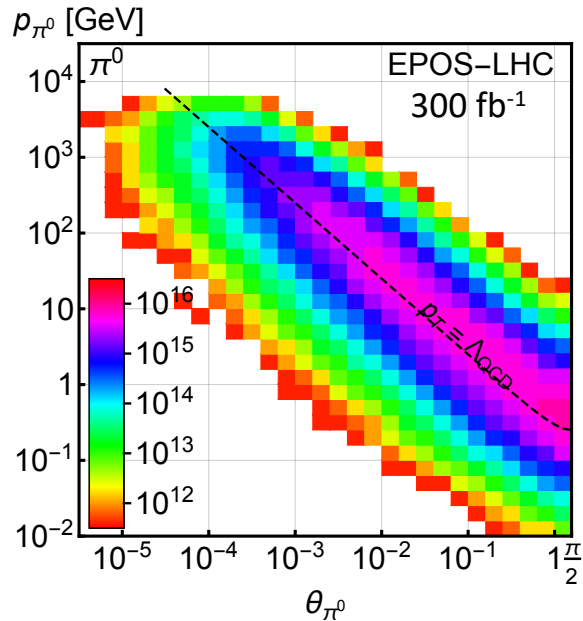
$$\Gamma_e \equiv \Gamma(A' \rightarrow e^+e^-) = \frac{\epsilon^2 e^2 m_{A'}}{12\pi} \left[ 1 - \left( \frac{2m_e}{m_{A'}} \right)^2 \right]^{1/2} \left[ 1 + \frac{2m_e^2}{m_{A'}^2} \right] .$$

For  $E_{A'} \gg m_{A'} \gg m_e$ , the dark photon decay length :

$$L = c\beta\tau\gamma \approx (80 \text{ m}) \left[ \frac{10^{-5}}{\epsilon} \right]^2 \left[ \frac{E_{A'}}{\text{TeV}} \right] \left[ \frac{100 \text{ MeV}}{m_{A'}} \right]^2$$

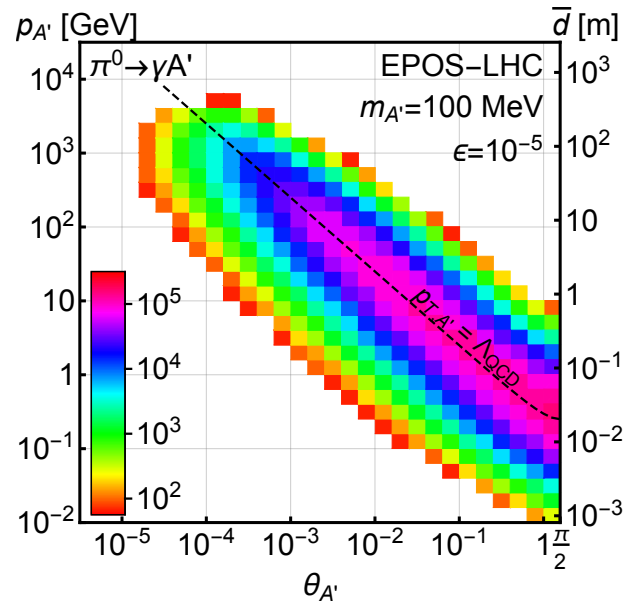
# SIGNALS: DARK PHOTONS

Pions at the IP



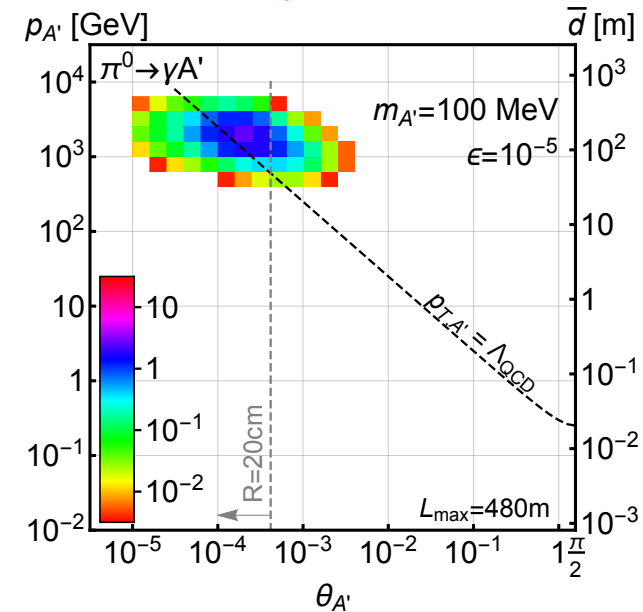
- Enormous event rates:  $N_{\pi} \sim 10^{15}$  per bin
- Production is peaked at low transverse momentum  $\sim 250$  MeV

A's at the IP



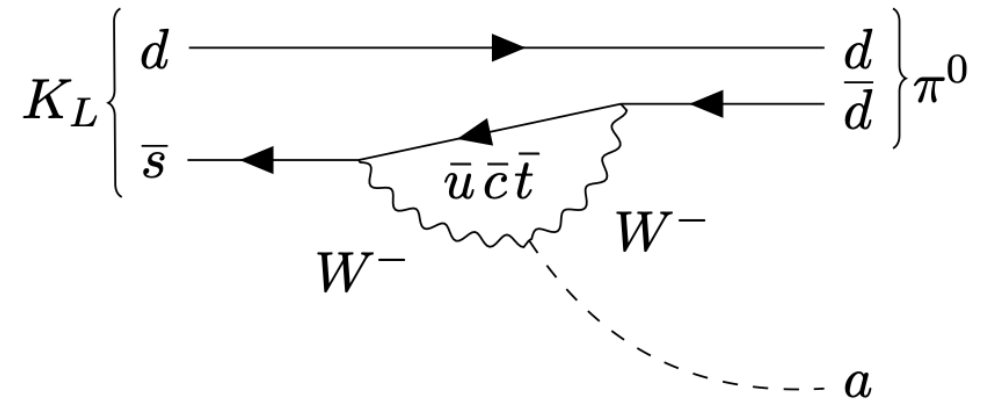
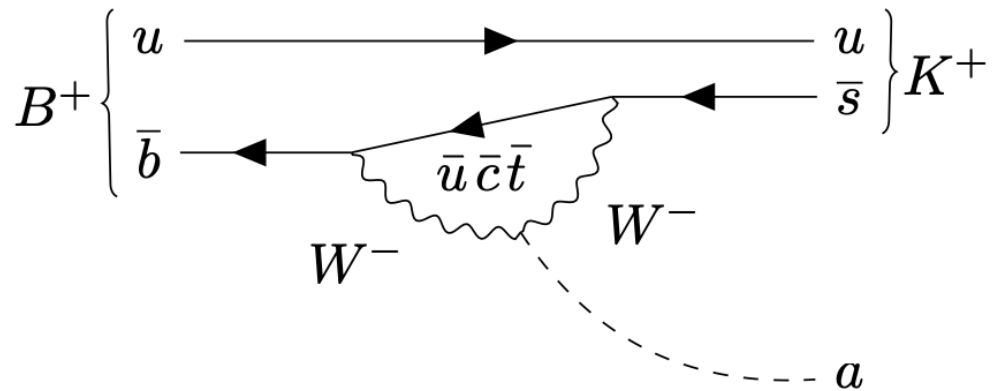
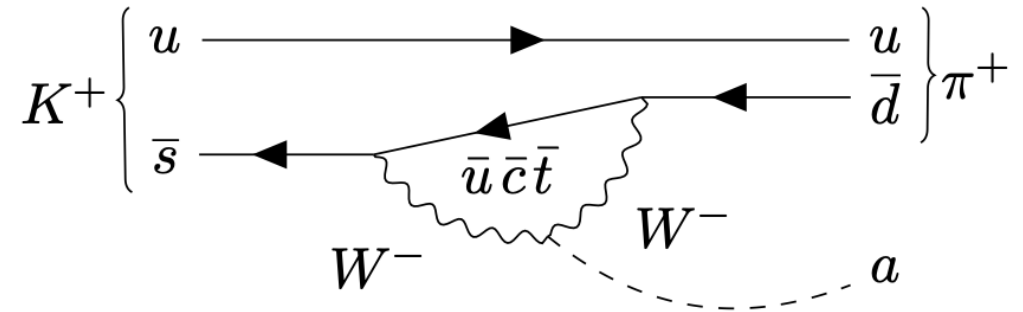
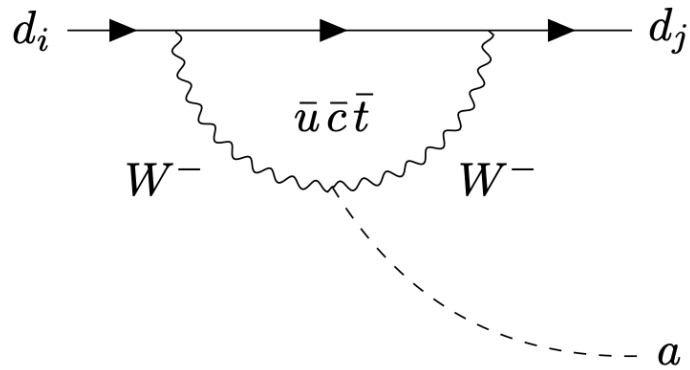
- Rates highly suppressed by  $\epsilon^2 \sim 10^{-10}$
- But still  $N_{A'} \sim 10^5$  per bin; LHC is a dark photon factory!

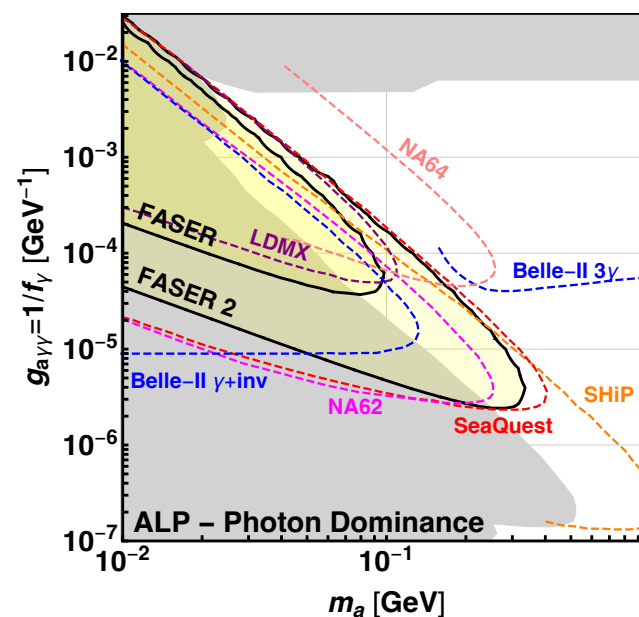
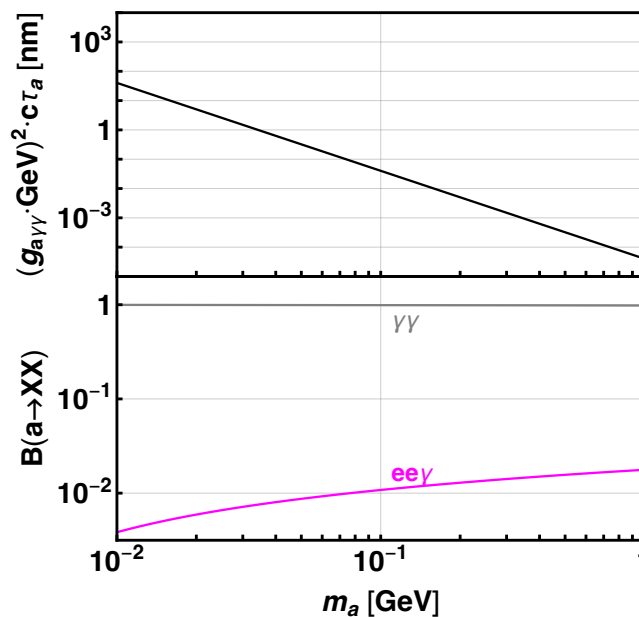
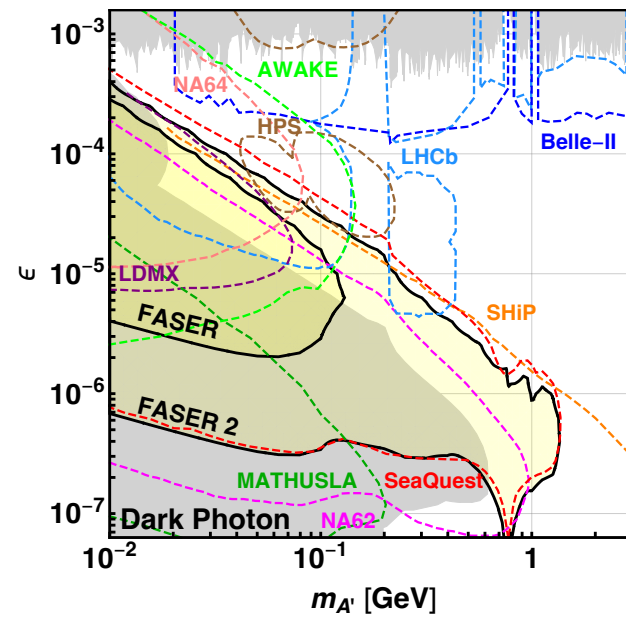
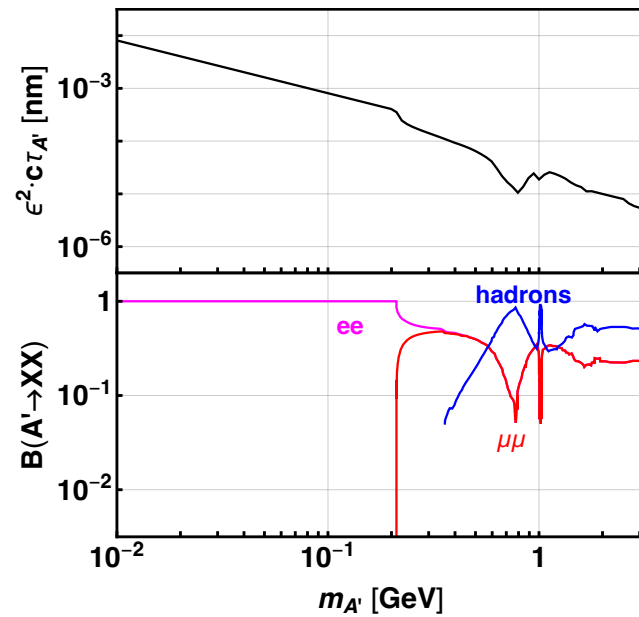
A's decay in FASER



- Rates suppressed again, but still  $N_{A'} \sim 100$  signal events
- Signal is  $E \sim \text{TeV}$  A's within 20 cm of the line of sight

# ALP production beyond primakoff







# Reach for Heavy Neutral Leptons

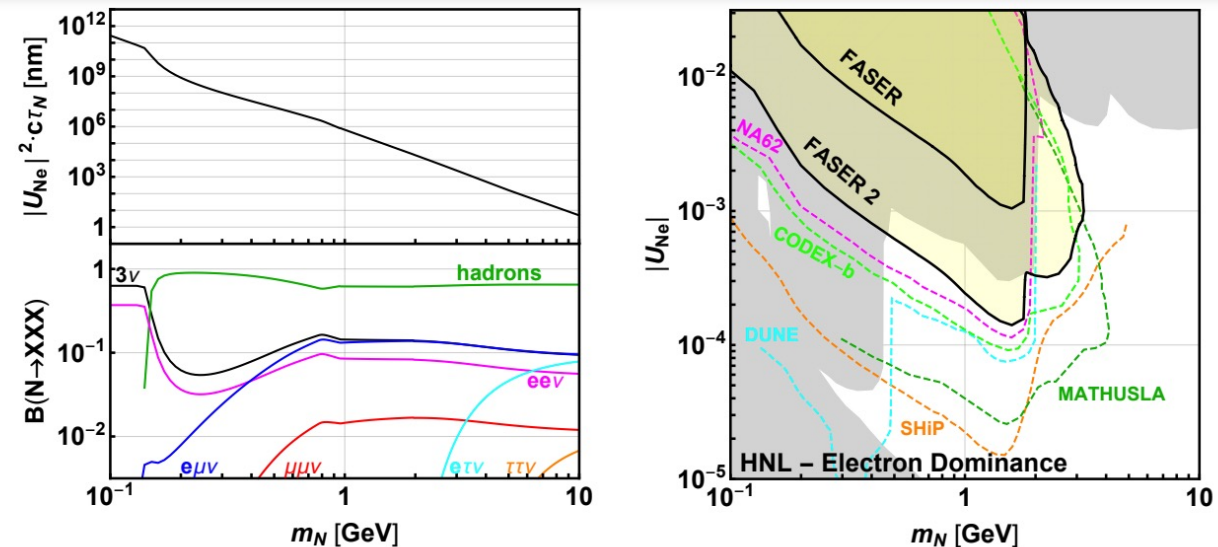


FIG. 12. **Benchmark Model F1.** The decay length (top left panel), decay branching fractions (bottom left panel), and **FASER's reach** (right panel) for the HNL that mixes only with the electron neutrino  $\nu_e$ . The gray shaded regions are excluded by current limits, and the colored contours are the projected sensitivities for other proposed experiments. See the text for details.

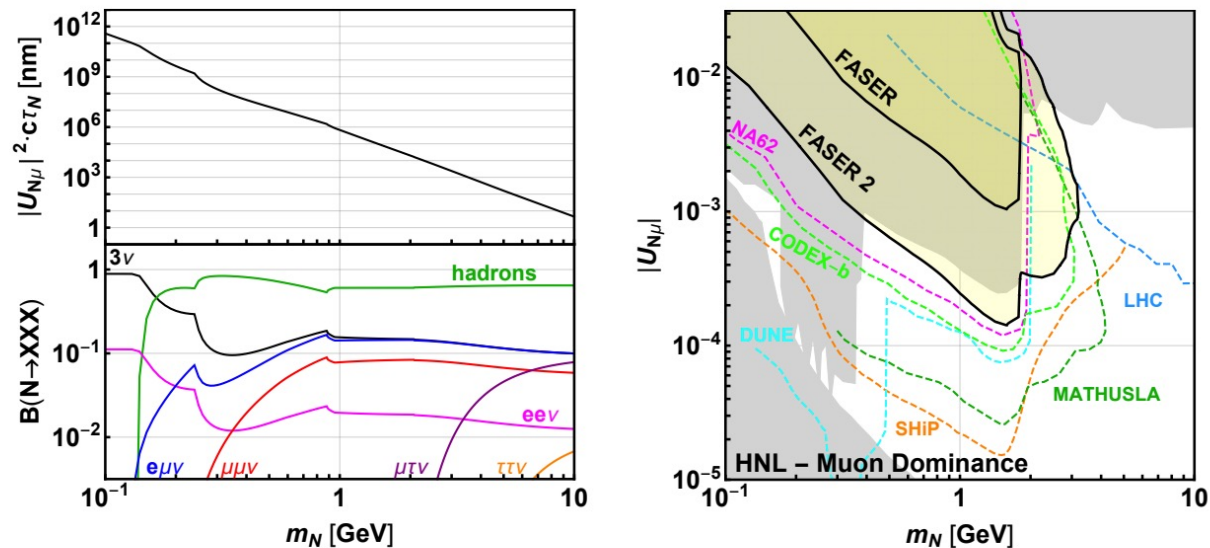


FIG. 13. **Benchmark Model F2.** As in Fig. 12, but for an HNL that only mixes with  $\nu_\mu$ .

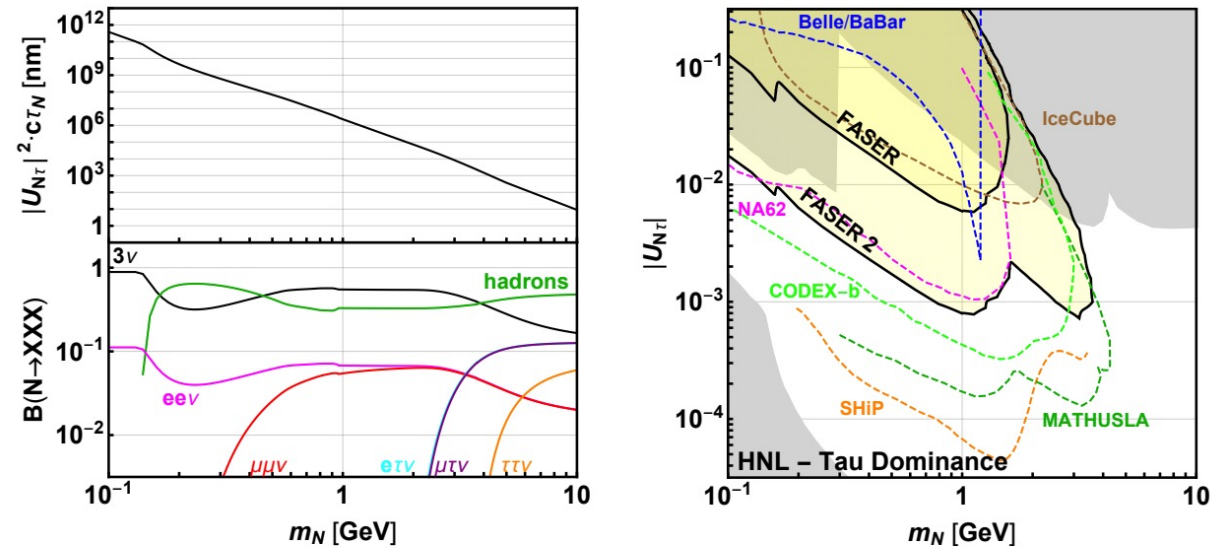
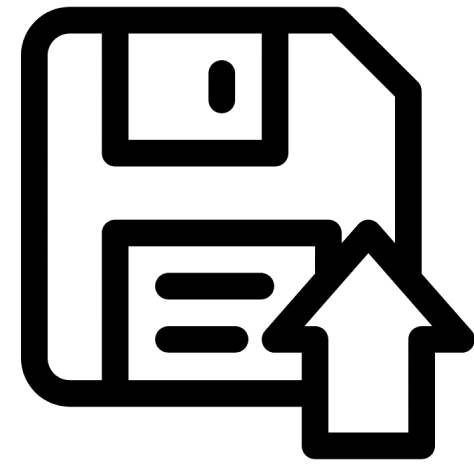
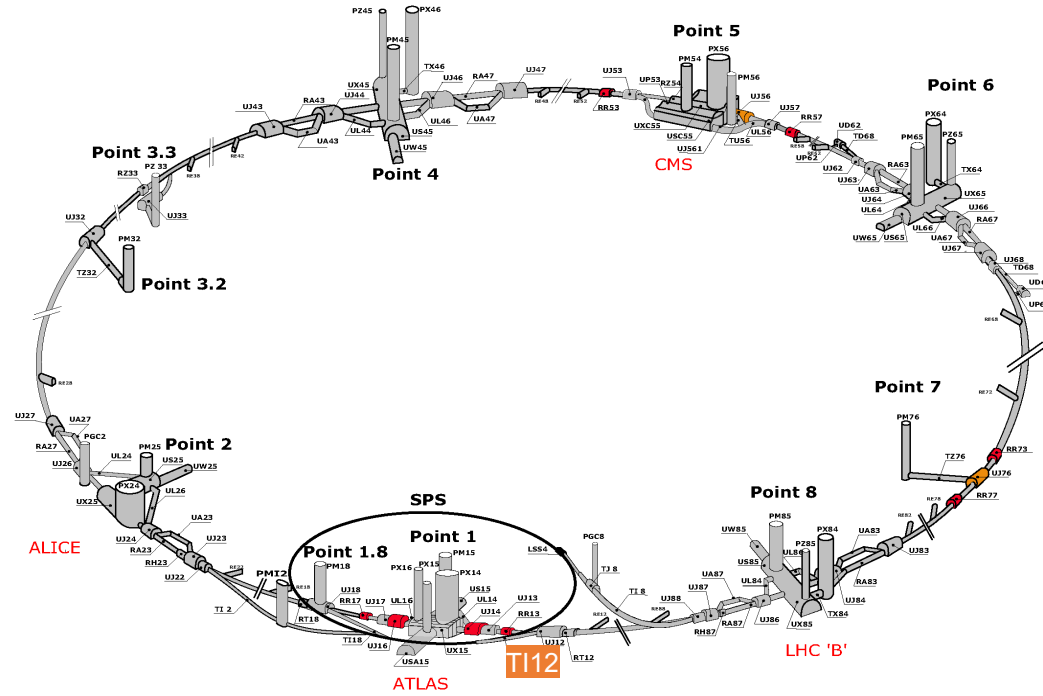


FIG. 14. **Benchmark Model F3.** As in Fig. 12, but for an HNL that only mixes with  $\nu_\tau$ .

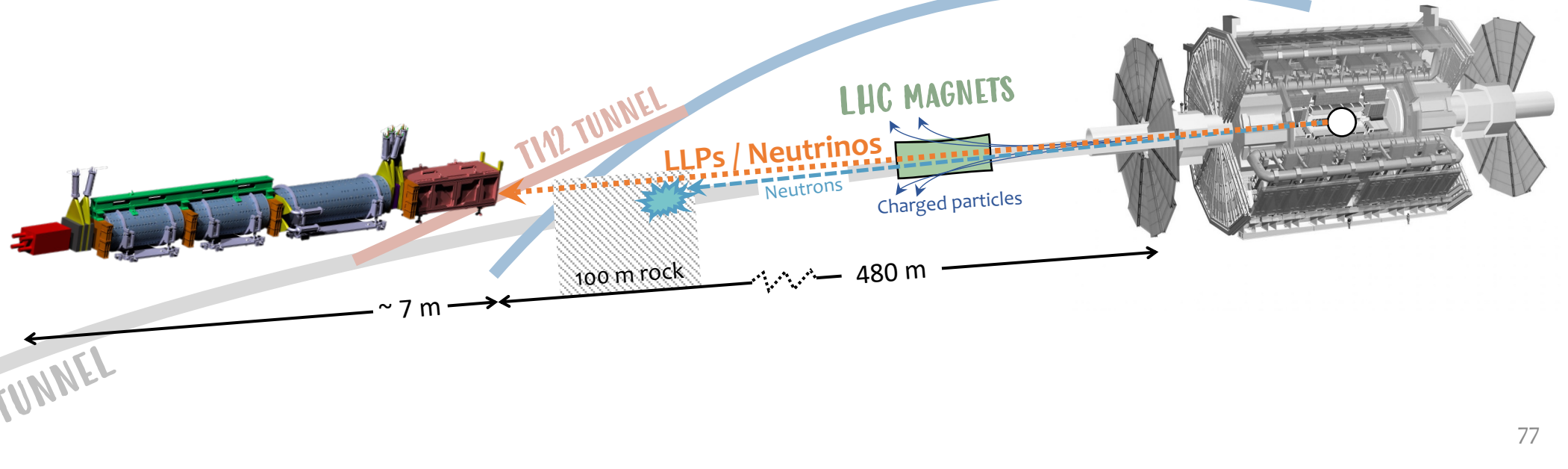


# THE *FASER* DETECTOR

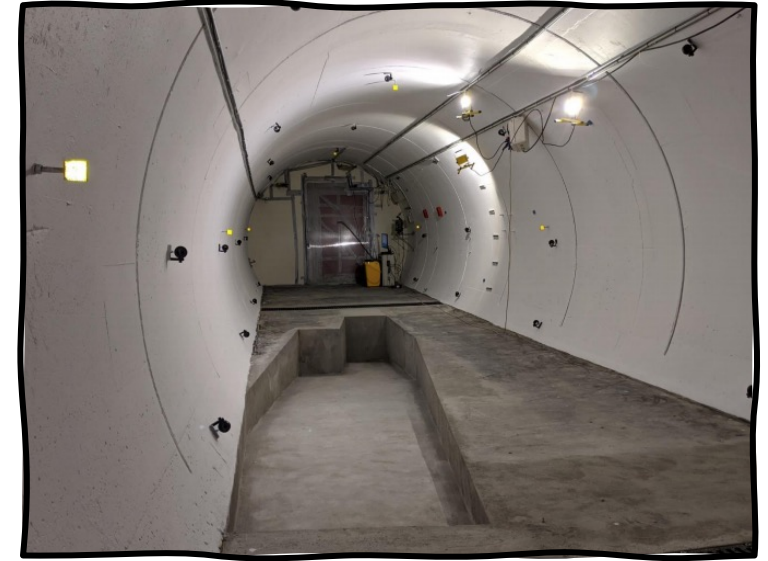
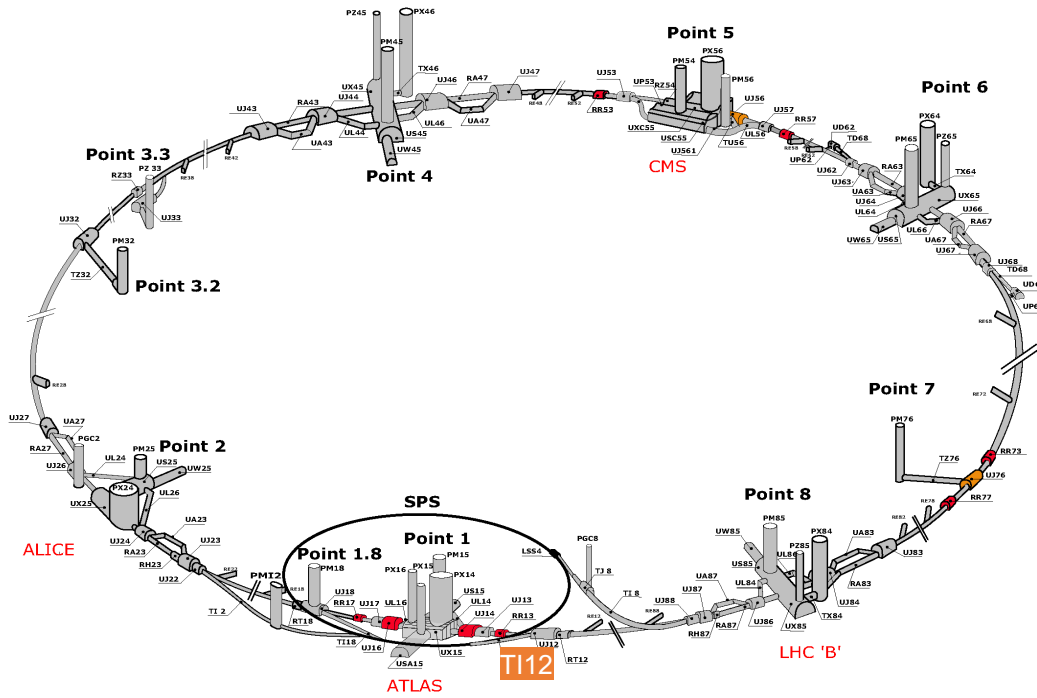
# LOCATION



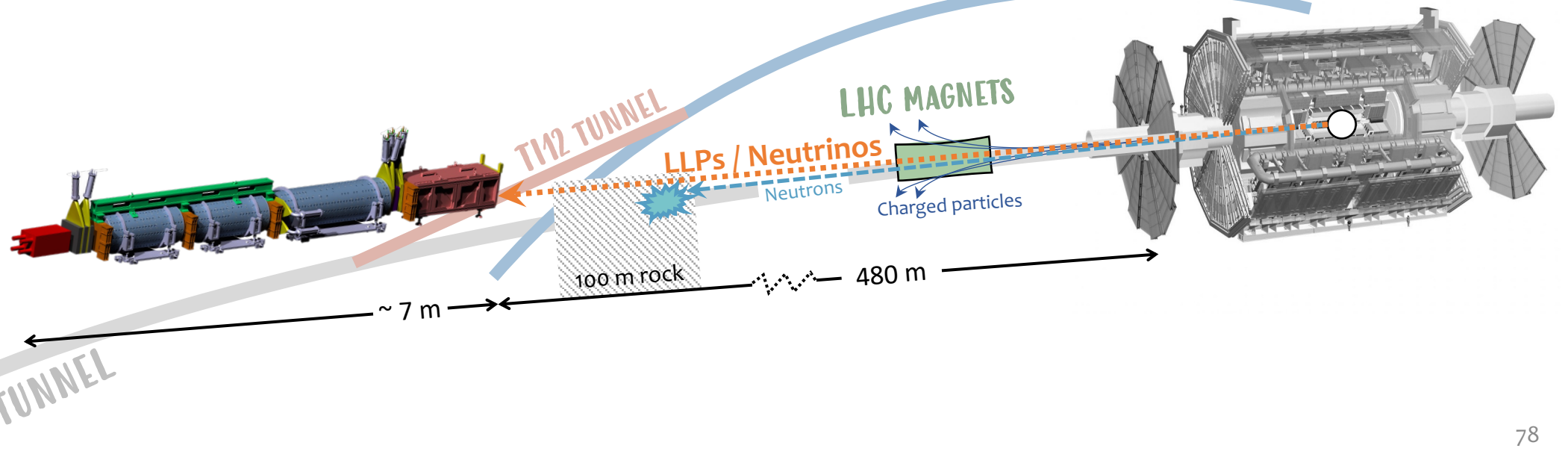
SPS TUNNEL



# LOCATION

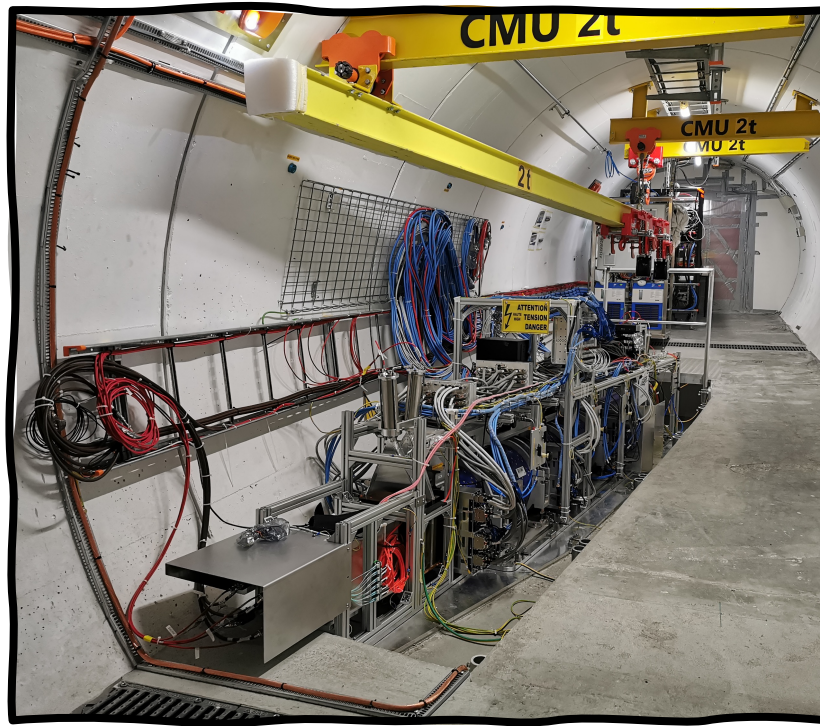


SPS TUNNEL

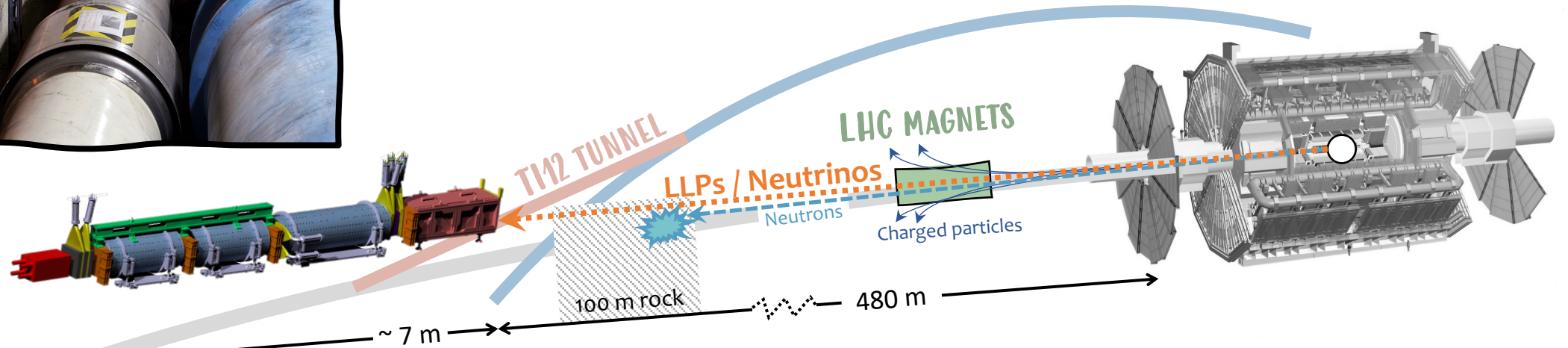




# LOCATION



SPS TUNNEL

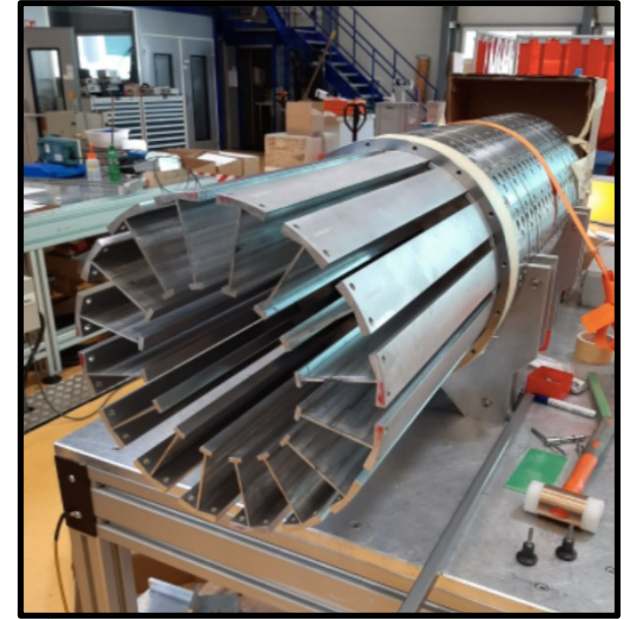
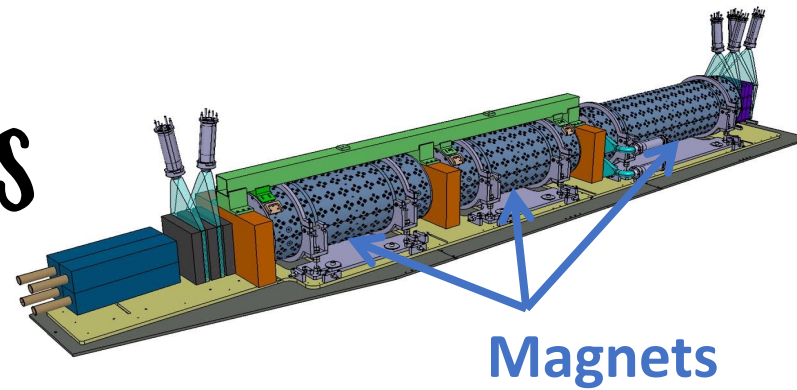


LHC TUNNEL

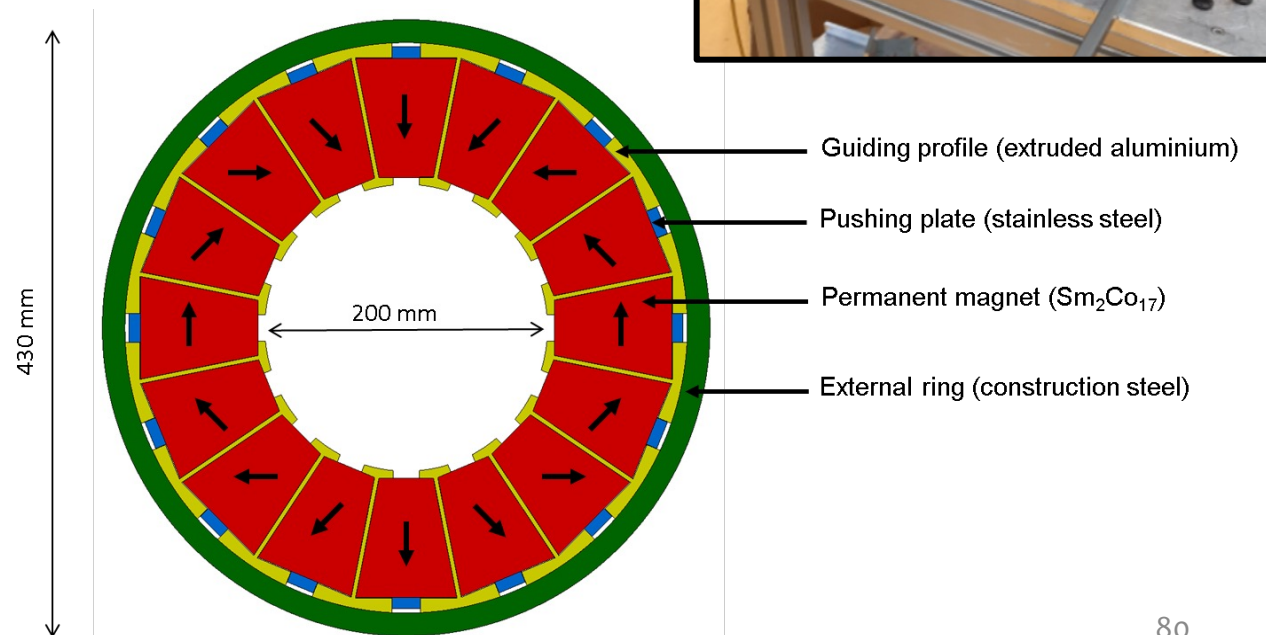
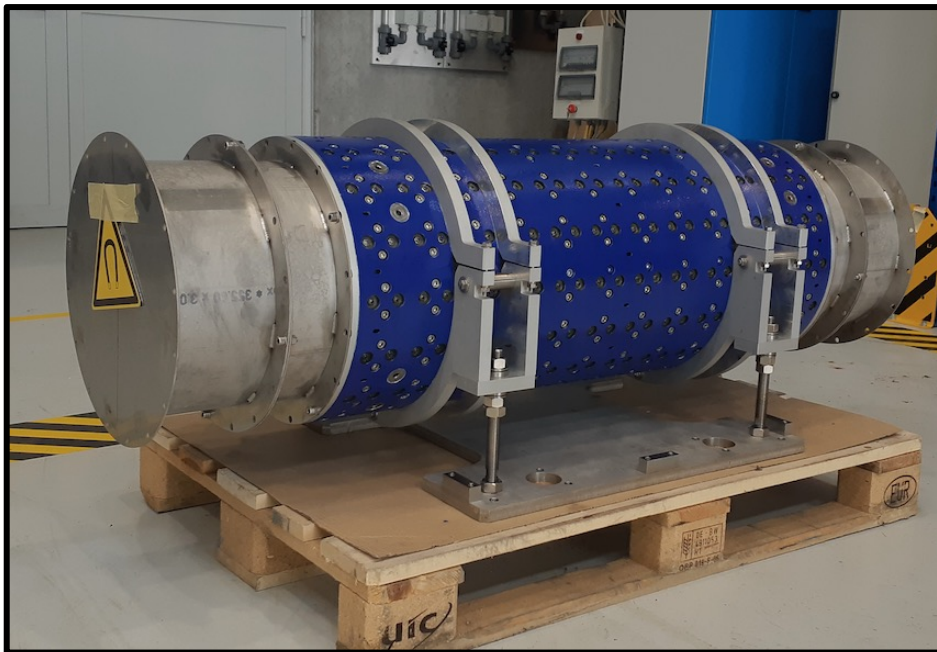




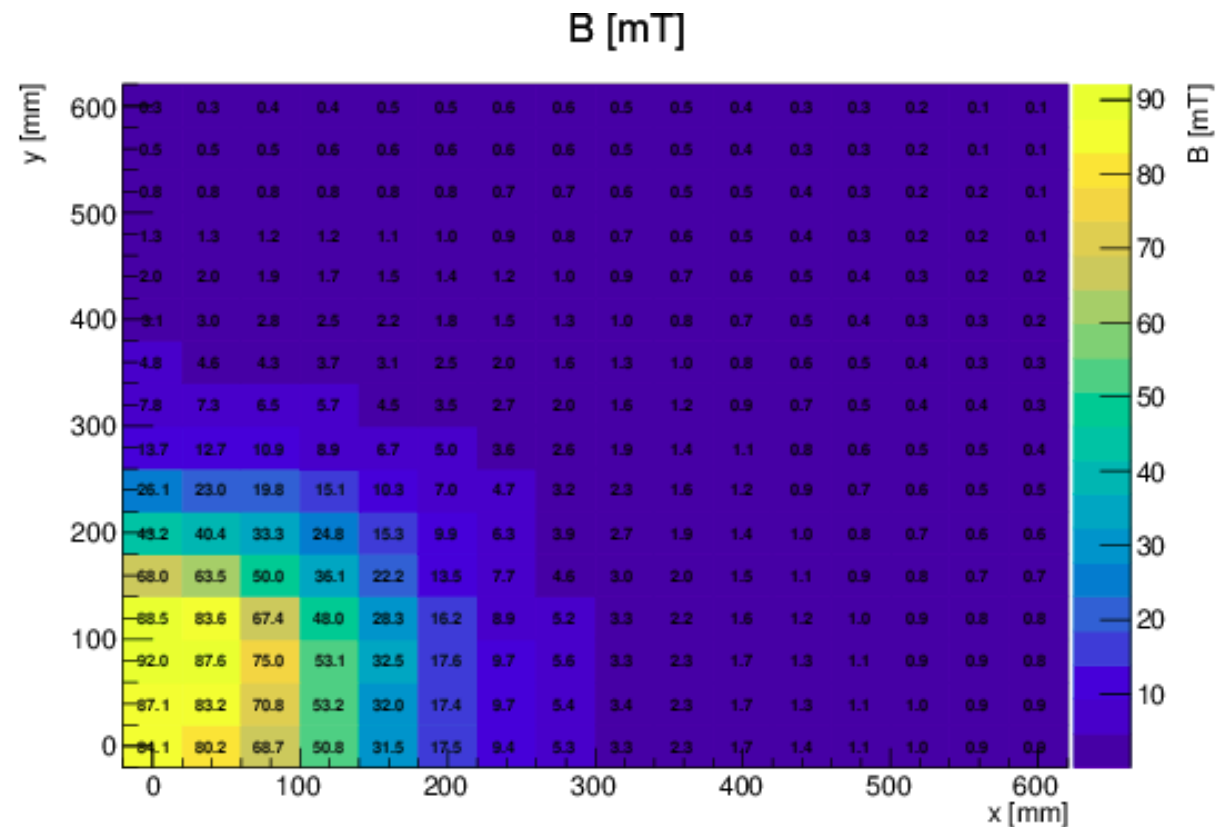
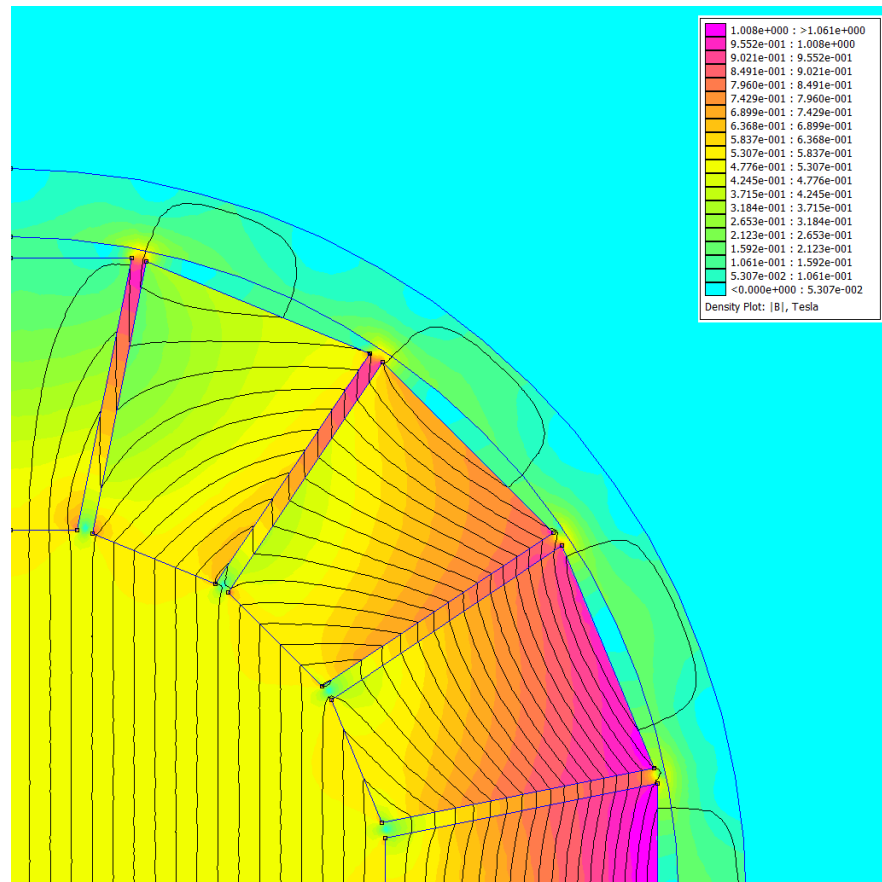
# MAGNETS



- Field of 0.55 T; permanent dipole
- Halbach array design with fixed-field magnets
  - Maximizes field without need for too much support infrastructure
  - Allows for a compact design, reducing amount of digging



# Magnets

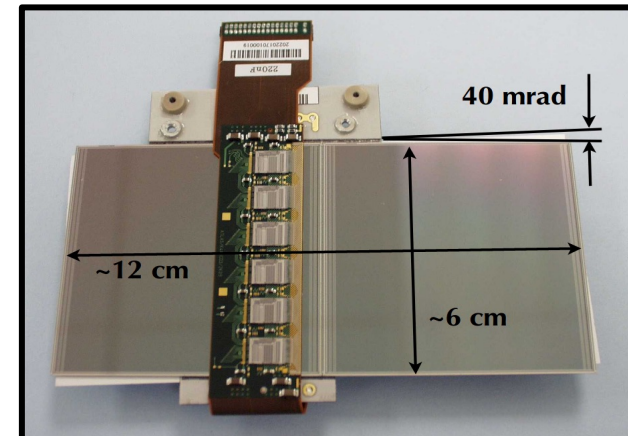
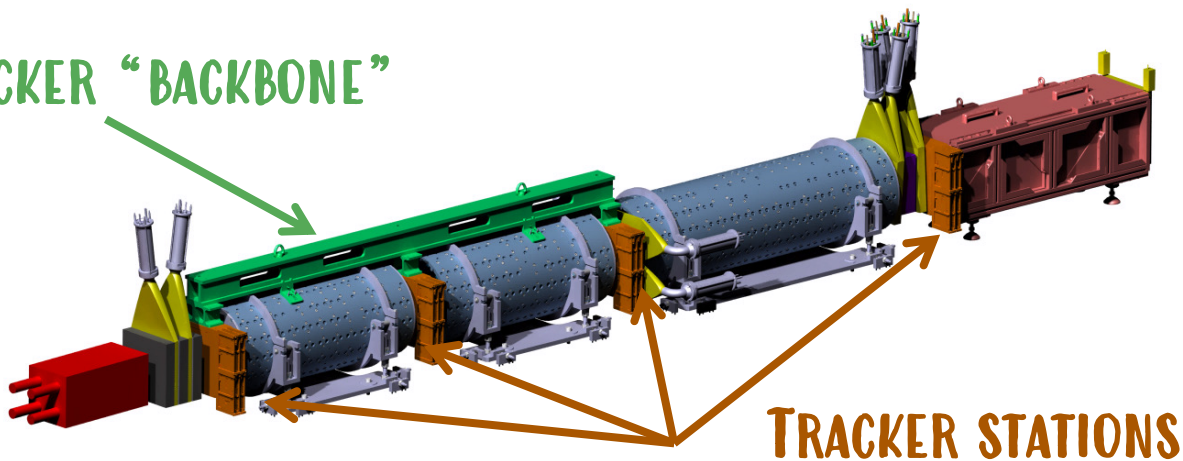


# TRACKER

THANKS!

- FASER uses ATLAS SCT spare modules
- 4 tracker stations x 3 tracker layers x 8 modules
  - 96 modules and  $O(10^5)$  channels in total
- Mechanical stability by “backbone” fixed on magnets
- Read out with custom GPIO board

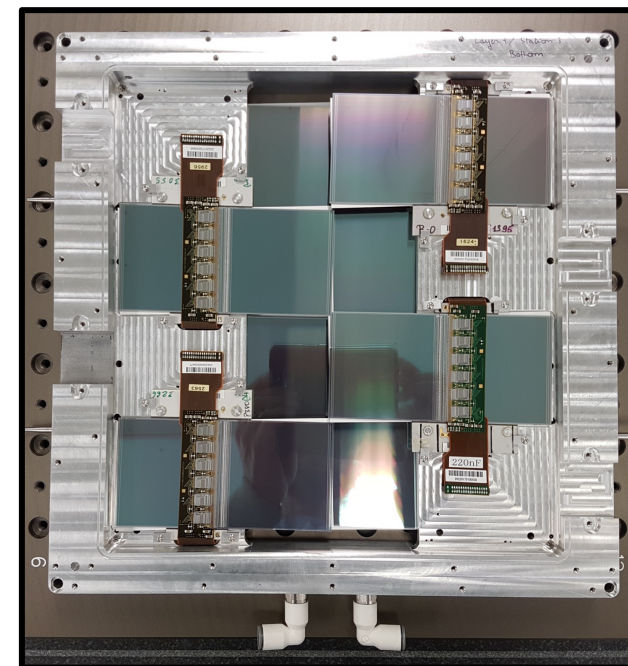
TRACKER “BACKBONE”



SCT module

80  $\mu\text{m}$  strip pitch / 40 mrad angle  
17  $\mu\text{m}$  / 580  $\mu\text{m}$  track resolution

Tracker layer



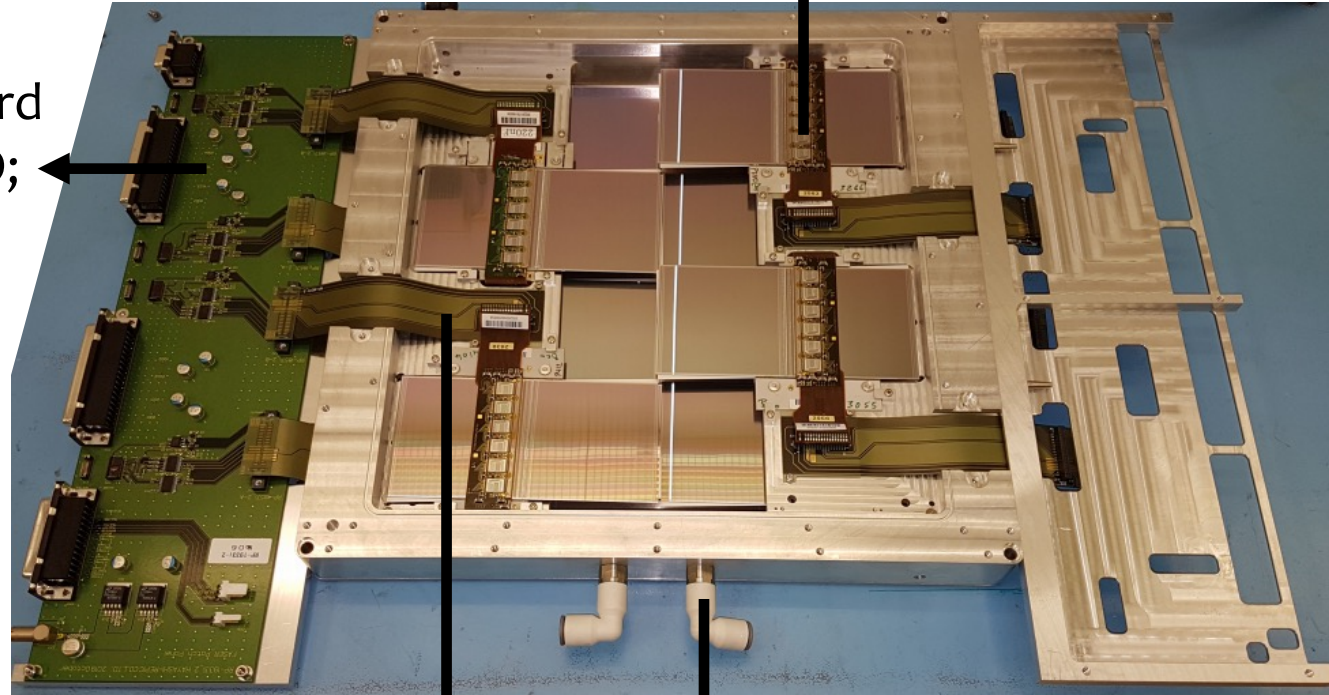


# TRACKER

SCT module ASICs,  
require  $\sim 5 \text{ W / module}$

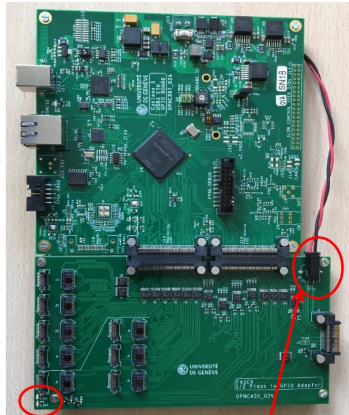
Low radiation in  
TI12 and much  
lower rates  
than ATLAS  
allow for  
simplifications  
in services and  
readout.

Patch panel to custom board  
based on home-made GPIO;  
Power (HV/LV), monitoring  
and readout lines.

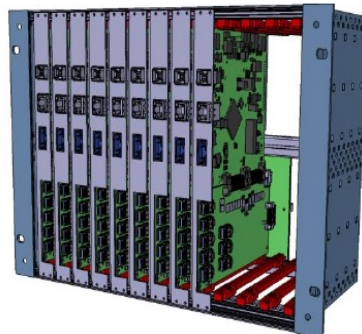


FLEX cables

Detector cooling via water  
chiller operating at  $10\text{-}15^\circ\text{C}$



2 front panels LEDs  
24V discrete wire to  
TRB adapter



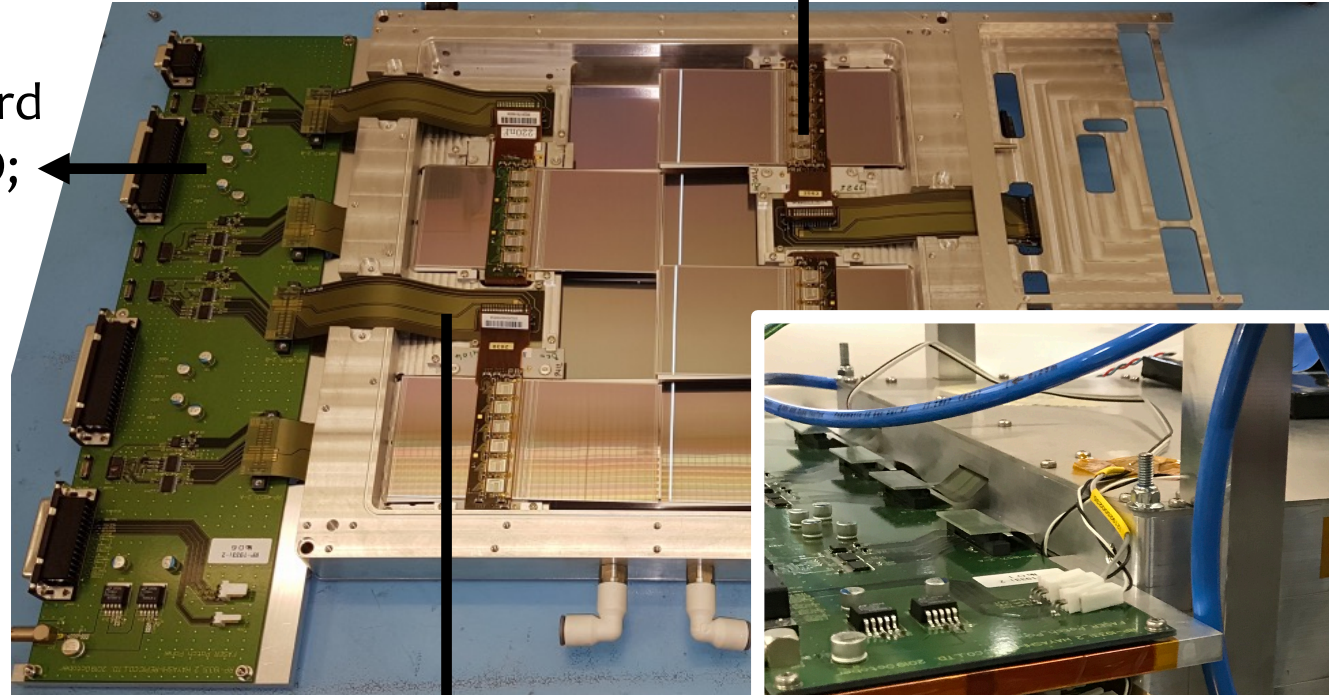
Into  
custom-made  
mini-crate



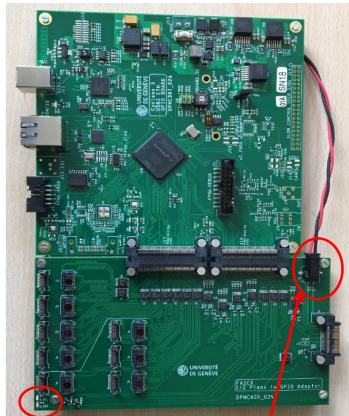
# TRACKER

SCT module ASICs,  
require  $\sim 5 \text{ W} / \text{module}$

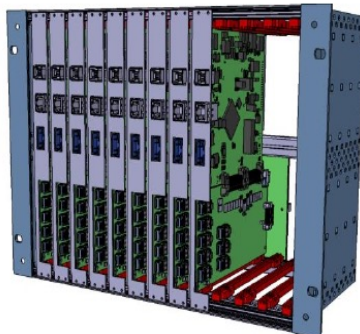
Patch panel to custom board  
based on home-made GPIO;  
Power (HV/LV), monitoring  
and readout lines.



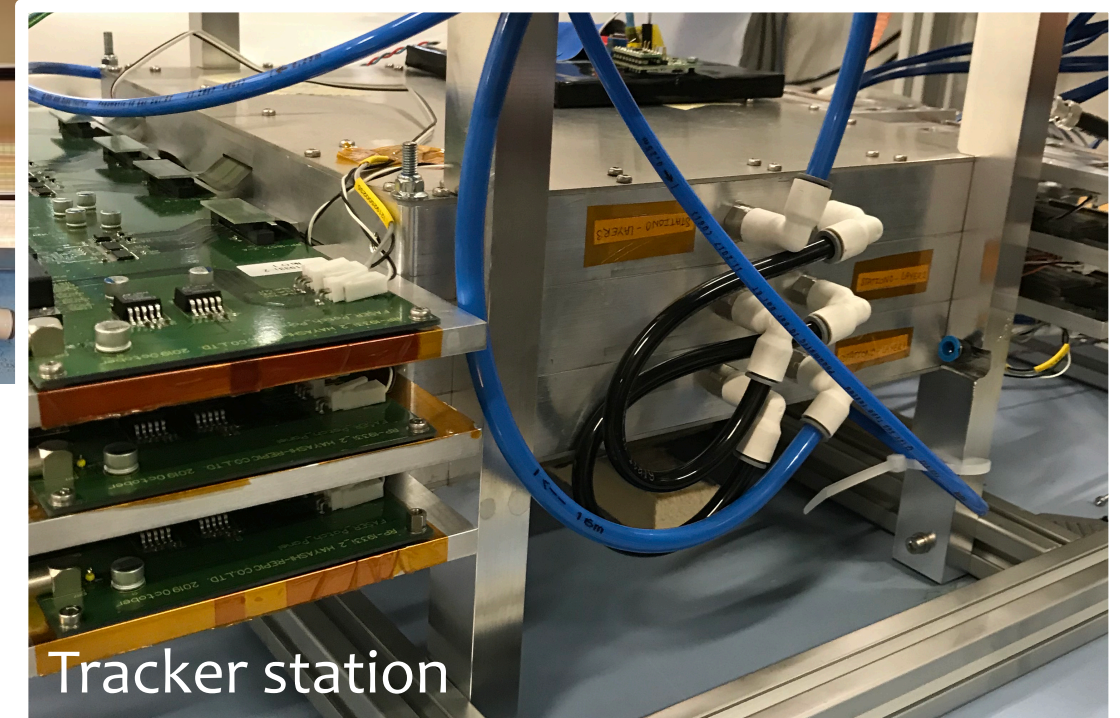
FLEX cables



2 front panels LEDs  
24V discrete wire to  
TRB adapter



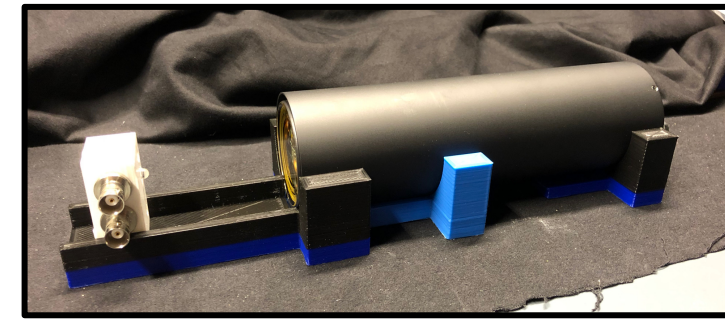
Into  
custom-made  
mini-crate



Tracker station

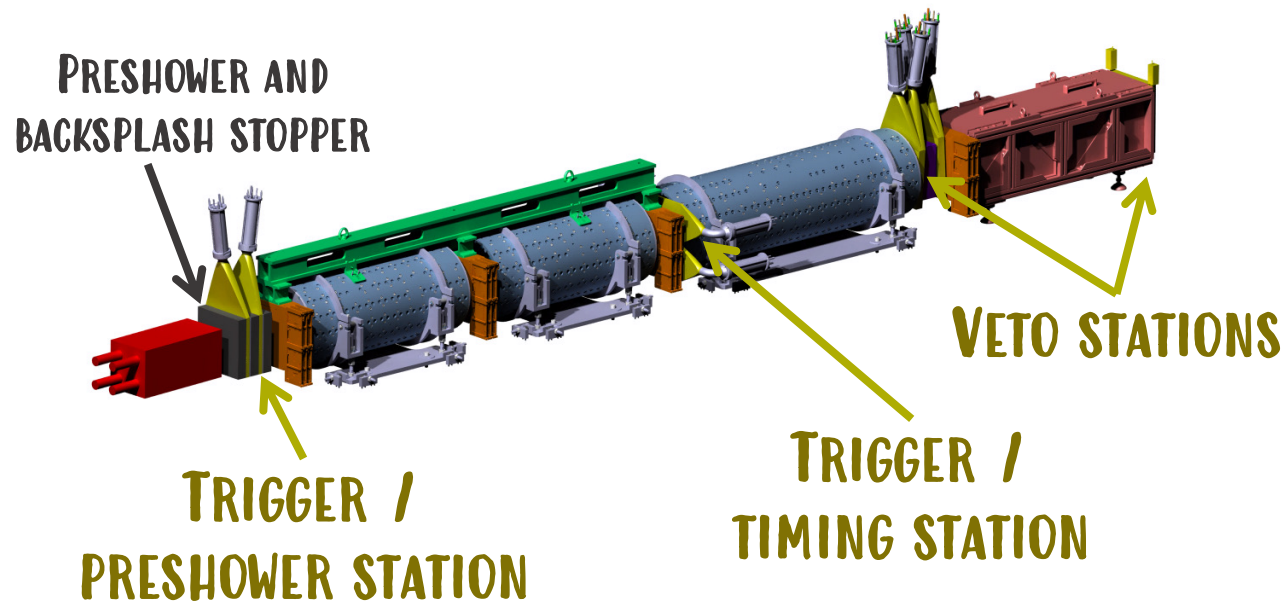


# SCINTILLATORS



Scintillator PMTs

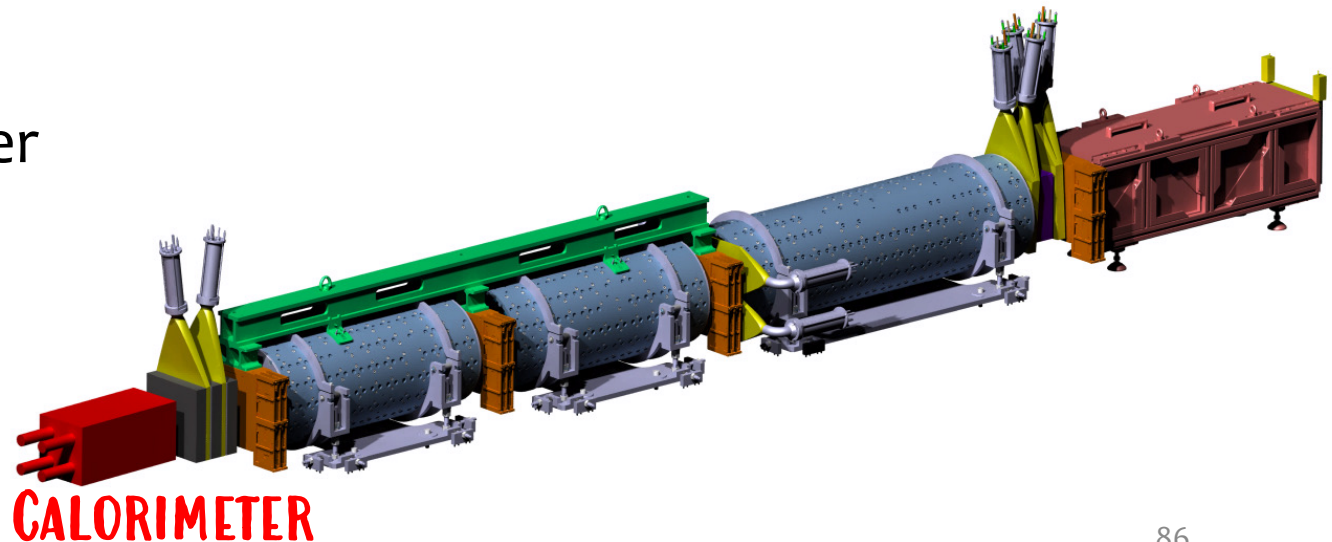
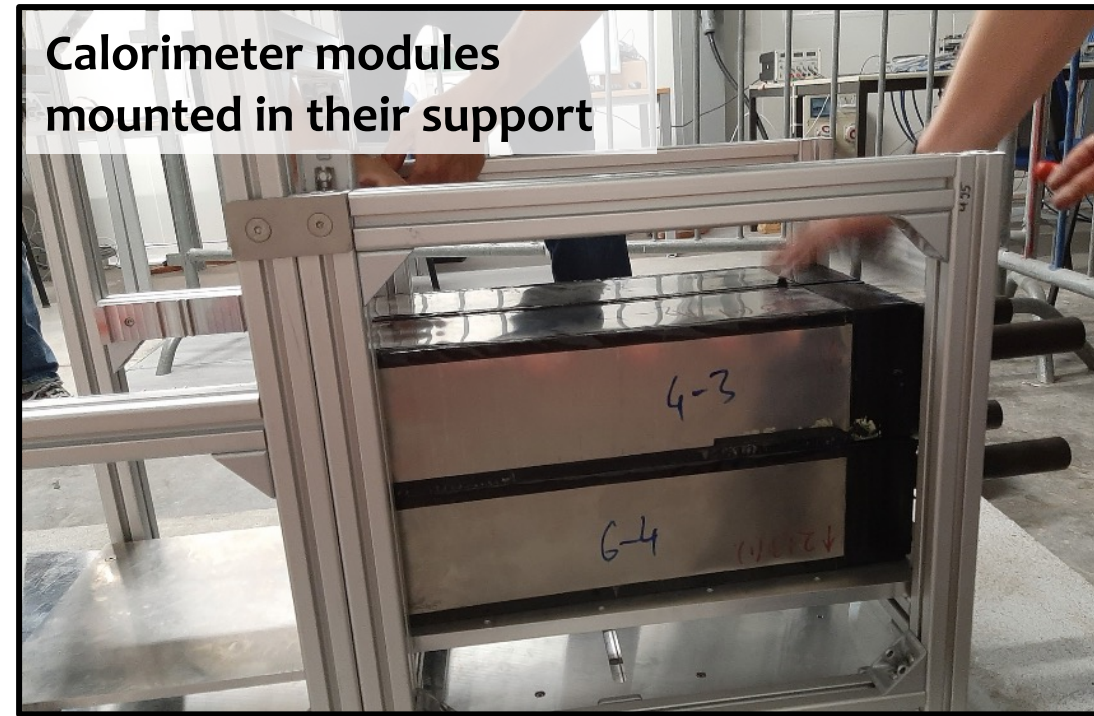
- Three stations all providing triggering capability:
  - Very high efficiency veto station for incoming charged particles (x6 planes)
  - Timing station; precise timing ( $\sim$  ns) wrt IP (x1 plane)
  - Pre-shower station; coincidence with timing station (x2 planes)
- Read out with PMTs and CAEN digitizer



# CALORIMETER

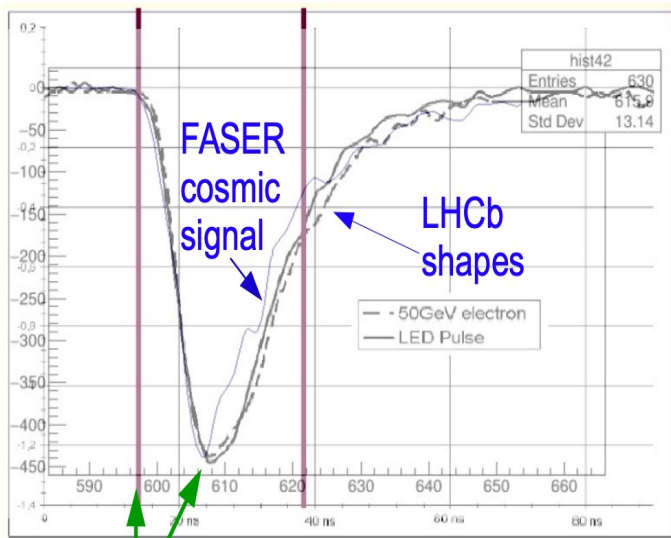
THANKS!

- FASER uses 4 LHCb spare outer ECAL modules
  - 25 radiation lengths long
  - Lead/scintillator calorimeter
- Energy resolution  $\sim 1\%$  for TeV deposits
  - No longitudinal shower information
- Provides triggering capability
- Read out with PMTs and CAEN digitizer

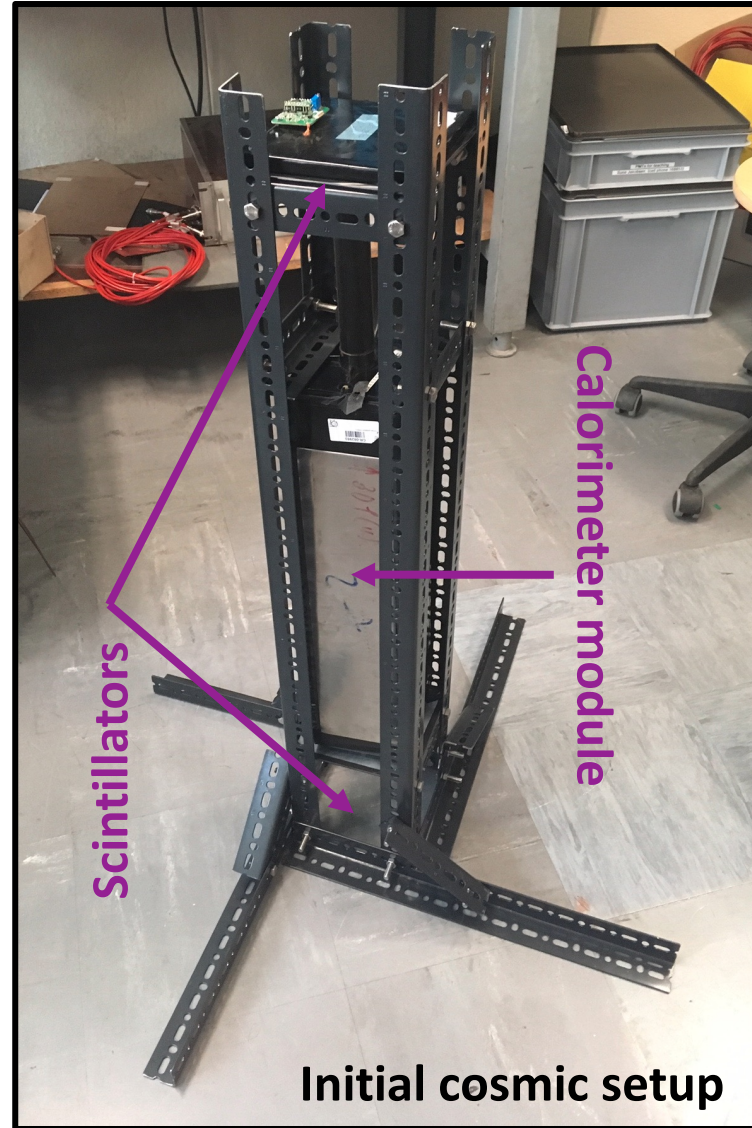




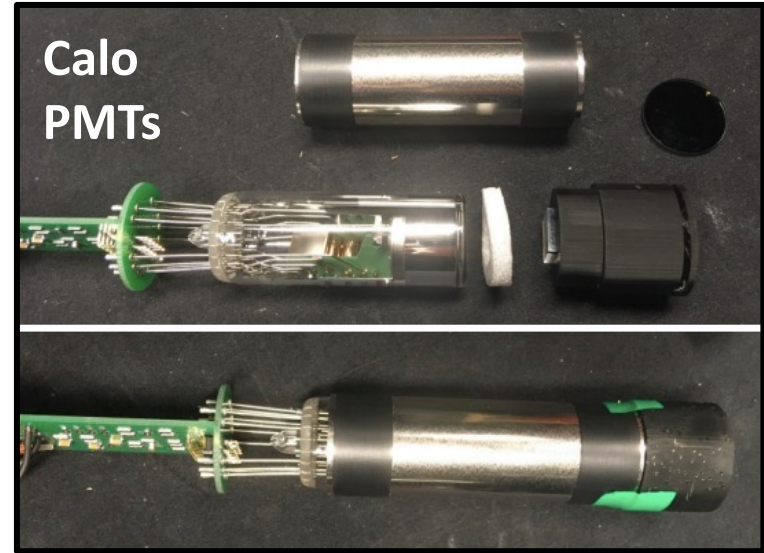
# Calorimeter



~10 ns rise time

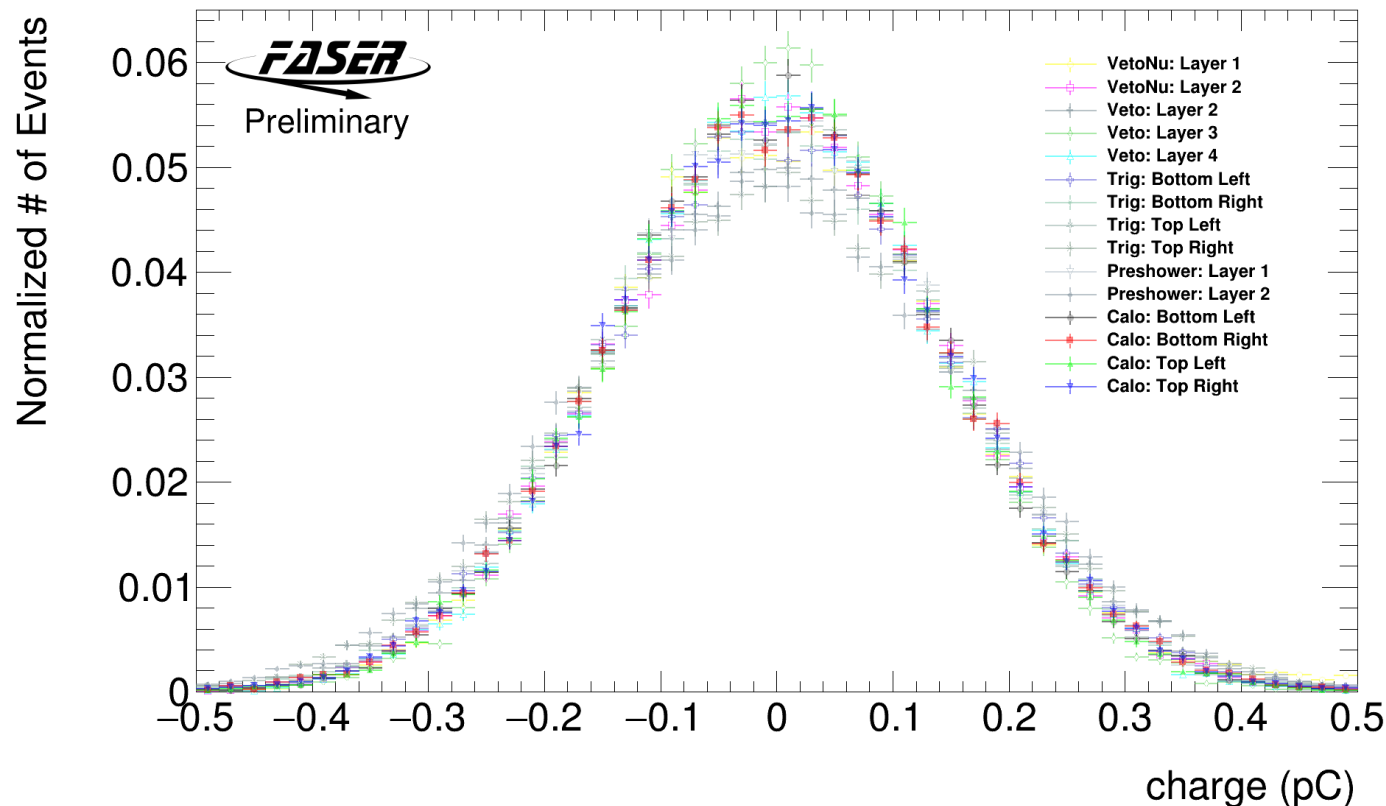


Initial cosmic setup



# NOISE DISTRIBUTIONS FOR ALL SCINTILLATOR AND CALO CHANNELS

- The pedestal subtracted charge distributions of randomly triggered events are shown for all scintillators and calorimeter modules. The charge is derived from the integration of the waveform over the standard 120 ns reconstruction window. Normalization of the distributions are done by dividing by the total number of events. The plot shows that the noise levels are similar across all scintillators and calorimeter channels, regardless of different PMT types and HV settings. Dominated by the digitizer noise, the total noise of each sub detector falls within the range of  $0.15 \pm 0.02$  pC.



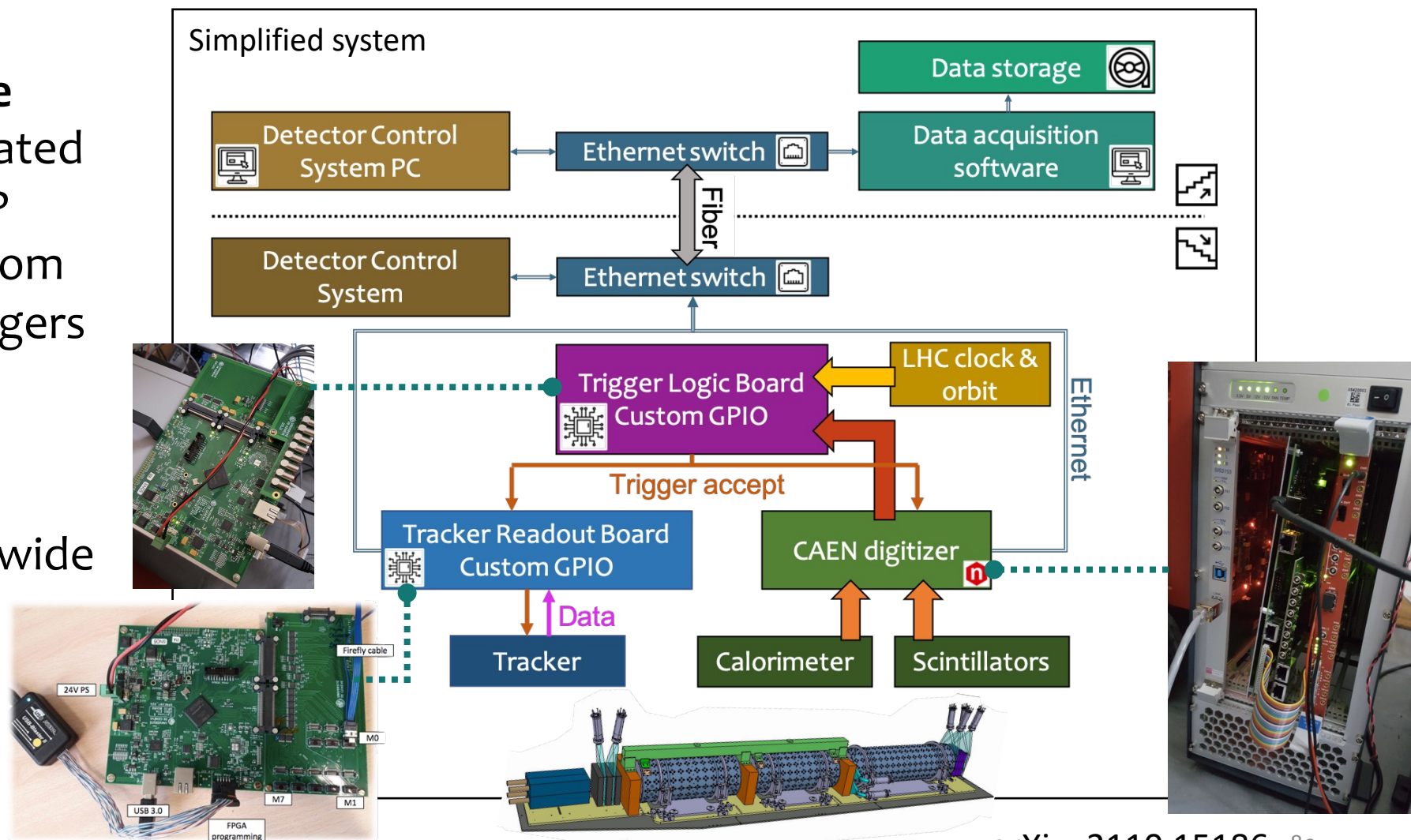


# TRIGGER & DATA ACQUISITION

- Expected **trigger rate** about **500 Hz**, dominated by muons from the IP
  - L1A includes random and software triggers

- Expected **bandwidth** about **15 MB / s**, dominated by PMTs' wide signal ( $\sim 1 \mu\text{s}$ )

- All TDAQ electronics are placed in T112





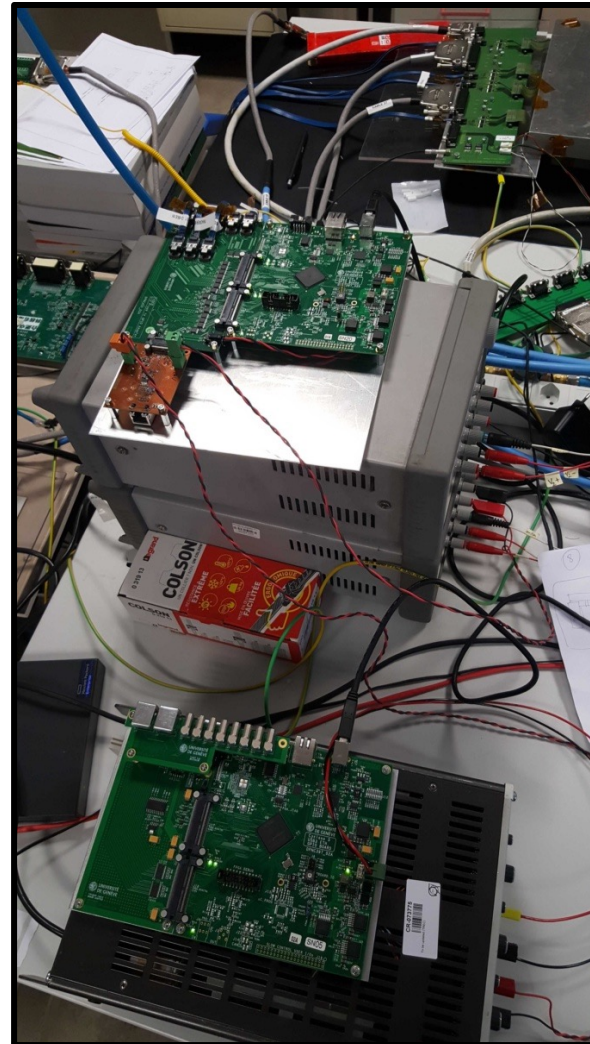


# Trigger & Data acquisition

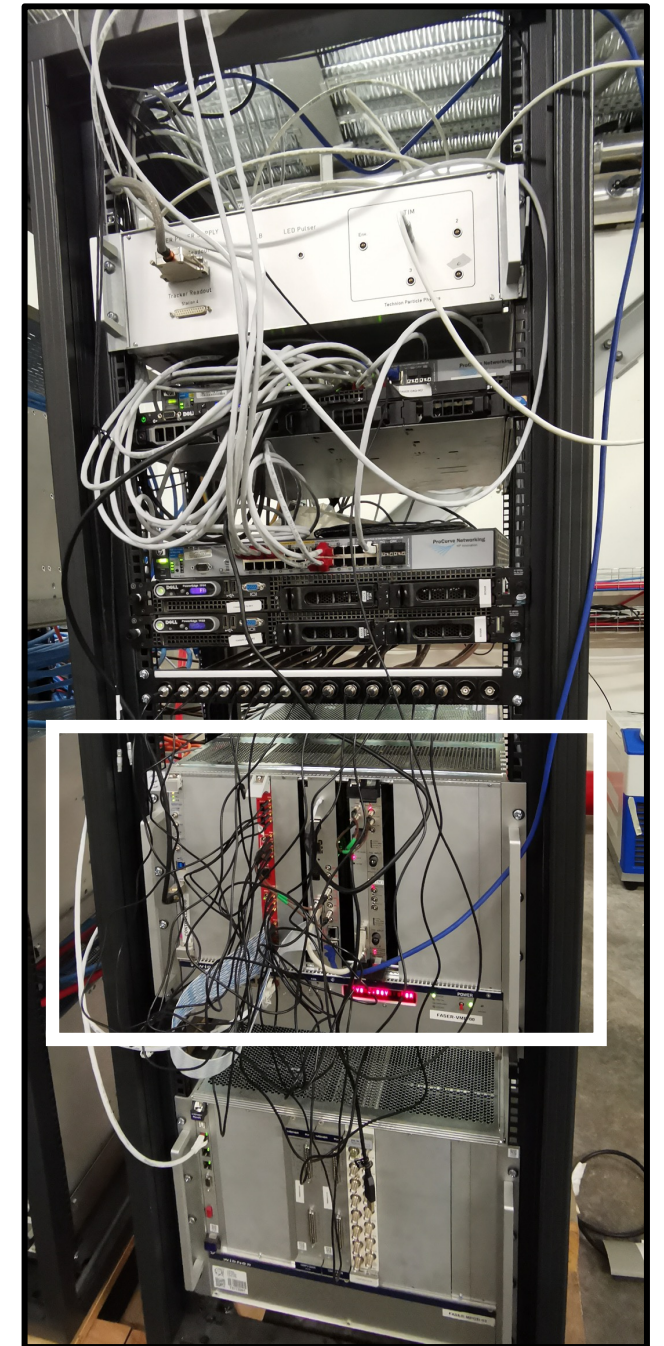


Initial Run Control application, produced by summer intern

- L1A includes random and software triggers
- Expected **bandwidth** about **15 MB / s**, dominated by PMTs' wide signal ( $\sim 1 \mu\text{s}$ )
- All TDAQ electronics will be placed in TI12

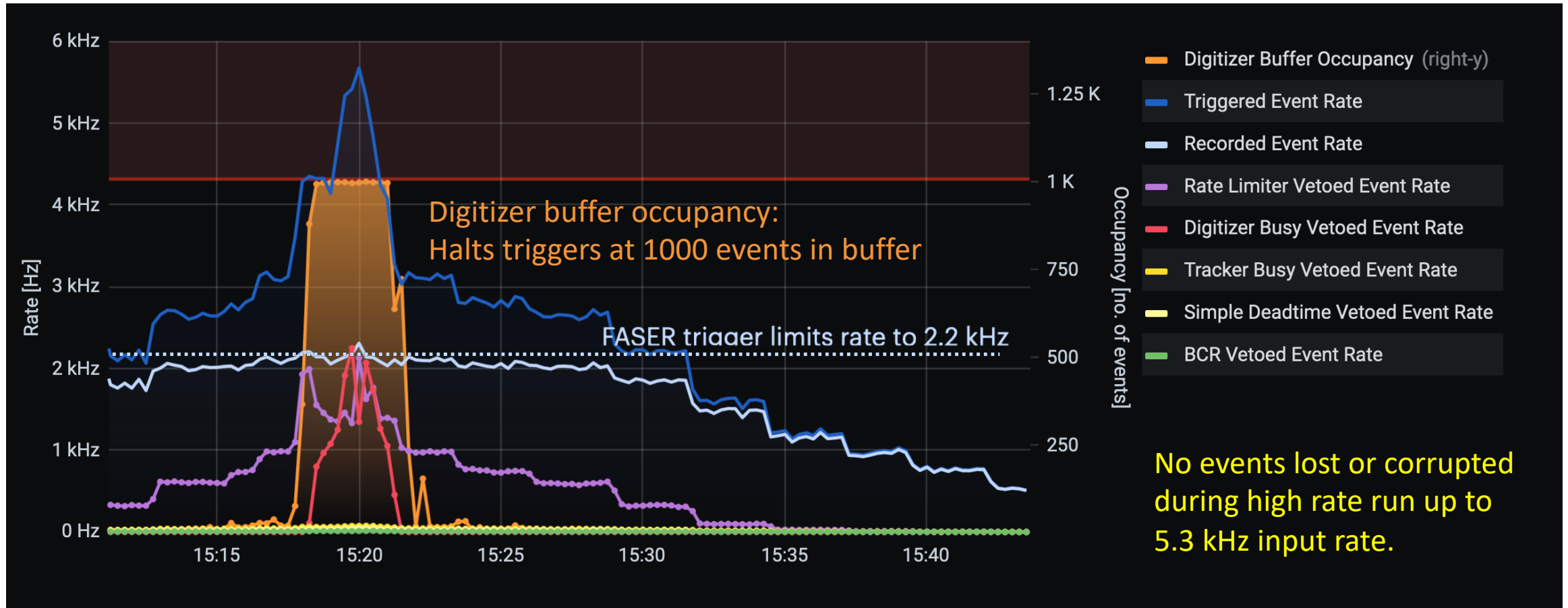


All boards connected together for tests

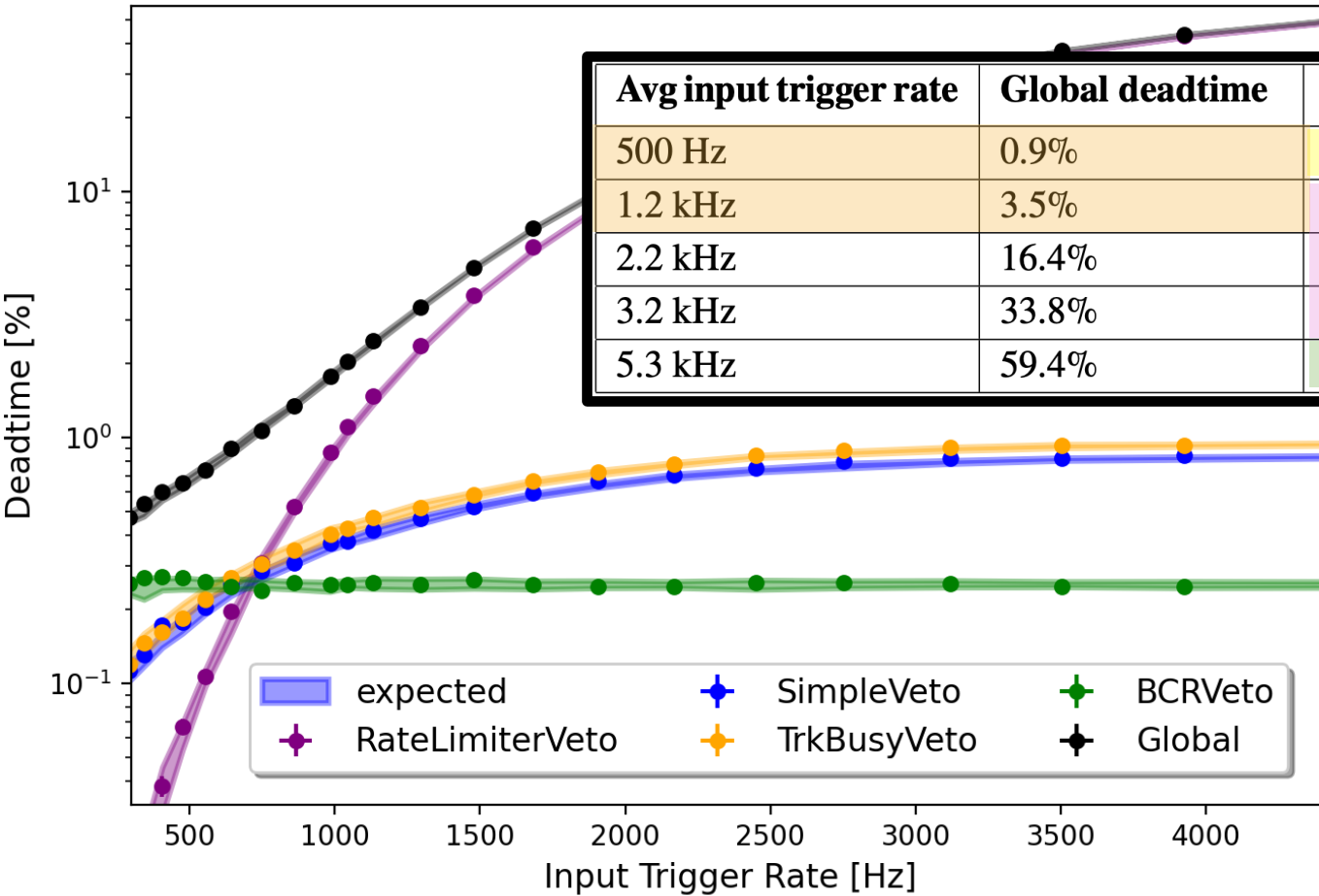


TDAQ boards in the final VME crate

# High rate tests



# Precise deadtime measurements



Avg input trigger rate	Global deadtime	Max deadtime source	Data throughput
500 Hz	0.9%	tracker	10 MBytes/s
1.2 kHz	3.5%	rate limiter	26 MBytes/s
2.2 kHz	16.4%	rate limiter	47 MBytes/s
3.2 kHz	33.8%	rate limiter	47 MBytes/s
5.3 kHz	59.4%	digitizer	47 MBytes/s

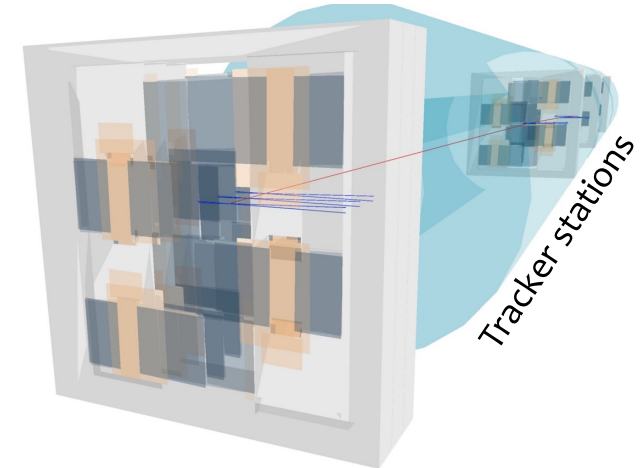
Expected Run3 rates

- Tracker only reads out one event at a time
- Rate is limited to 2.2 kHz
- Digitizer read-out limitations

At 500 Hz, deadtime expected to be < 1 %



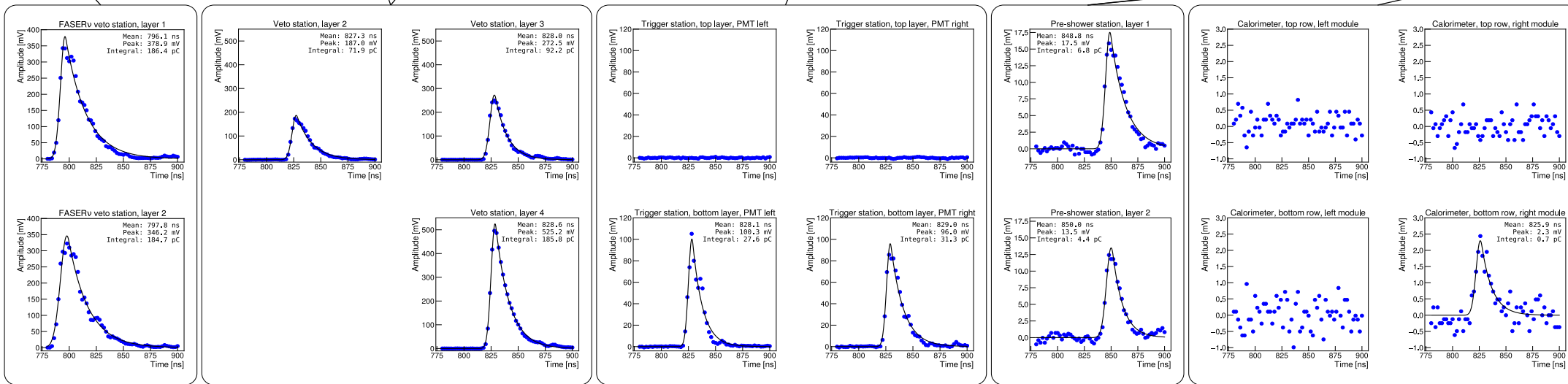
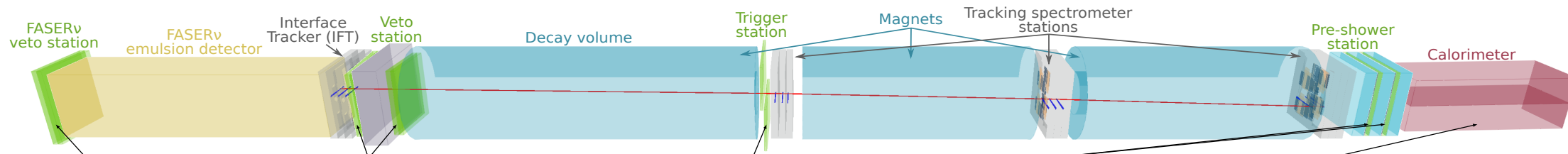
# EXAMPLE TYPICAL EVENT: MUON



Run 8336  
Event 1477982  
2022-08-23 01:46:15

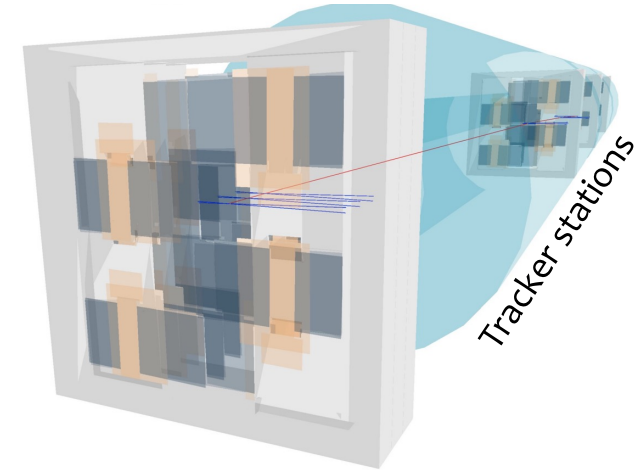
Collision event with muon traversing FASER

← To ATLAS IP





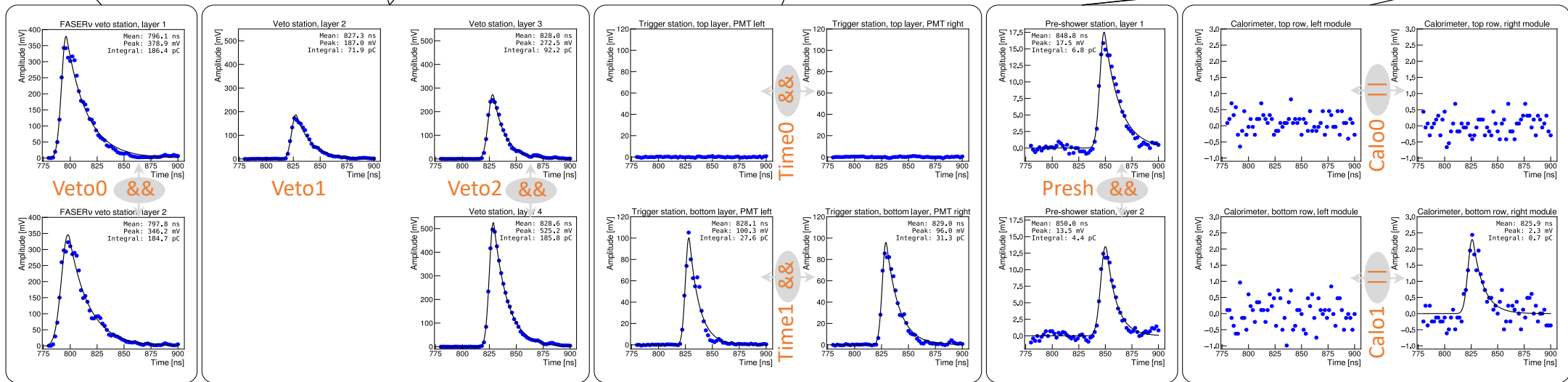
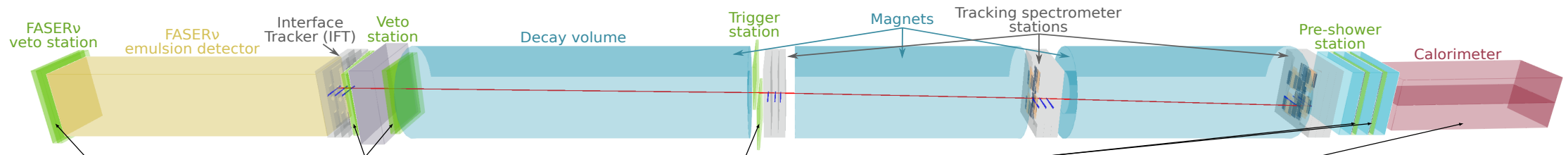
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Event 1477982  
2022-08-23 01:46:15

Collision event with muon traversing FASER

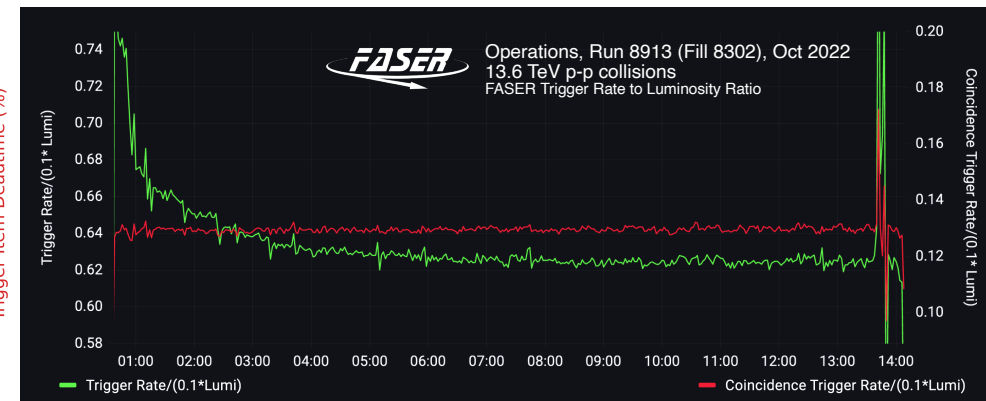
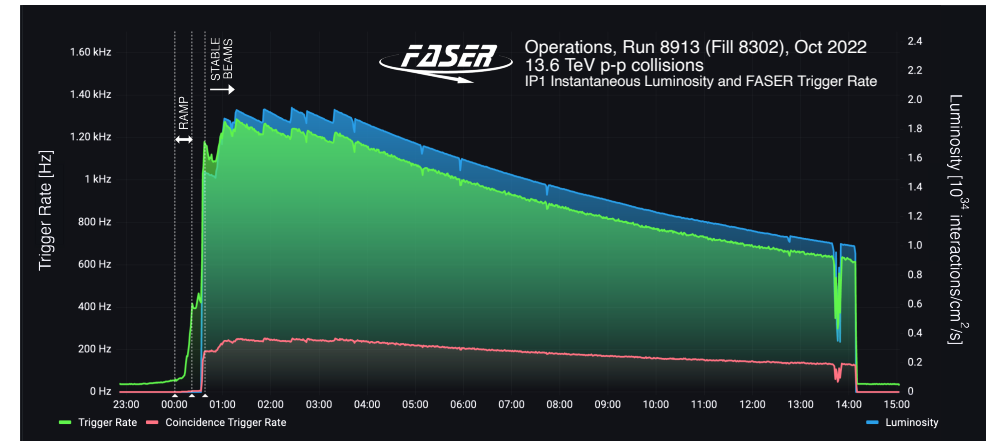
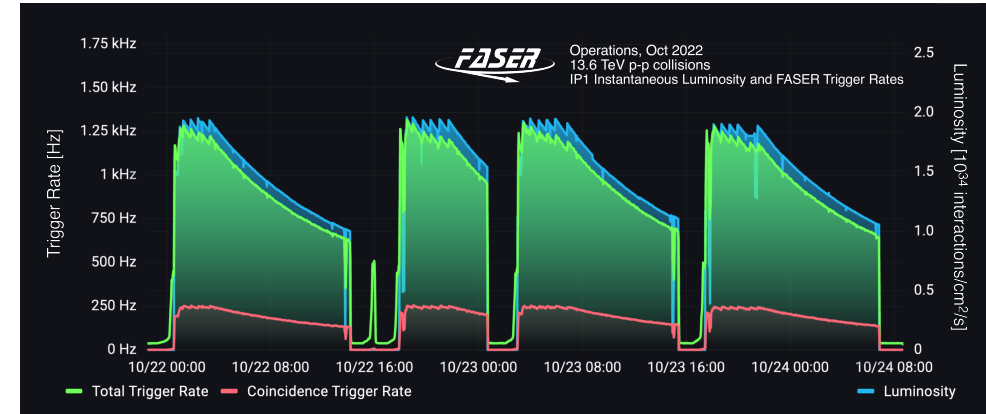
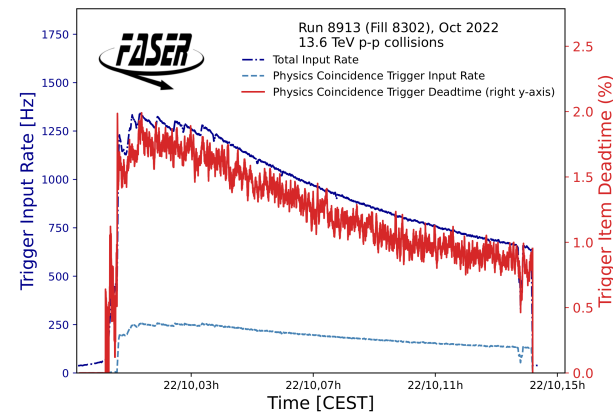
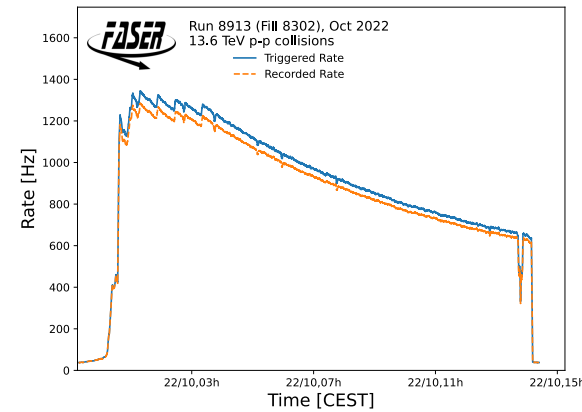
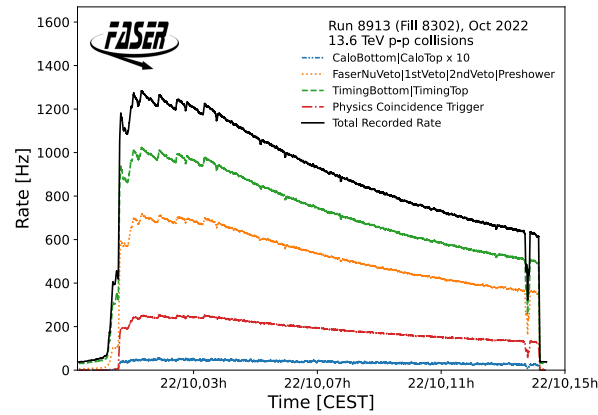
← To ATLAS IP



Trigger bits combined to trigger items -- e.g. "coincidence trigger" (incoming muons): (Veto0 || Veto2) && Presh

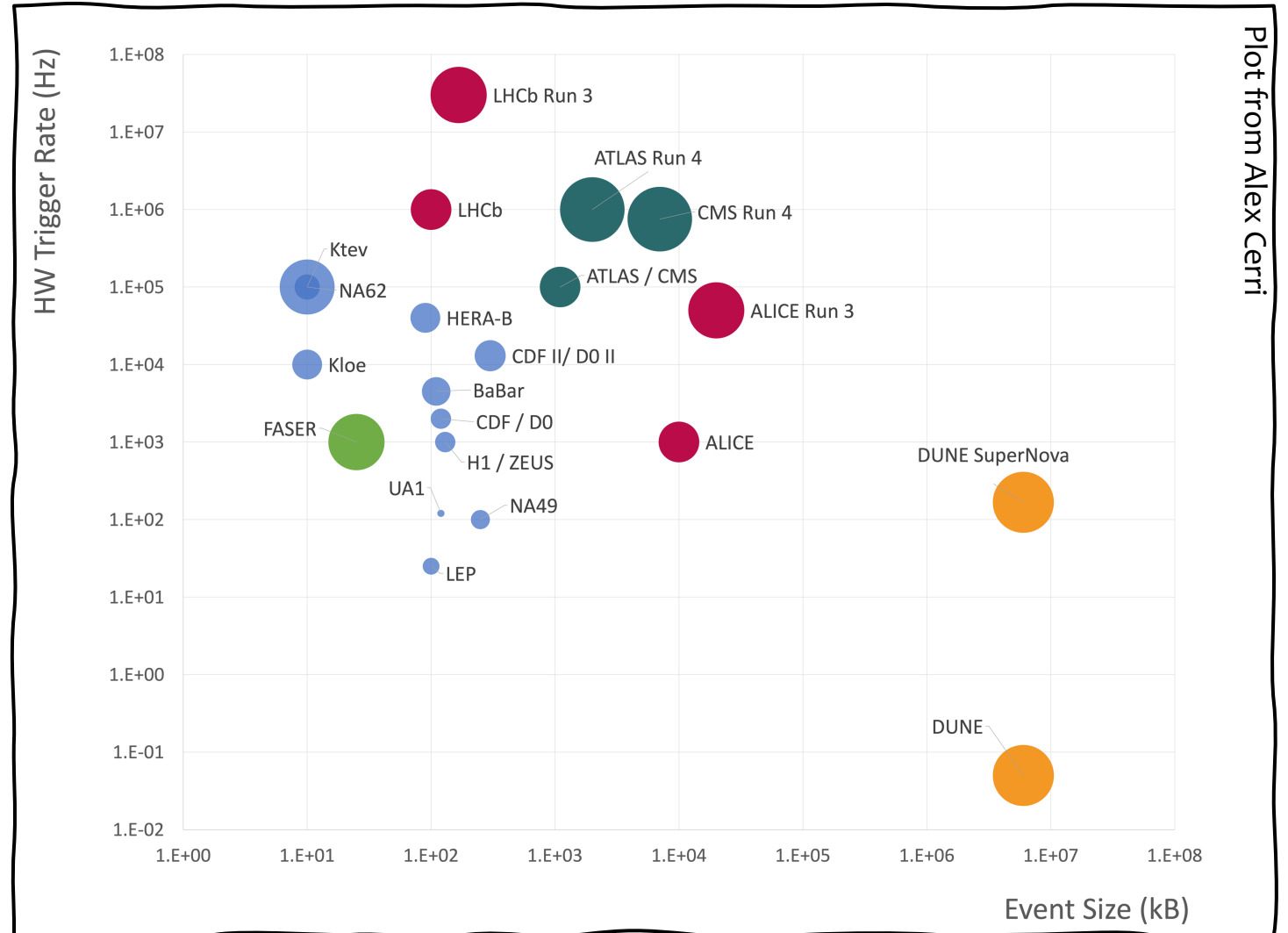
# TDAQ OPS

The rate of a coincidence trigger (requiring a signal in the veto scintillators at the front and the preshower scintillator layer at the back of FASER, likely signifying the passage of an energetic muon from the direction of IP1) is shown in red. The trigger rate trend generally follows the luminosity trend but it is evident that the trigger rate falls off more strongly at the beginning of fills than the luminosity. This is due to higher beam-induced backgrounds at the beginning of the fill. The dip in rate towards the end of the fill is due to an emittance scan at IP1.



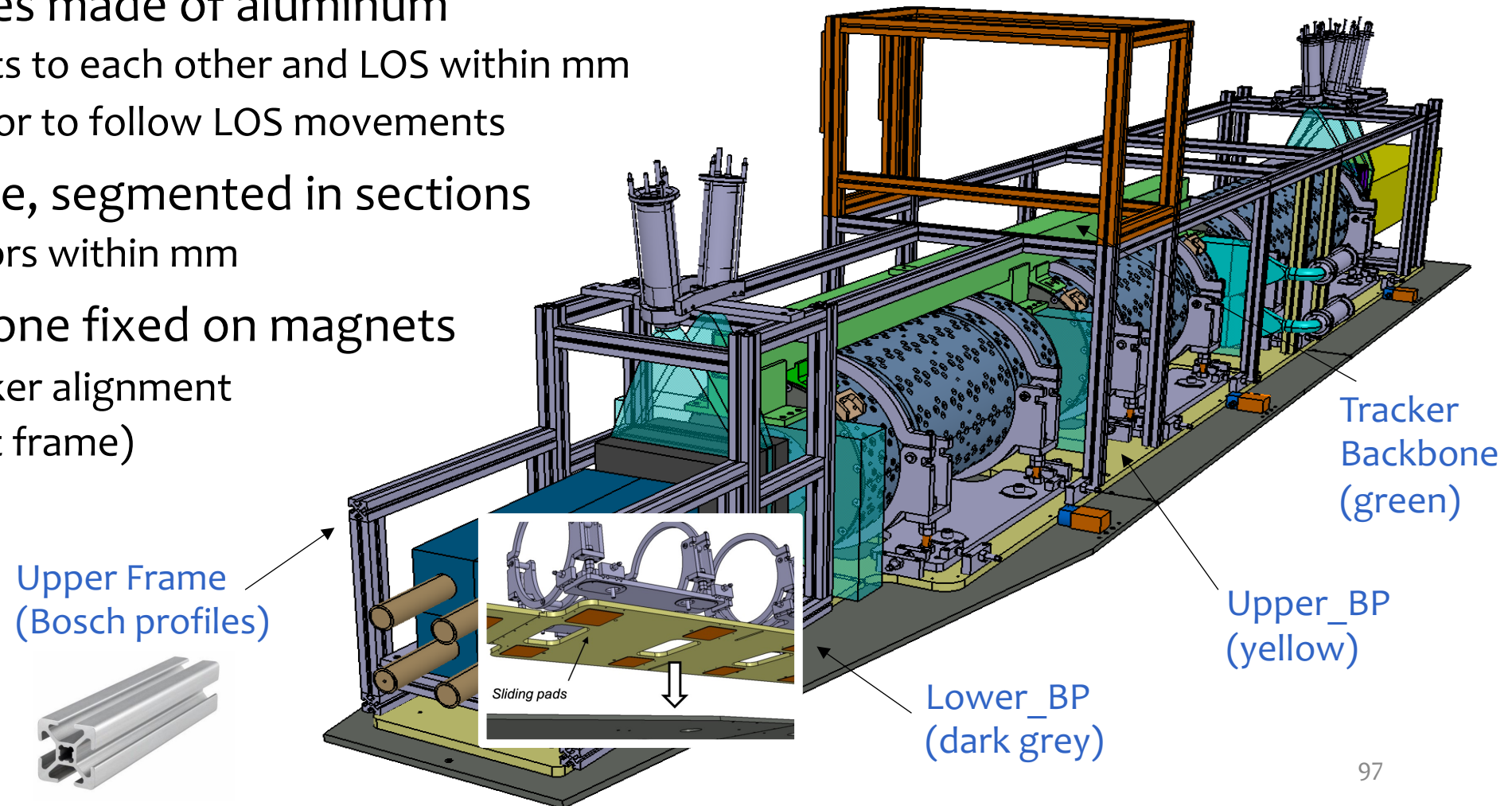
# ...WHY THIS ARCHITECTURE?

- Compare trigger rate and event size to other LHC experiments!



# DETECTOR SUPPORT STRUCTURE

- Two base-plates made of aluminum
  - Align magnets to each other and LOS within mm
  - Allow detector to follow LOS movements
- An upper frame, segmented in sections
  - Align detectors within mm
- Tracker backbone fixed on magnets
  - ensures tracker alignment ( $<100 \mu\text{m}$  wrt frame)

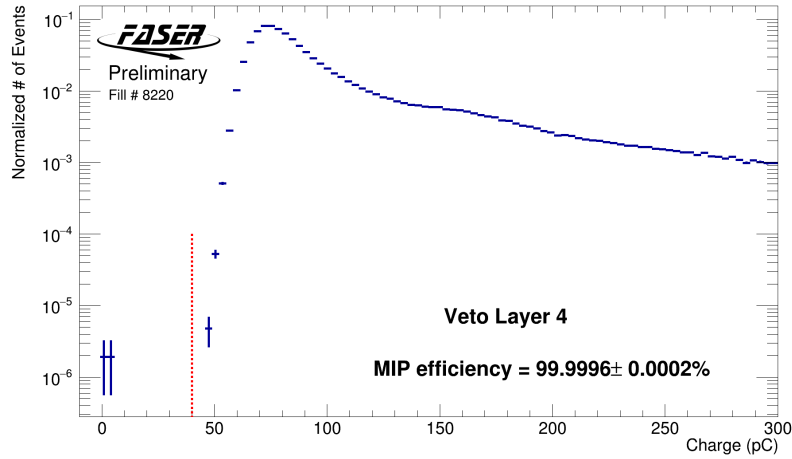




# PERFORMANCE ASSESSMENT

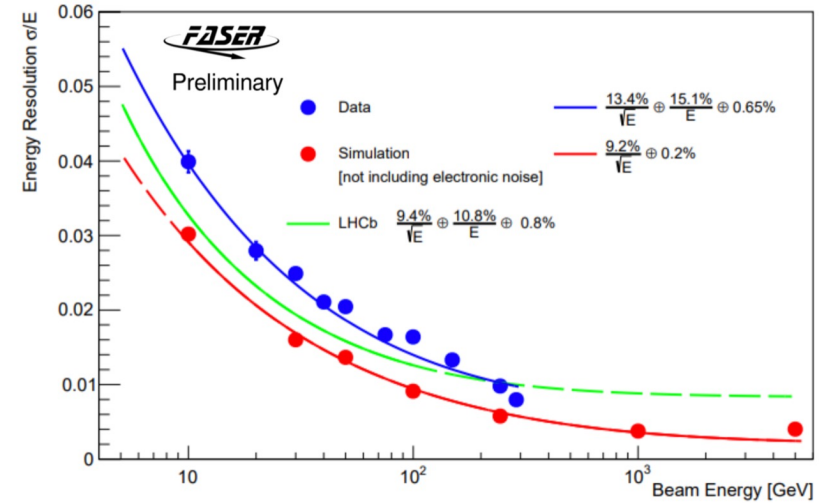
## Veto layer charge distribution Efficiency > 99.99% per layer

VETO SCINTILLATORS



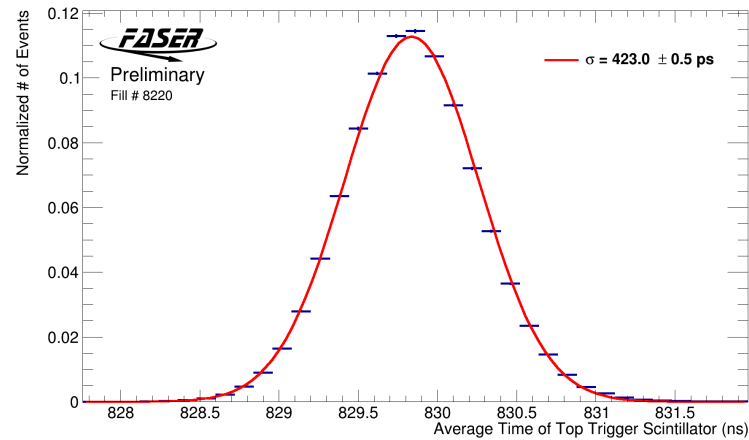
## Calorimeter calibration at testbeam Resolution confirmed ~1% for high energy electrons

CALORIMETER



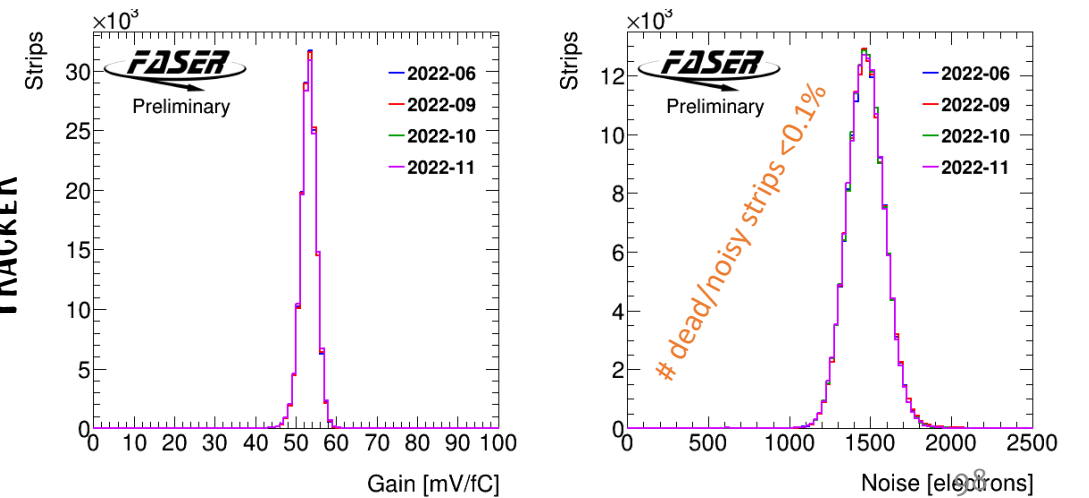
## Timing distribution of top trigger scintillator Resolution < 0.5 ns

TIMING SCINTILLATORS



## Gain and noise distributions of four tracking stations Confirm stable data taking conditions

TRACKER



# STUDYING BACKGROUNDS

- Before data taking, expected dominant background source:
  - high energy muons from IP
- With first data we see two extra background sources

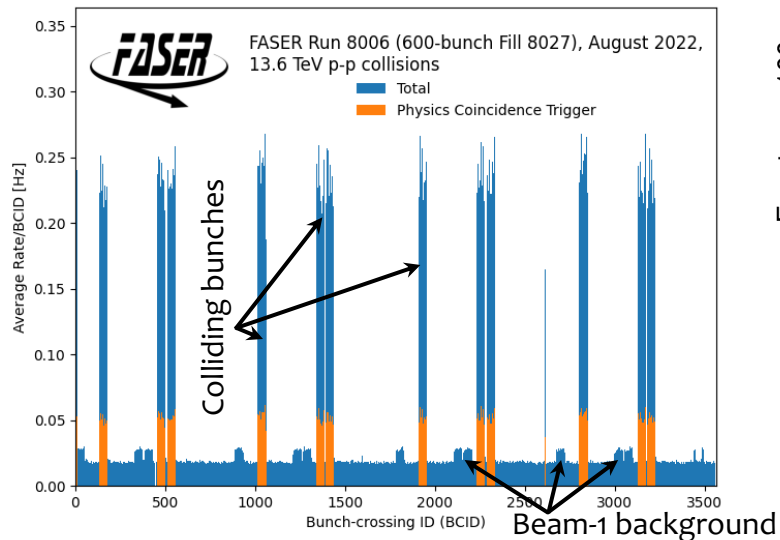
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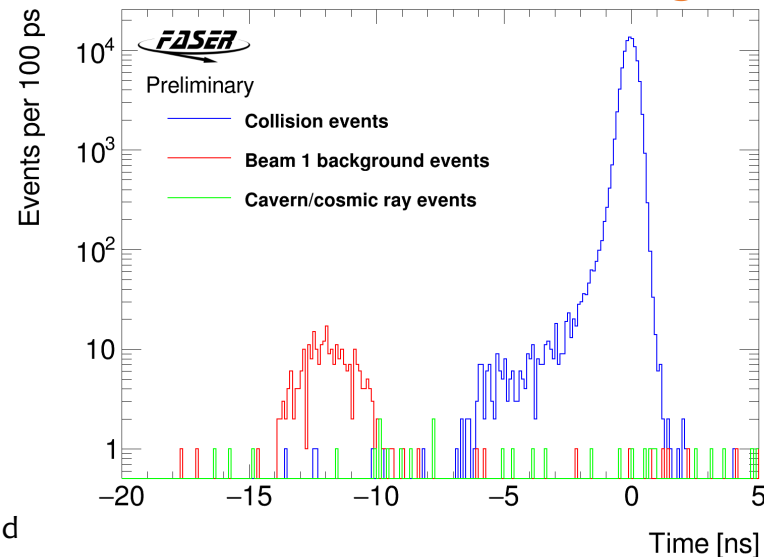
## SOURCE A

- Caused by beam-1 on the way to ATLAS passing from the back of FASER
- Early by about  $3.2 \mu\text{s}$  compared to particles from IP
- First observed with pilot beams in Nov 2021
- Concrete shielding installed to reduce it

Average rate per bunch-crossing ID



Calorimeter time separation of collision events and Beam-1 background



# STUDYING BACKGROUNDS

- Before data taking, expected dominant background source:
  - high energy muons from IP
- With first data we see two extra background sources

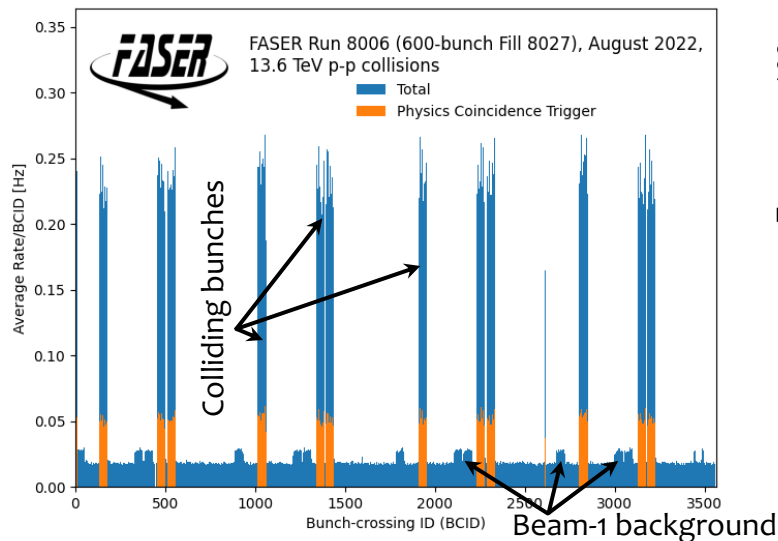
## SOURCE B

- Correlated to beam
- Only fires single scintillators
- Likely low energy neutrons

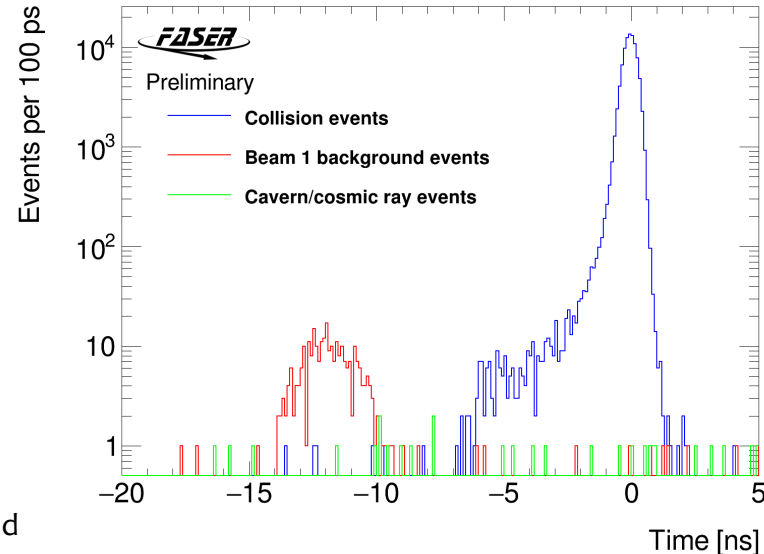
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Average rate per bunch-crossing ID



Calorimeter time separation of collision events and Beam-1 background





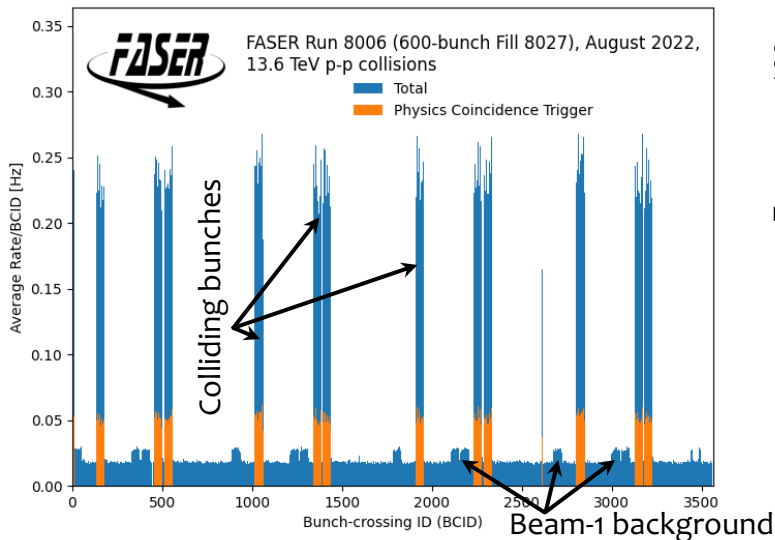
# STUDYING BACKGROUNDS

- Before data taking, expected dominant background source:
  - high energy muons from IP
- With first data we see two extra background sources

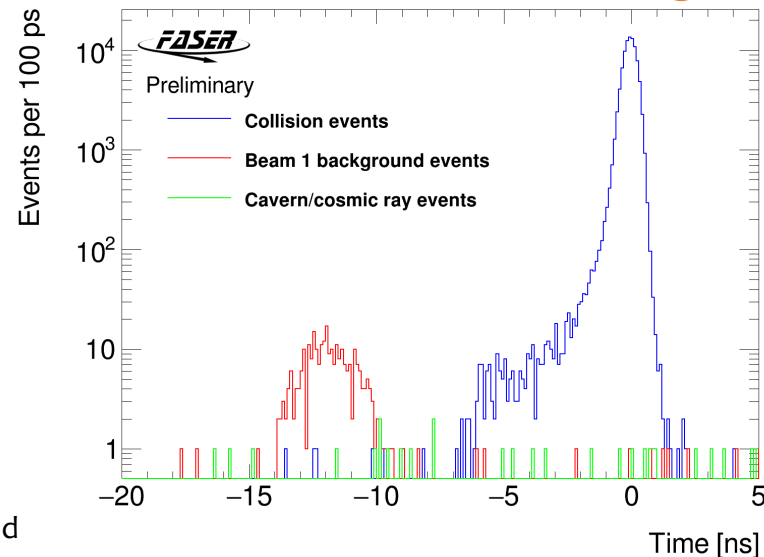
## SOURCE A

- Caused by beam-1 on the way to ATLAS passing from the back of FASER
- Early by about  $3.2 \mu\text{s}$  compared to particles from IP
- First observed with pilot beams in Nov 2021
- Concrete shielding installed to reduce it

Average rate per bunch-crossing ID



Calorimeter time separation of collision events and Beam-1 background

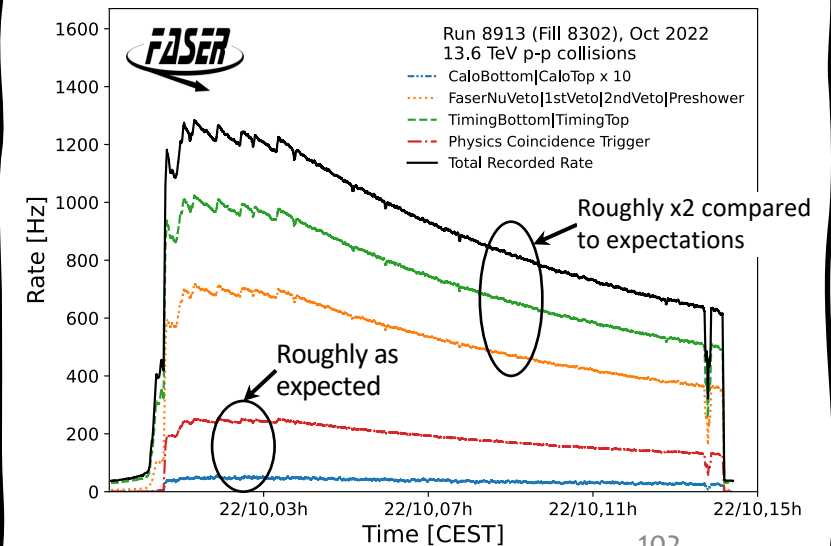


## SOURCE B

- Correlated to beam
- Only fires single scintillators
- Likely low energy neutrons

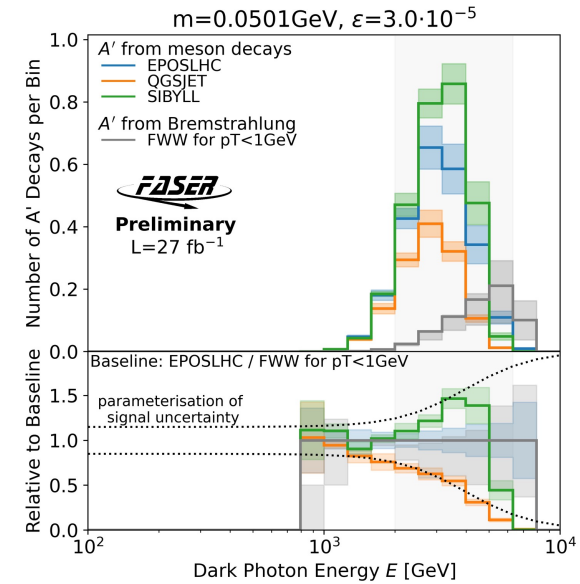
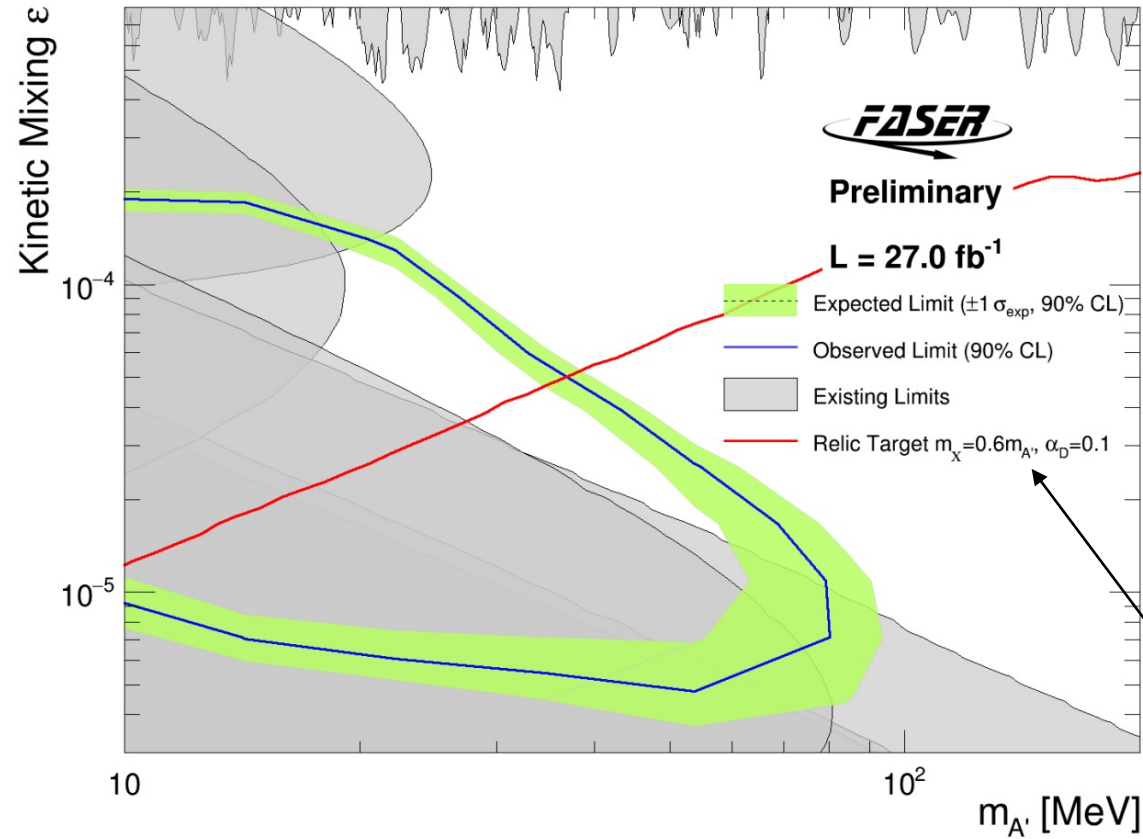
- Additional backgrounds give total trigger rate x2 than expected
- Rate of muons from IP roughly consistent with expectation
- Extra rate not problematic

FASER recorded rate per trigger item

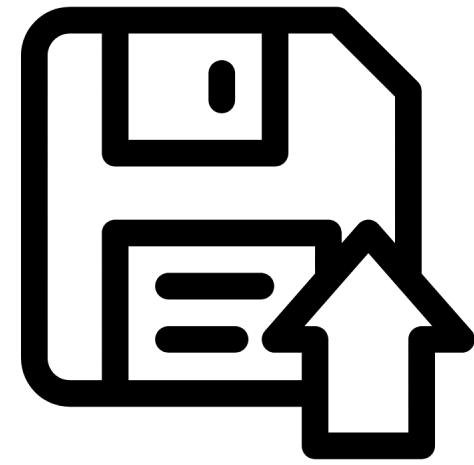


# EXCLUSION REACH

On previously unexplored phase-space



- Parameters for which DM  $\chi$  annihilates via  $\chi\chi \rightarrow A' \rightarrow ff$
- Model dependent *line*, but region favoured by DM relic density



# THE *FASER* DETECTOR

**Preshower upgrade**

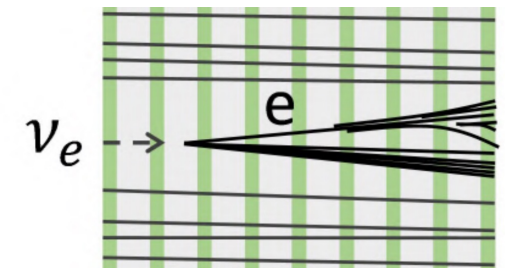
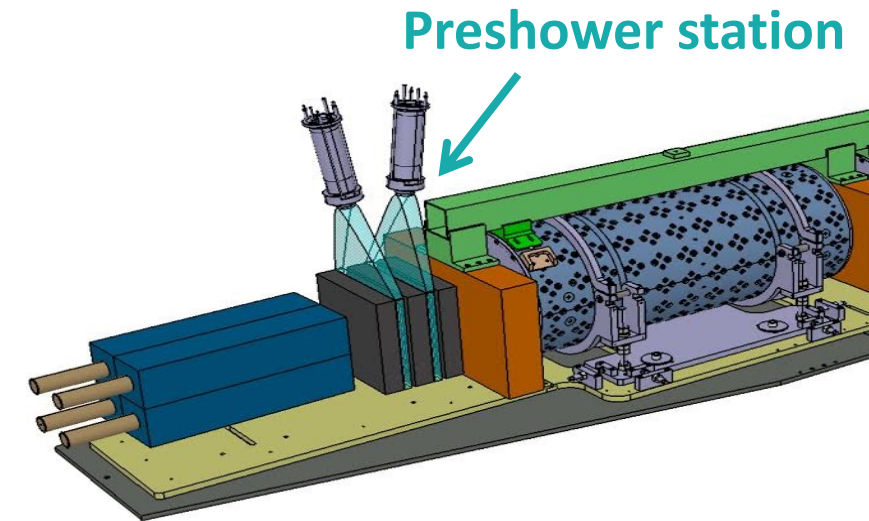
# CURRENT DETECTOR LIMITATION

- **Current FASER pre-shower**

- two layers of scintillators, each preceded by a 1X0 lead-radiator plane
- will create a photon shower to help distinguish photons from electrons coming from deep inelastic scattering (DIS) of very energetic neutrinos in the calorimeter.

- **Limitations:**

- no information about the multiplicity of the photons or the topology of the event
- neutrino DIS events produced in the 2 X0 lead of the present pre-shower will be undistinguishable from a photon (LLP) signature
  - About 10 such events expected in 150/fb

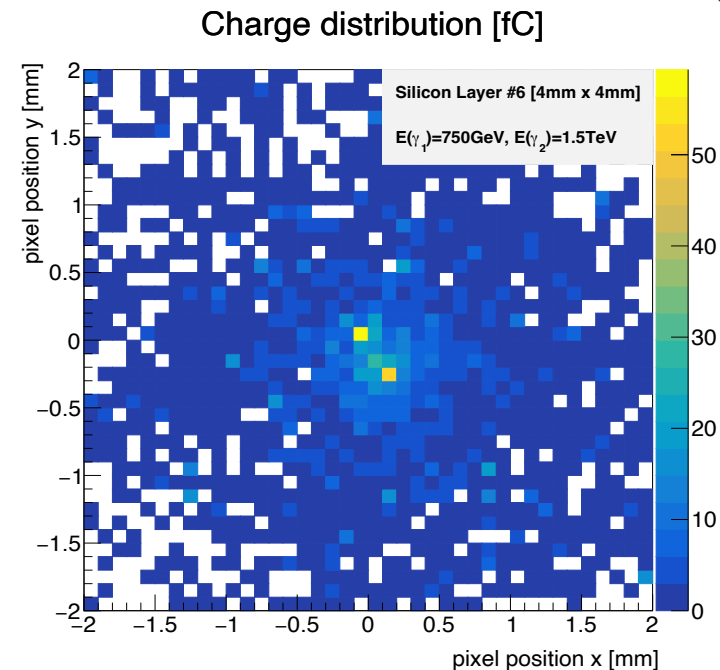
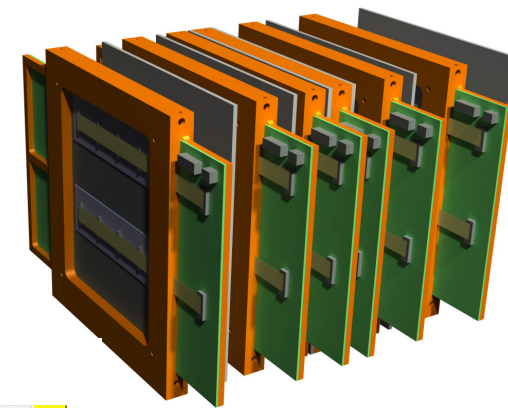
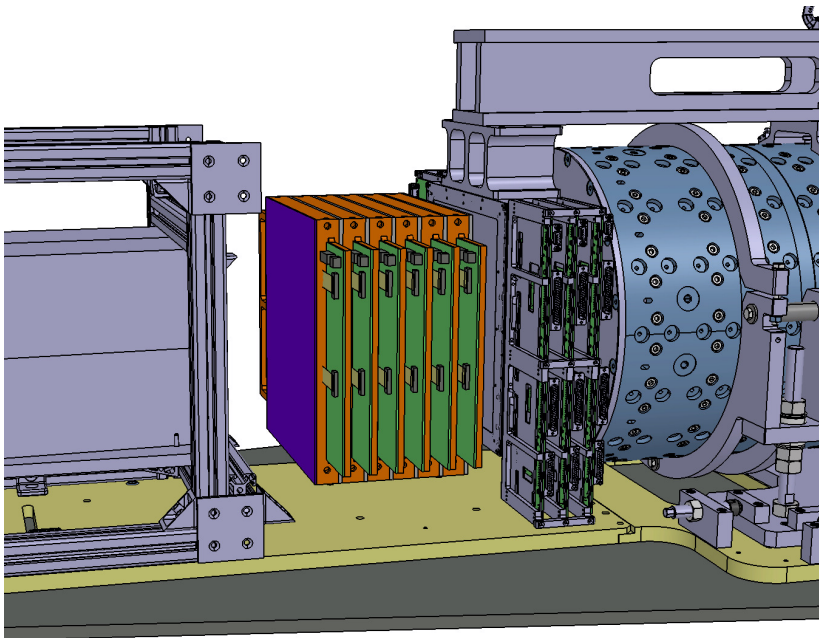


Cartoon of  $\nu_e$  DIS  
in FASERv



# UPGRADE TO ENABLE $2-\gamma$ PHYSICS

- Existing pre-shower to be replaced with a high-resolution silicon pre-shower detector using monolithic pixel ASICs
  - hexagonal pixels of  $65\ \mu\text{m}$  side

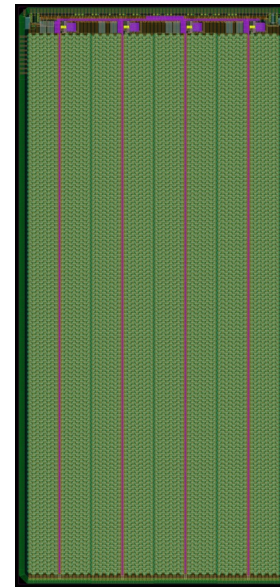
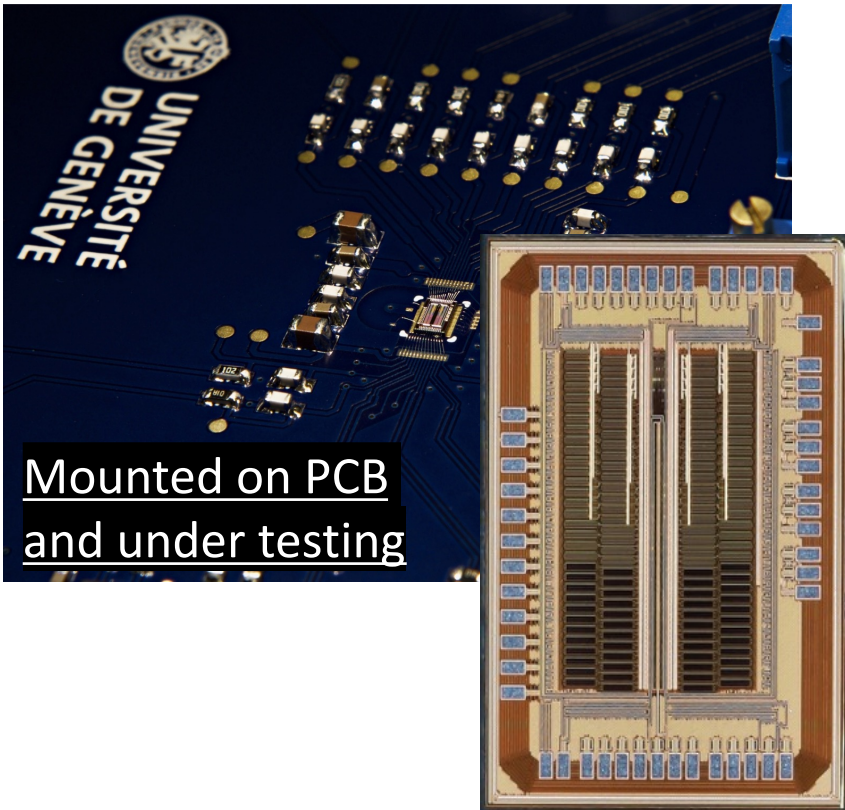


- Distance between two photons:  $200\ \mu\text{m}$
- Distinguishable!

Detector to be used for  
2024 & 2025 data taking  
(70% of Run3 data)

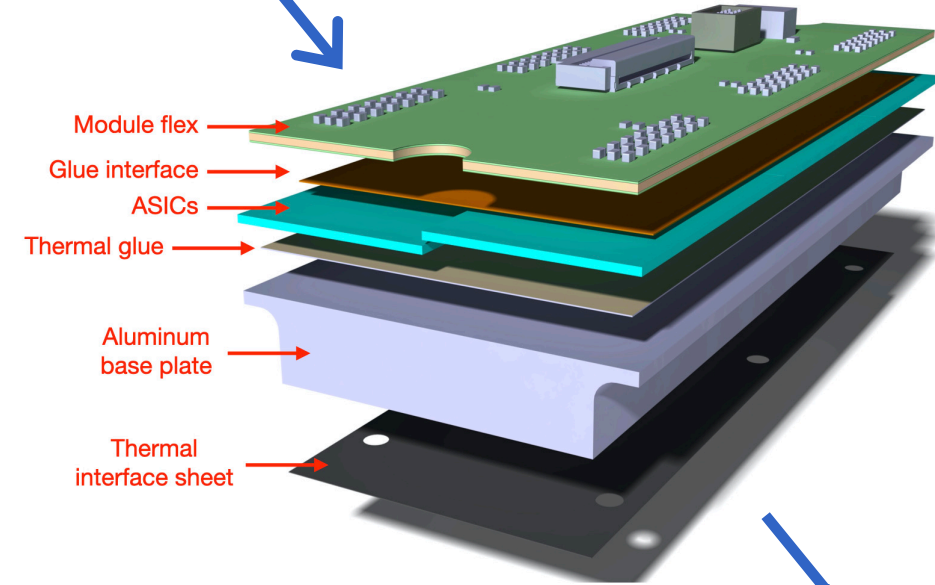
# FROM PIXELS TO LAYERS OF MODULES

Prototype ASIC available end 2020

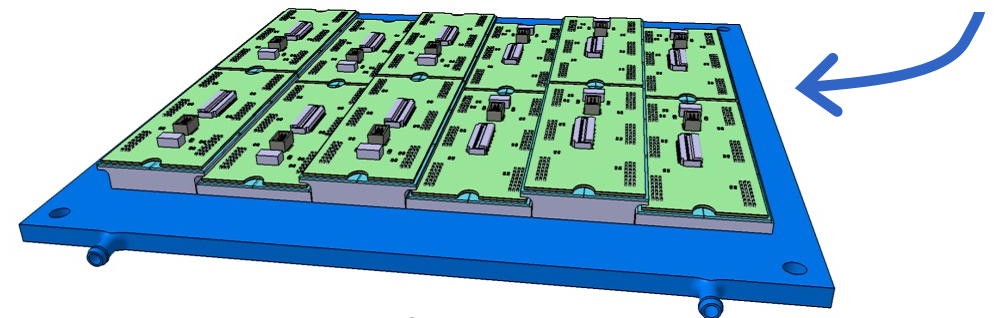


Pixel chips into sensor wafers

...into modules of pixels



...into layers of modules



Proposed layout of 1 layer with 12 modules.



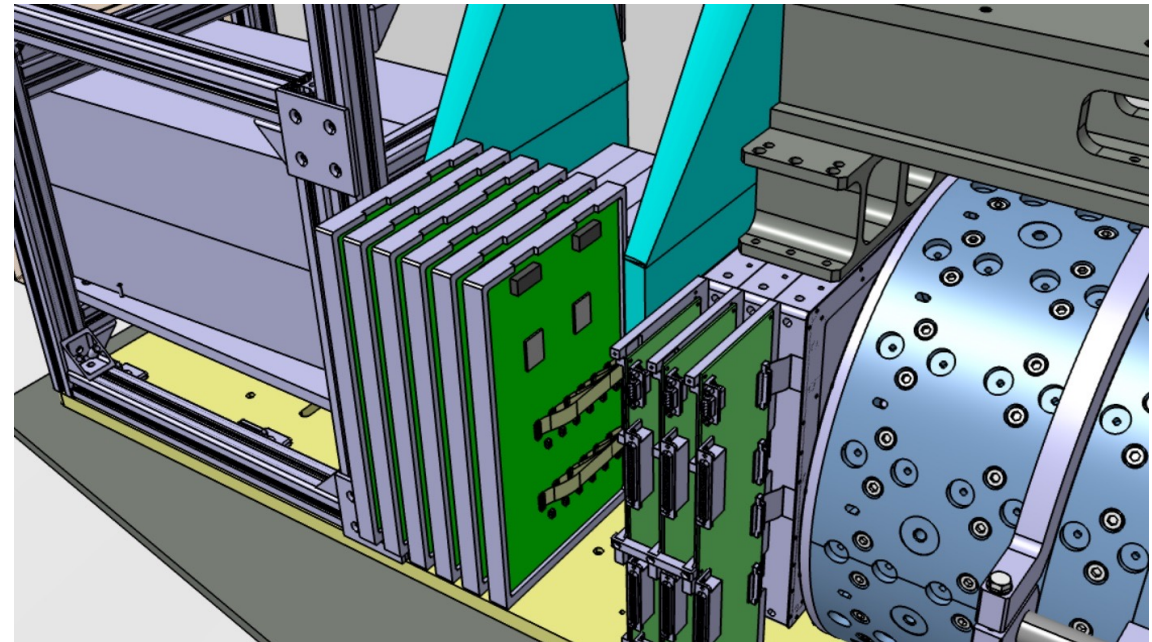
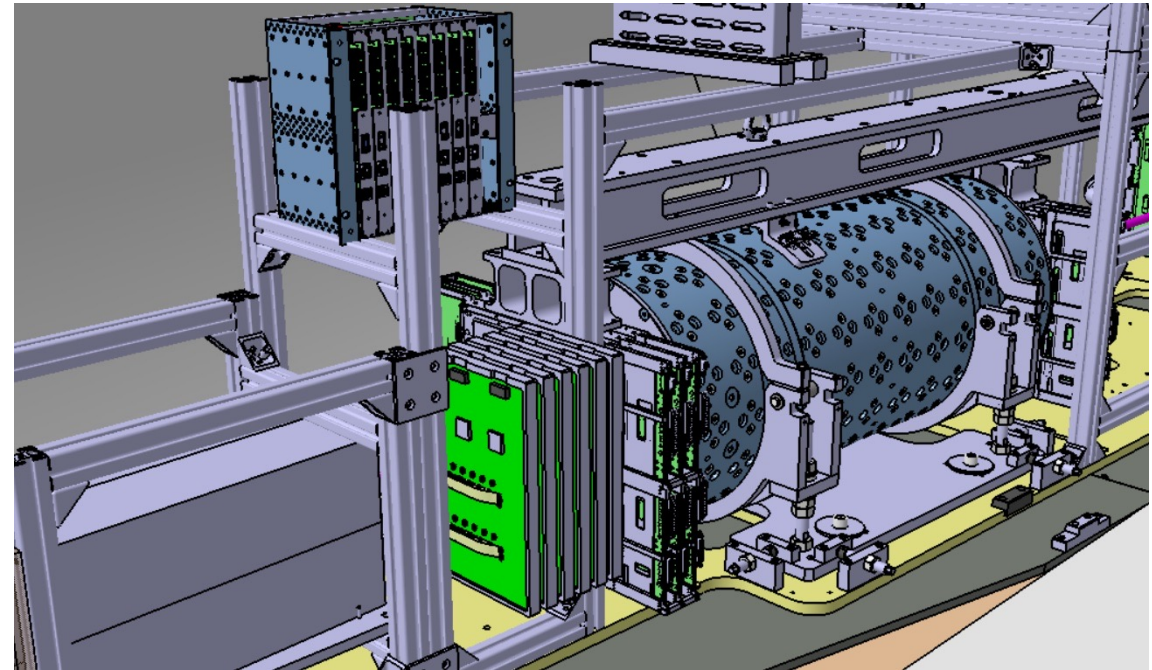
# UPGRADED PRE-SHOWER

## Status:

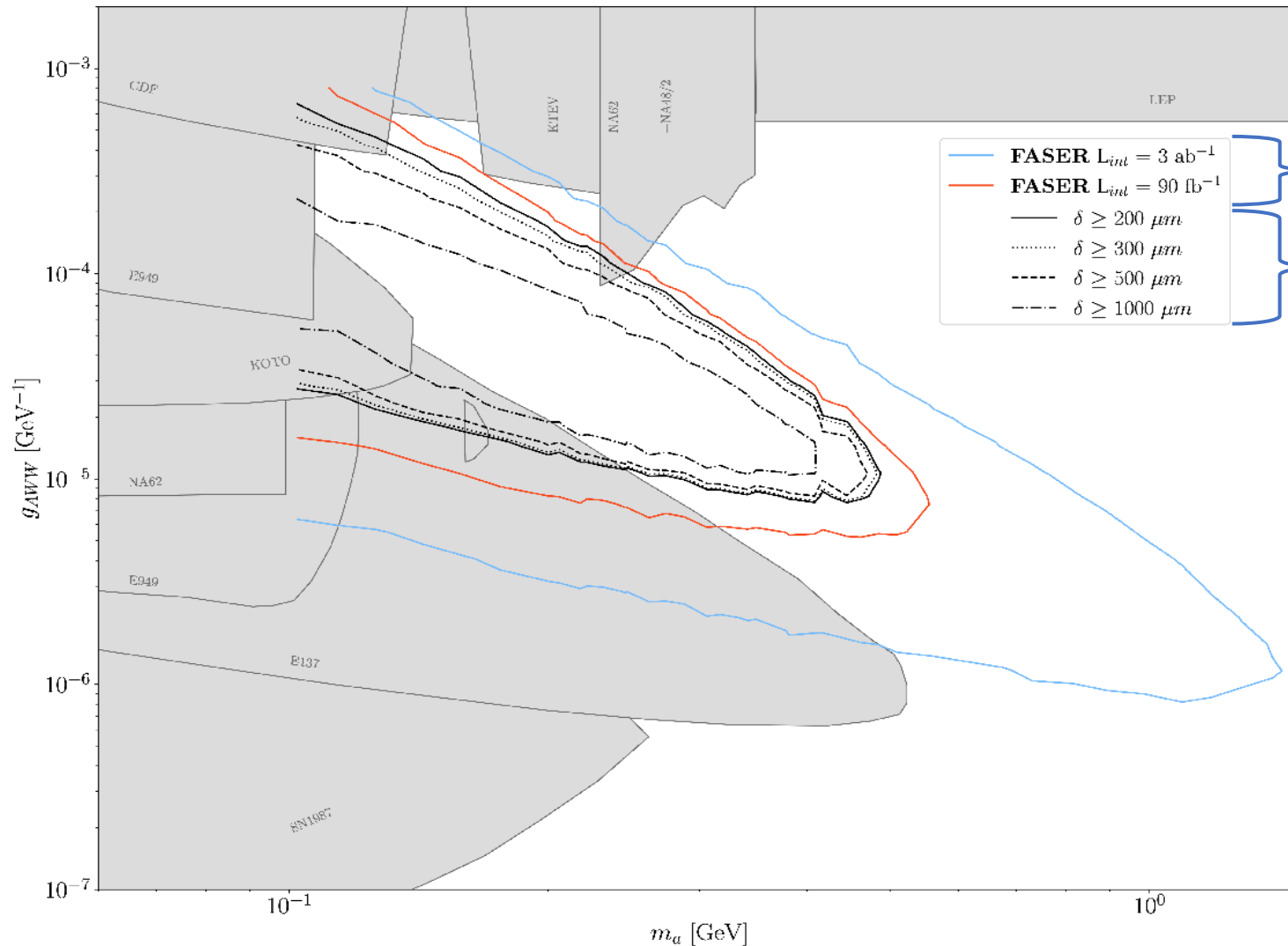
- Preparing the detailed technical proposal
- Pre-production ASICs in foundry, expected ~ March
- Design of modules, planes, mechanics, read-out in progress
- Simulation and reconstruction in progress

## Plan:

- Install the detector end of 2023 for data taking in 2024 and the rest of Run3



# REACH FOR ALPS



2-photon pairs with  
 $E > 250 \text{ GeV}$  and  $\delta_{\gamma\gamma} > 0.2 \text{ mm}$   
 Zero background events assumed

Ideal detector (100% efficiency)

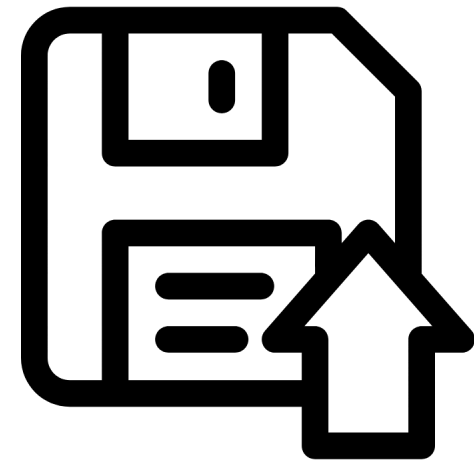
Realistic detector

(65-75% efficiency for a  $\delta_{\gamma\gamma} = 0.2 \text{ mm}$ ,  
 85-90% efficiency for  $\delta_{\gamma\gamma} \geq 0.3 \text{ mm}$ )

→ 2-photon pairs can be  
 resolved and studied!

**Otherwise: Indistinguishable  
 from background**

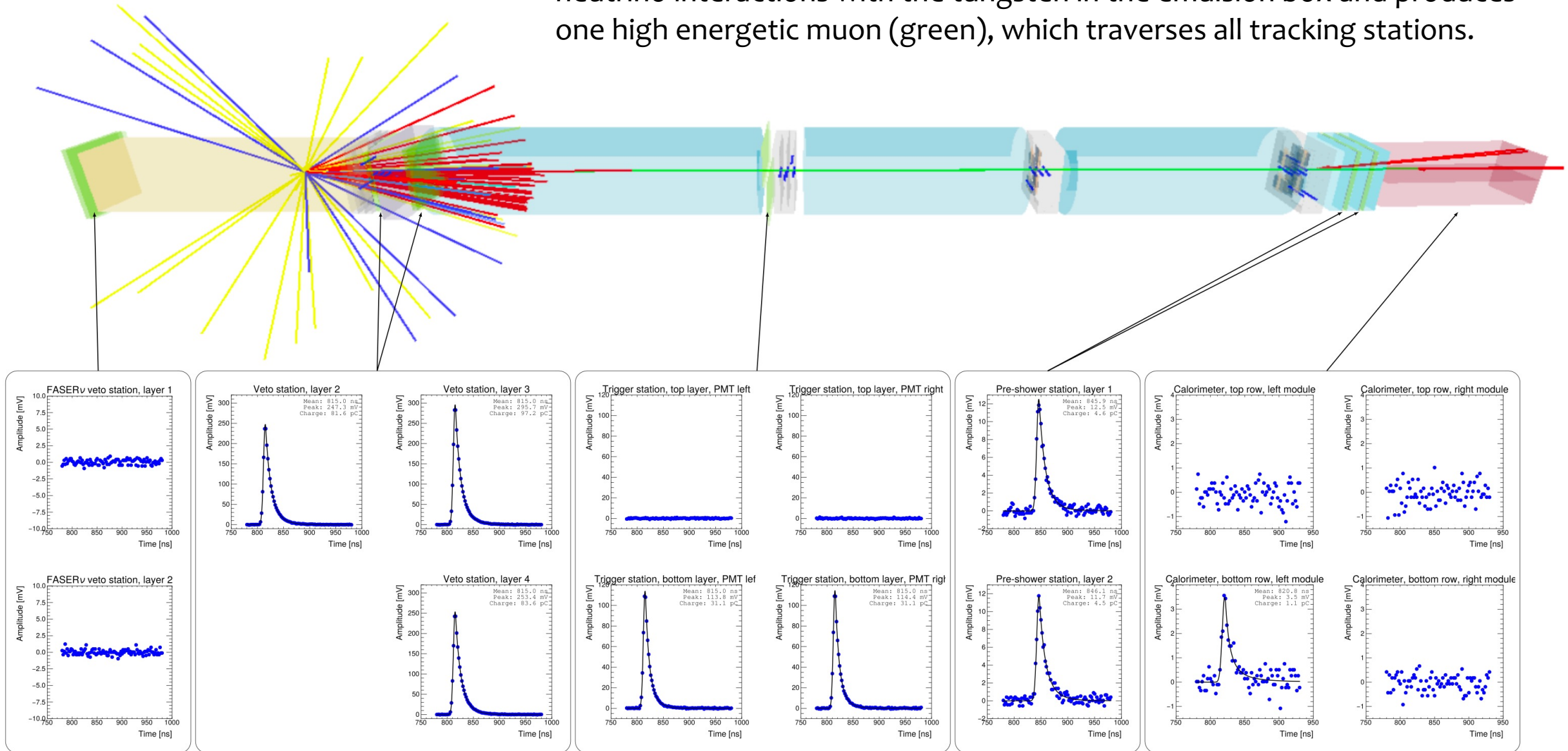




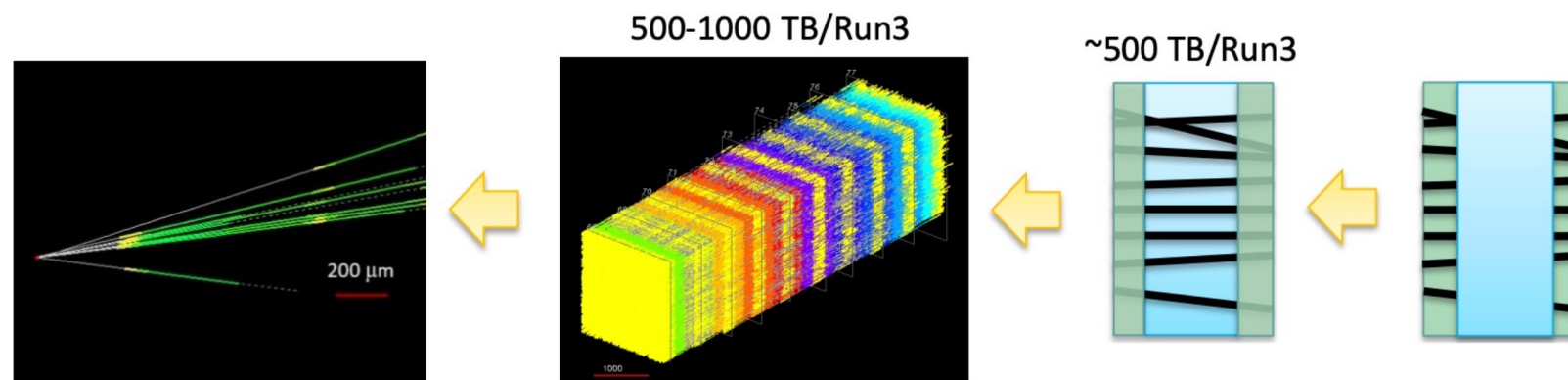
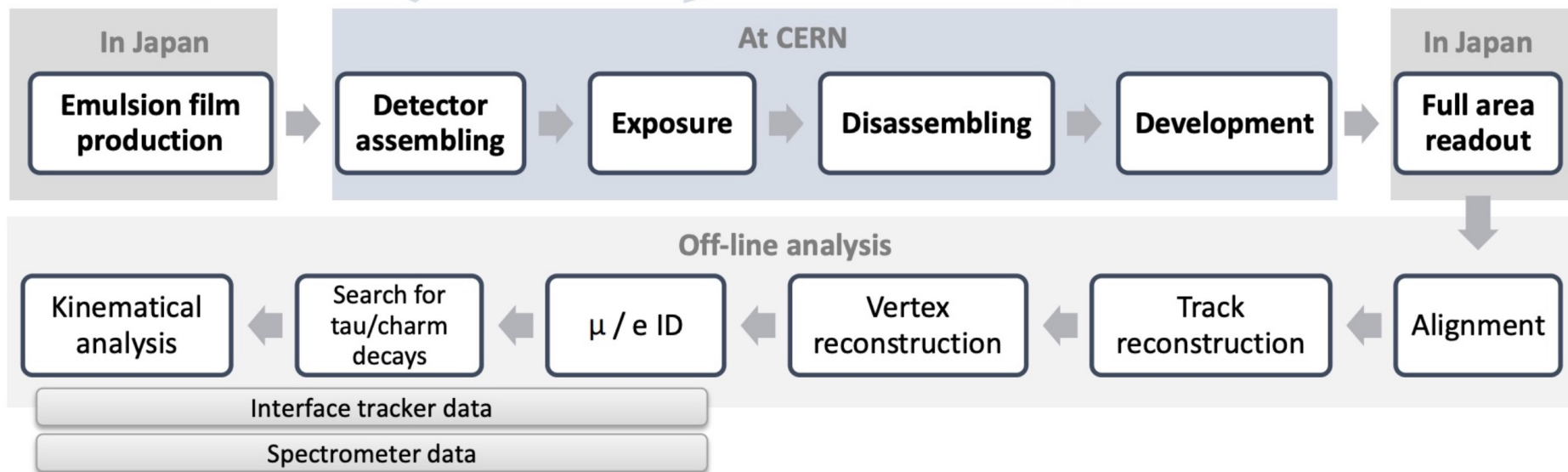
***FAZER***

**FAZERmu**

Event display for a typical charged current interaction where a muon neutrino interacts with the tungsten in the emulsion box and produces one high energetic muon (green), which traverses all tracking stations.



# FASERv Workflow





# EMULSION

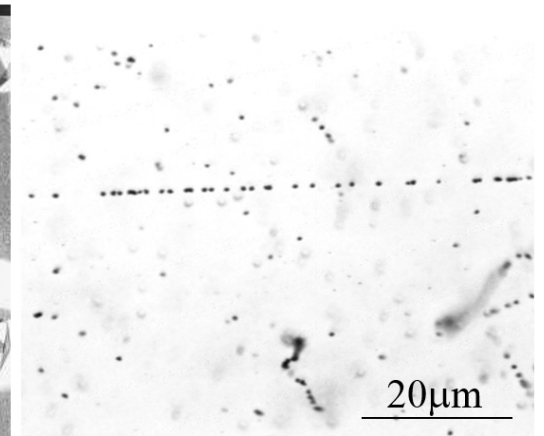
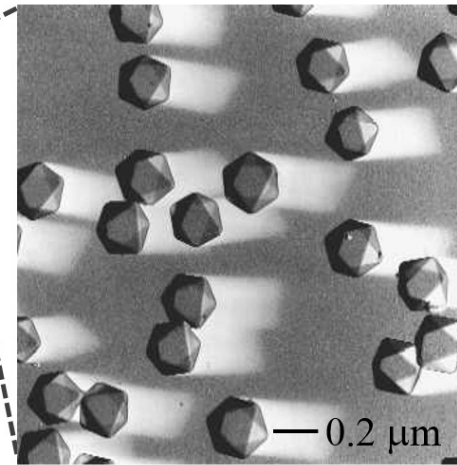
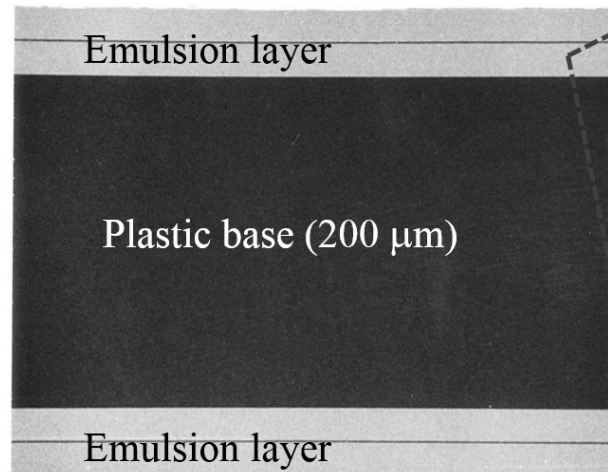


Photo of an emulsion film (left), its cross-sectional view (left center), electron microscope image of the silver halide crystals (right center), and a minimum ionising particle track from a 10 GeV/c  $\pi$  beam (right).



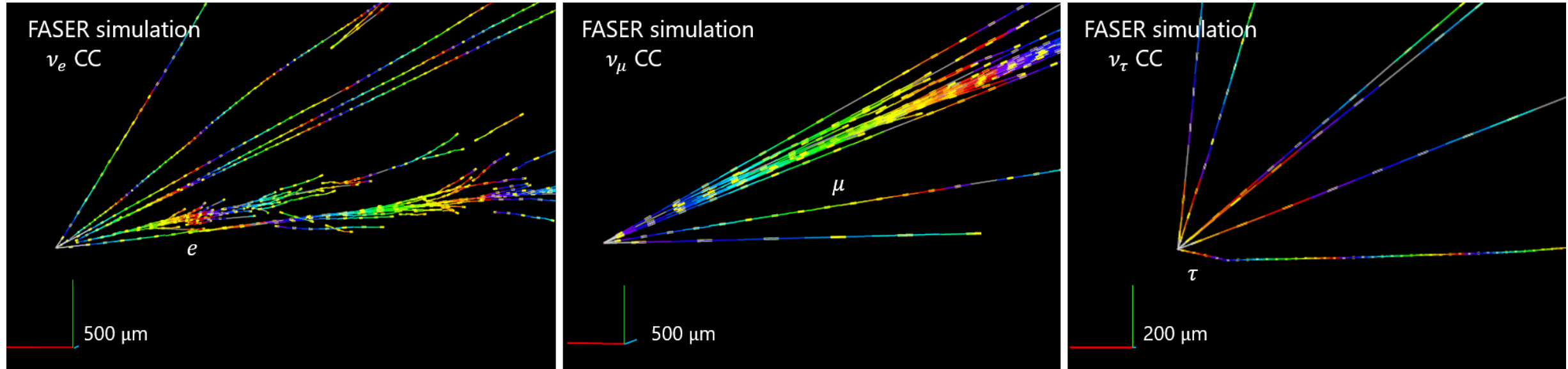


# NEUTRINO PRODUCTION

Type	Particles	Main Decays	E	Q	S	P
Pions	$\pi^+$	$\pi^+ \rightarrow \mu\nu$	✓	✓	✓	—
Kaons	$K^+, K_S, K_L$	$K^+ \rightarrow \mu\nu, K \rightarrow \pi l\nu$	✓	✓	✓	—
Hyperons	$\Lambda, \Sigma^+, \Sigma^-, \Xi^0, \Xi^-, \Omega^-$	$\Lambda \rightarrow p l\nu$	✓	✓	✓	—
Charm	$D^+, D^0, D_s, \Lambda_c, \Xi_c^0, \Xi_c^+$	$D \rightarrow K l\nu, D_s \rightarrow \tau\nu, \Lambda_c \rightarrow \Lambda l\nu$	—	—	✓	✓
Bottom	$B^+, B^0, B_s, \Lambda_b, \dots$	$B \rightarrow D l\nu, \Lambda_b \rightarrow \Lambda_c l\nu$	—	—	—	✓

TABLE I. Decays considered for the estimate of forward neutrino production. For each type in the first column, we list the considered particles in the second column and the main decay modes contributing to neutrino production in the third column. In the last four columns we show which generators were used to obtain the meson spectra: EPOS-LHC (E) [59], QGSJET-II-04 (Q) [60], SIBYLL 2.3C (S) [61–64], and PYTHIA 8 (P) [66, 67], using both the MONASH-tune [68] and the minimum bias A2-tune [69].

# EVENT DISPLAYS OF SIMULATED NEUTRINO INTERACTION VERTICES

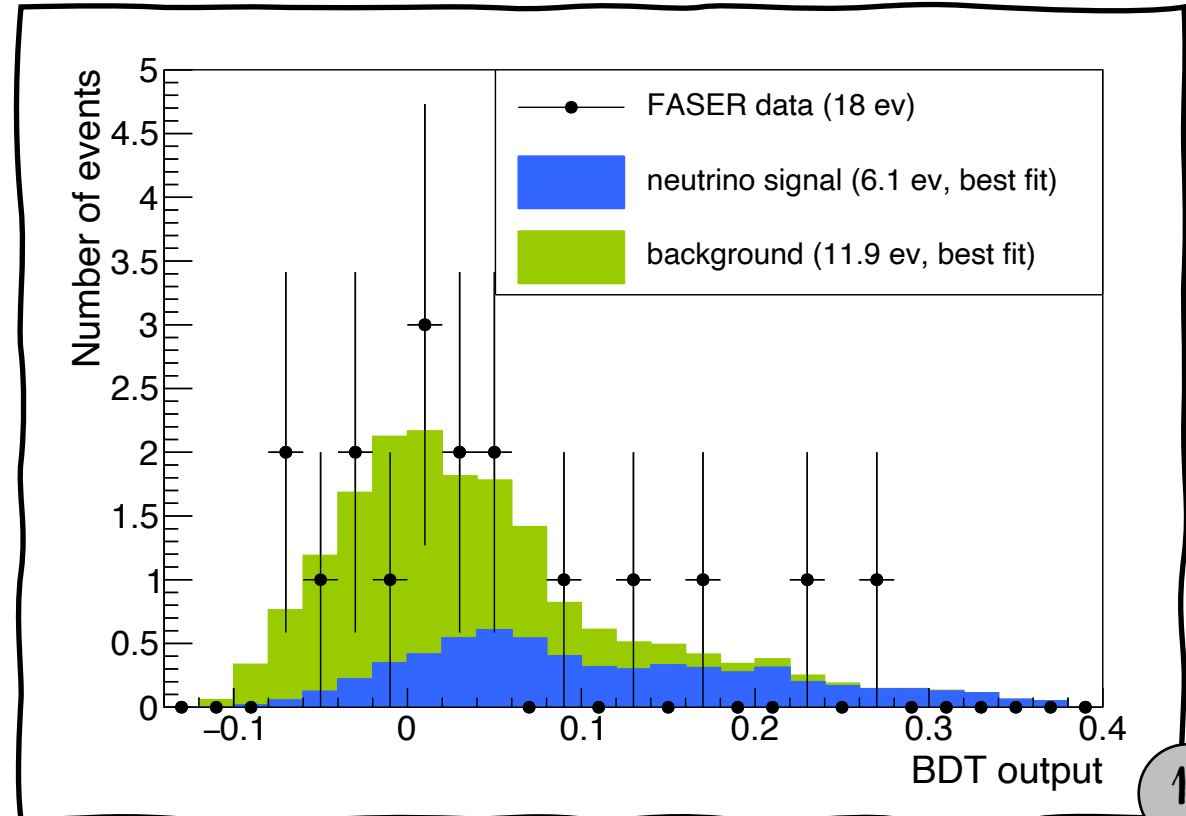
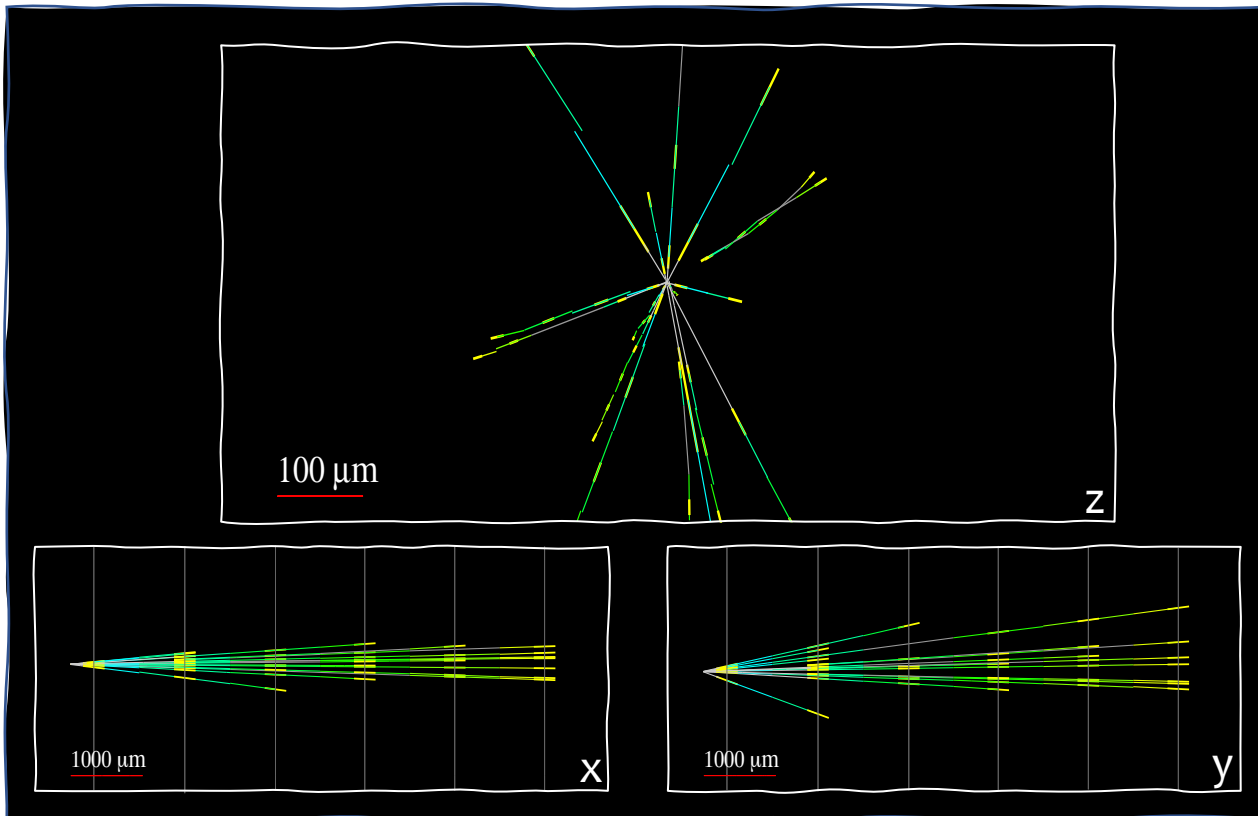
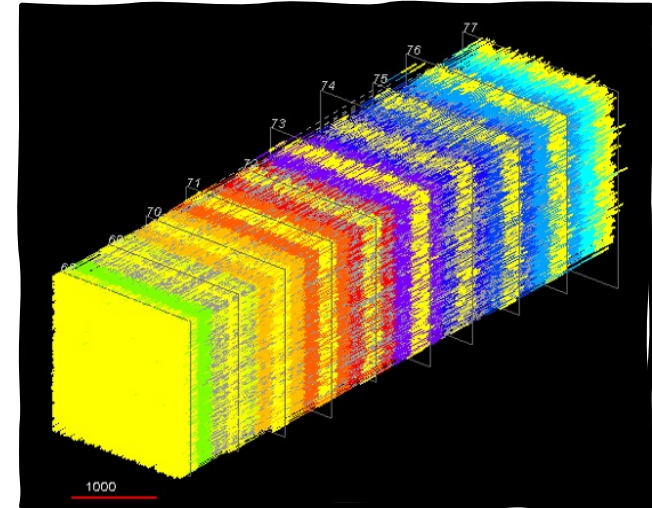


Event displays of simulated neutrino interaction vertices for 433 GeV  $\nu_e$  CC, 664 GeV  $\nu_\mu$  CC, and 831 GeV  $\nu_\tau$  CC. Yellow line segments show the trajectories of charged particles in the emulsion films. The other colored lines are extrapolations of the track hits to the neighboring tungsten plates, and the colors change depending on the depth in the detector.



# PILOT RUN IN 2018

First candidate **collider neutrino** interactions!

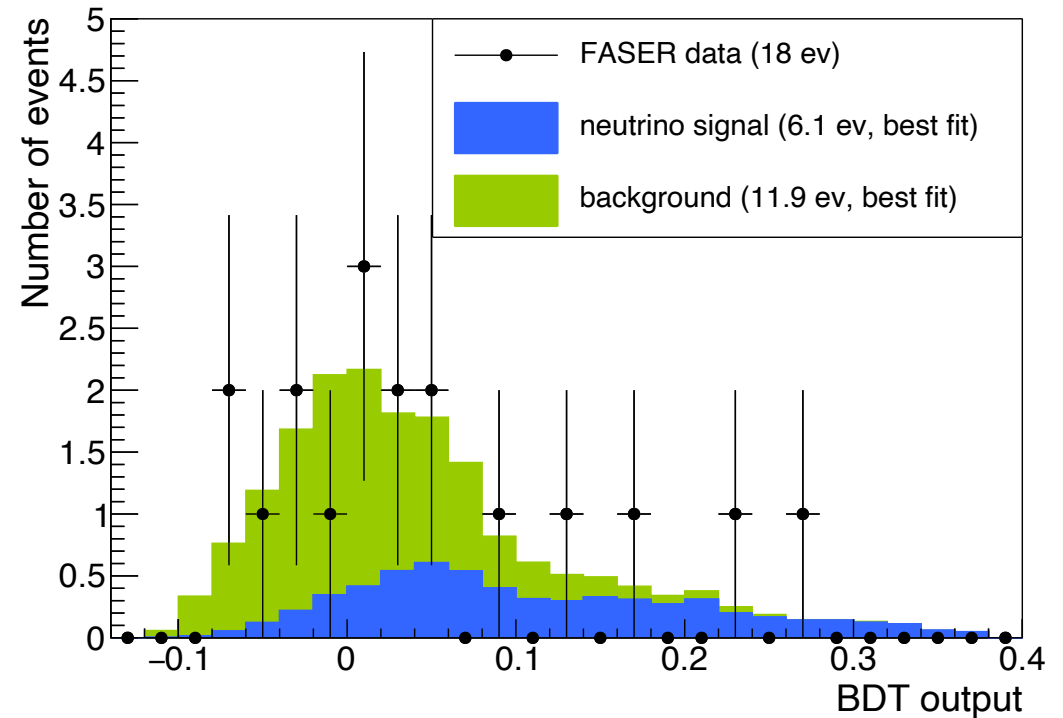
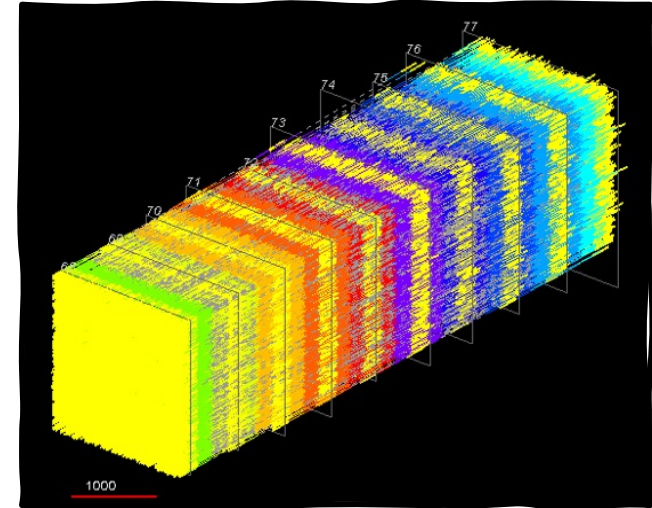




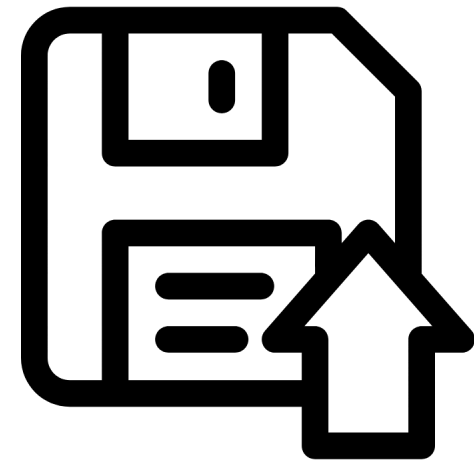
# PILOT RUN IN 2018

First candidate **collider neutrino** interactions!

- A 11 kg detector collected  $\sim 12/\text{fb}$
- About 3.3 neutrino interactions expected to have occurred after selections
- BDT developed to distinguish neutrino signal from neutral hadron background
  - The background-only hypothesis is rejected with significance of  $2.7\sigma$
- Excellent testbed for future data analysis







***FASE***

**RESULTS**

**Neutrinos**

# STATISTICAL ANALYSIS FOR SIGNIFICANCE ESTIMATION

Besides the signal category, we select:

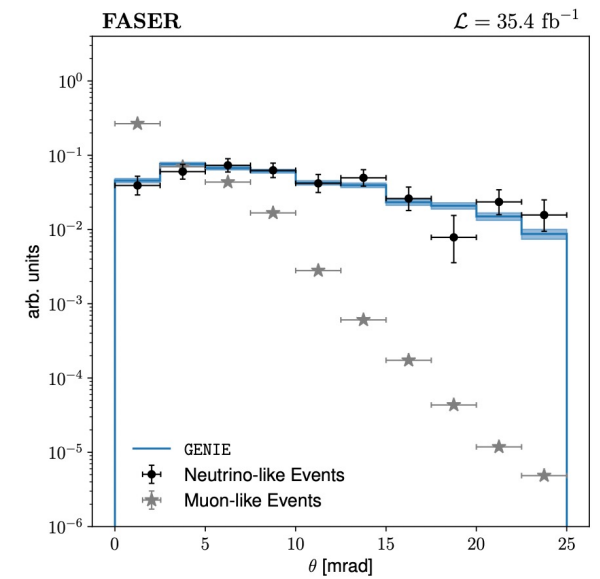
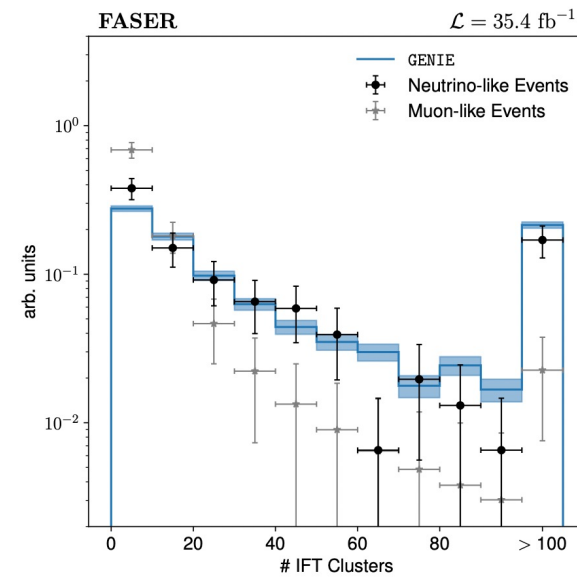
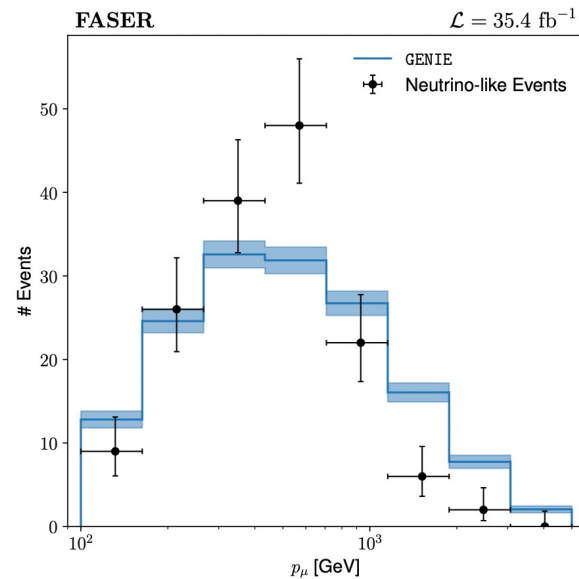
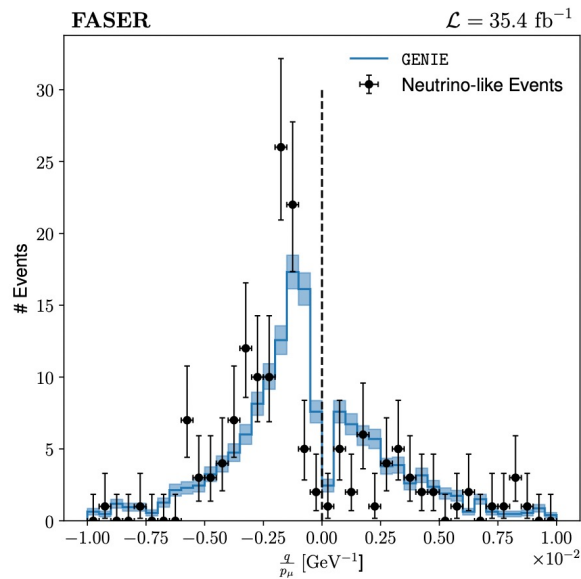
- $n_{10}$ : Events for which the first layer of the FASER $\nu$  scintillator produces a charge of  $>40$  pC in the PMT, but no signal with sufficient charge is seen in the second layer.
- $n_{01}$ : Analogous events for which more than 40 pC in the PMT was observed in the second layer, but not in the first layer.
- $n_2$ : Events for which both layers observe more than 40 pC of charge.

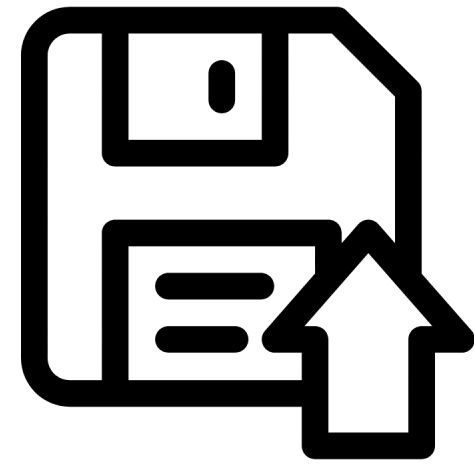
tor response and selection. The determined inefficiencies of the two FASER $\nu$  scintillators are  $p_1 = (6_{-3}^{+4}) \times 10^{-8}$  and  $p_2 = (9_{-3}^{+4}) \times 10^{-8}$ , showing values close to the expected performance [25].

Category	Events	Expectation
Signal	153	$n_\nu + n_b \cdot p_1 \cdot p_2 + n_{\text{had}} + n_{\text{geo}} \cdot f_{\text{geo}}$
$n_{10}$	4	$n_b \cdot (1 - p_1) \cdot p_2$
$n_{01}$	6	$n_b \cdot p_1 \cdot (1 - p_2)$
$n_2$	64014695	$n_b \cdot (1 - p_1) \cdot (1 - p_2)$

TABLE I. Observed event yields in  $35.4 \text{ fb}^{-1}$  of collision data and their relation to neutrino and background events.

# PROPERTIES OF OBSERVED EVENTS





***FASER***

**RESULTS**

**Dark Photon**



# EXPECTED DARK PHOTON PHASE-SPACE

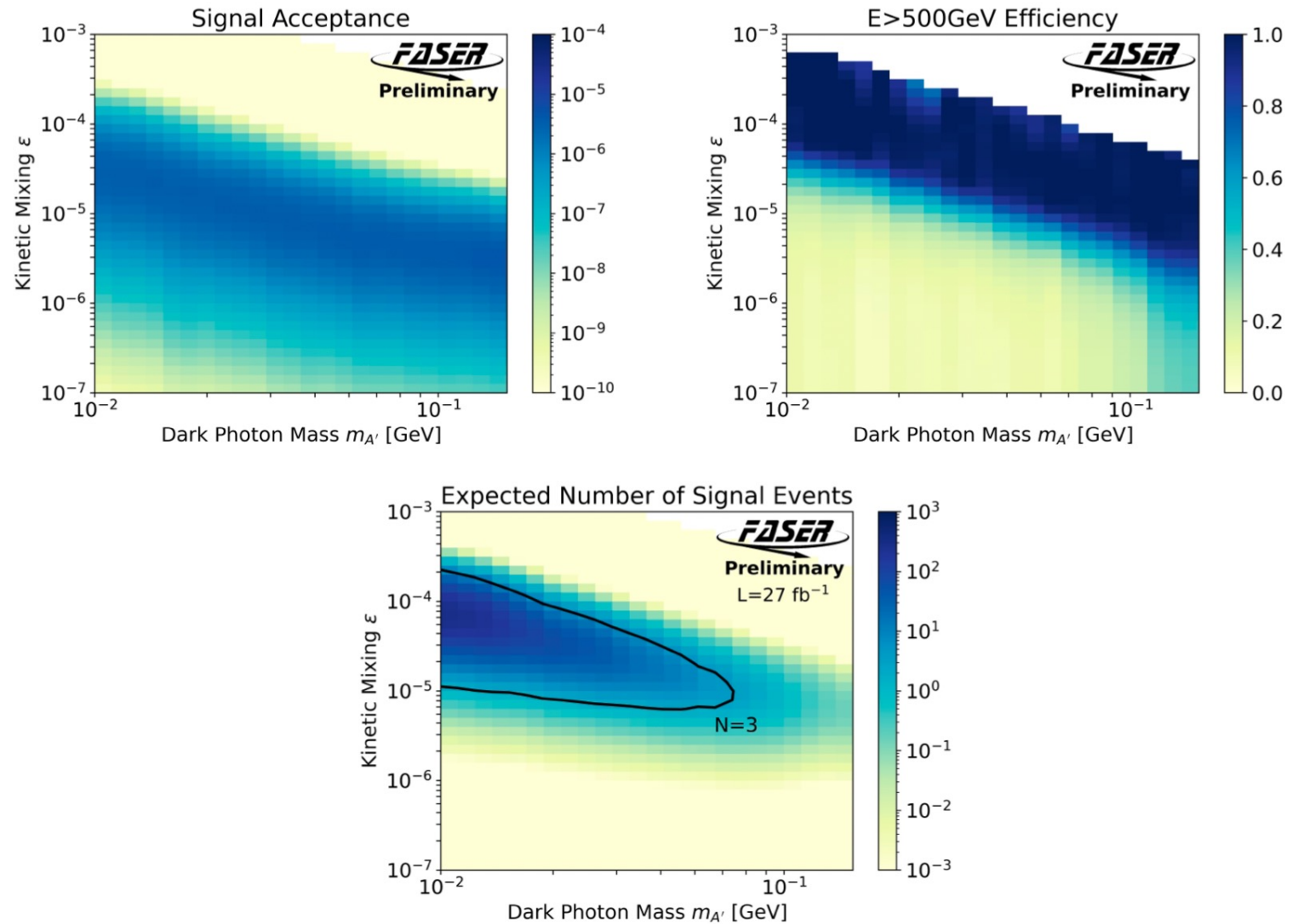
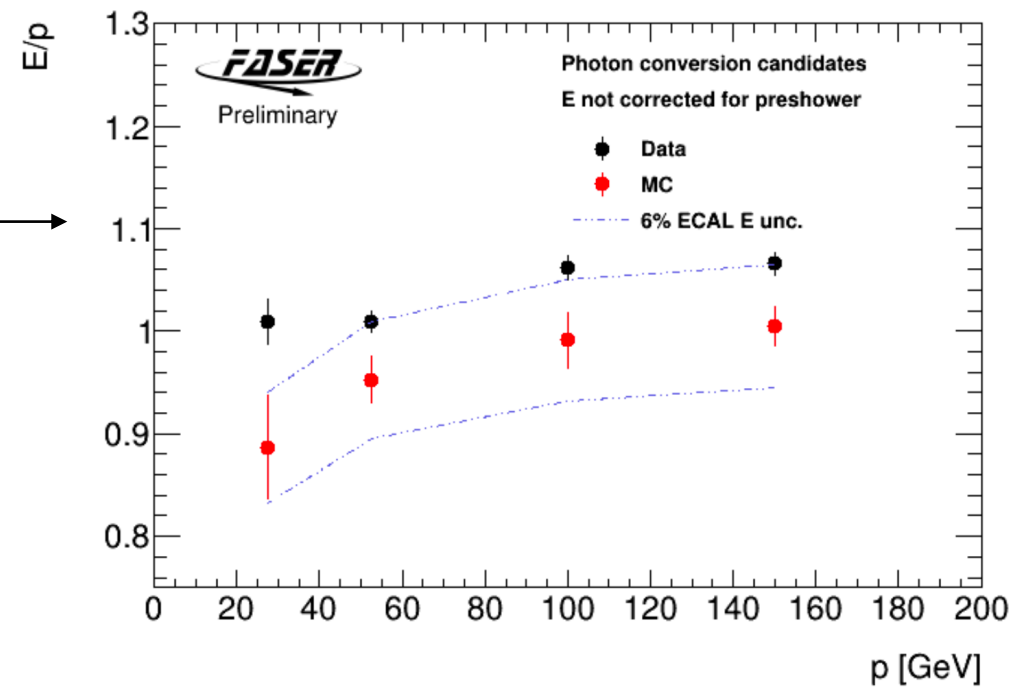
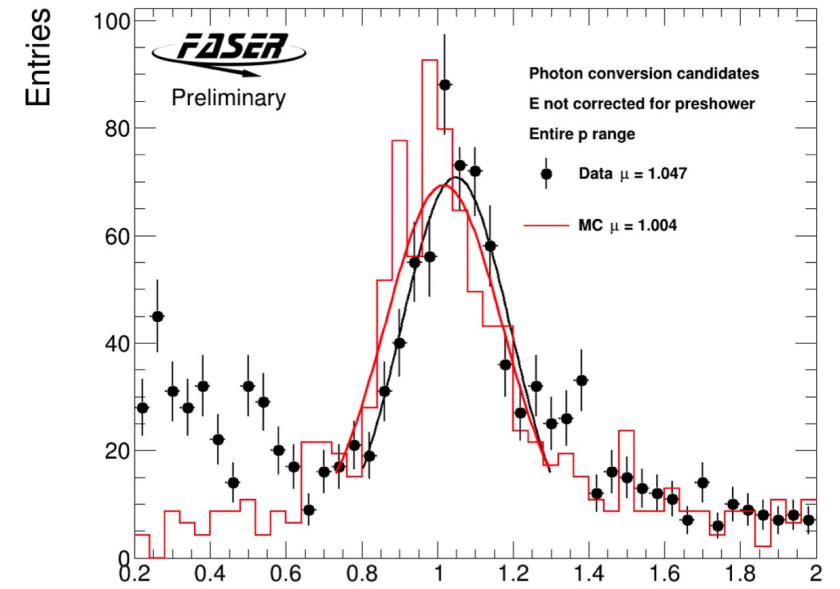
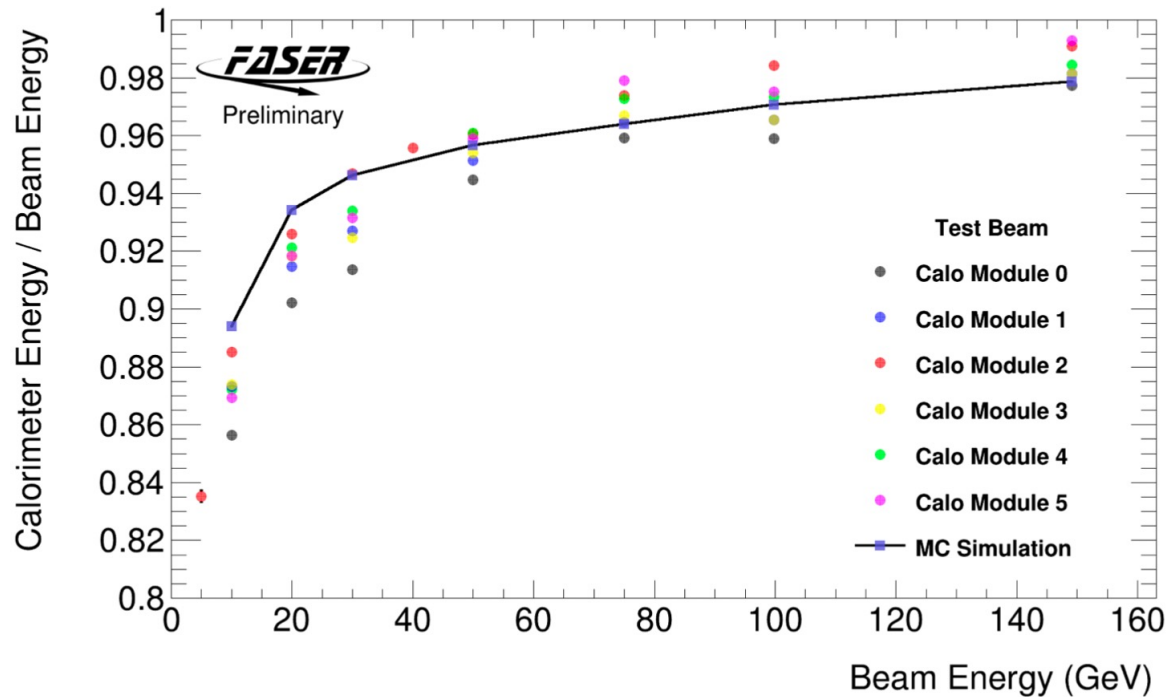


FIG. 1. Top left: The acceptance for dark photons to decay inside the FASER decay volume. Top right: The fraction of dark photons decaying inside the FASER decay volume that have energy greater than 500 GeV. Bottom: The expected number of dark photon events in FASER for  $27.0 \text{ fb}^{-1}$  of data, assuming a 50% signal efficiency, on top of the requirement that the  $A'$  energy is greater than 500 GeV.

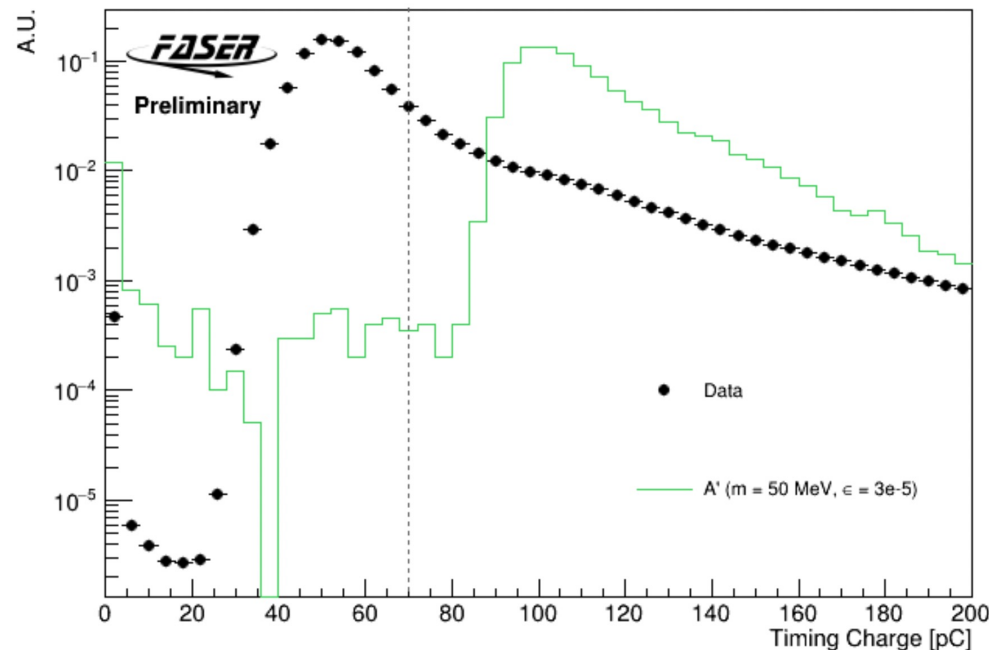
# CALO-RELATED PERFORMANCE ASSESSMENT

- Calorimeter energy scale and uncertainty estimated based on test beam data and in situ MIP calibration
- Validated using conversion events ( $\mu$  with  $e+e-$  pair)
- $E/p$  in data/MC agrees within 6%



# SELECTION IN TIMING LAYER

Selecting events with more than 70pC in timing layer is  $\sim 100\%$  efficient for signal, while also suppressing a large fraction of single track events



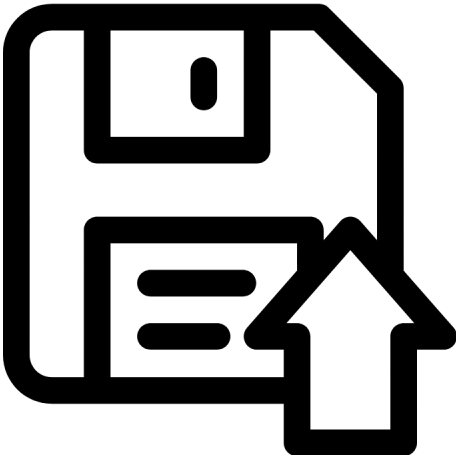
# SYSTEMATICS

Source	Value	Effect on signal yield
Theory, Statistics and Luminosity		
Dark photon cross-section	$\frac{0.15+(E_{A'}/4\text{TeV})^3}{1+(E_{A'}/4\text{TeV})^3}$	15-65% (15-45%)
Luminosity	2.2%	2.2%
MC Statistics	$\sqrt{\sum W^2}$	1-3% (1-2%)
Tracking		
Momentum Scale	5%	< 0.5%
Momentum Resolution	5%	< 0.5%
Single Track Efficiency	3%	3%
Two-track Efficiency	15%	15%
Calorimetry		
Calo E scale	6%	0-8% (< 1%)



# YIELDS

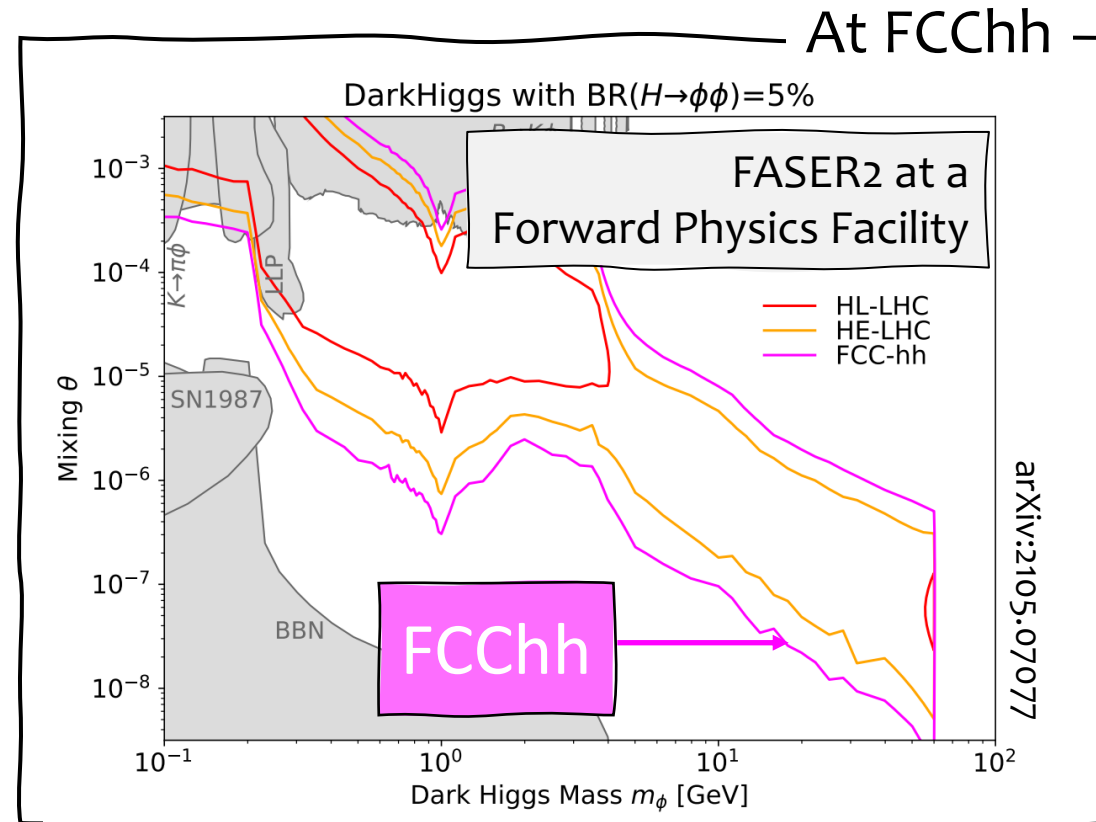
Cut	Data	
	Events	Efficiency
Good collision event	151750788	—
No Veto Signal	1235830	0.814%
Timing/Preshower Signal	313988	0.207%
$\geq 1$ good track	21329	0.014%
= 2 good tracks	0	0.000%
Track radius $< 95$ mm	0	0.000%
Calo energy $> 500$ GeV	0	0.000%



**BEYOND**



# OPPORTUNITIES FOR FIPs AT FCC



Significant opportunities open up, beyond what can be done with conventional collider detectors!  
Essential to account for them since the beginning, to minimize overheads later on.