

# Detectors for future neutrino facilities

## Missions of ISS study:

"--Evaluate the options for the neutrino detection systems "Provide a research-and-development program

Funding request for four years of detector R&D "2008-2011" (In Europe: DEvDET) IDS--> Neutrino factory detectors (Magnetized) NNN--> non magnetic detectors (Larg, Water Cherenkov)

The nice thing with neutrino beams is that one can have more than one detector on the same beam line!







	Beta beam <sup>18</sup> Ne: $\nu_e \rightarrow \nu_\mu$	T violation	Superbeam $\pi^+$ : $\nu_{\mu} \rightarrow \nu_{e}$
Superbeam & beta-beam: Non-MAGNETIC	CP violation	CPT	CP violation
	Beta beam $^{\circ}$ He: $\overline{\nu}_{e} \rightarrow \overline{\nu}_{\mu}$	T violation	Superbeam $\pi^+$ : $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\theta}$

## Nu-Fact: MAGNETIC

$\mu^{\scriptscriptstyle +} \to e^{\scriptscriptstyle +} \nu_{_{\boldsymbol{\theta}}} \overline{\nu}_{_{\boldsymbol{\mu}}}$	$\mu^- \to e^- \nu_\mu \overline{\nu}_e$	reaction		
$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\mu}$	$\nu_{\mu} \rightarrow \nu_{\mu}$	CC	Disappearance	
$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\theta}$	$\nu_{\mu} \rightarrow \nu_{e}$	СС	Appearance ('platinum' channel)	
$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\tau}$	$\nu_{\mu} \rightarrow \nu_{\tau}$	CC	Appearance (atmospheric oscillation)	
$\nu_{_{\vartheta}} \rightarrow \nu_{_{\vartheta}}$	$\overline{\nu}_{e} \rightarrow \overline{\nu}_{e}$	CC	Disappearance	
ν <sub>e</sub> →ν <sub>μ</sub>	$\nabla_{\Theta} \rightarrow \nabla_{\mu}$	CC	Appearance: 'golden' channel	
ν <sub>e</sub> →ν <sub>τ</sub>	$\overline{\nu}_e^{} \rightarrow \overline{\nu}_\tau^{}$	CC	Appearance: 'silver' channel	
$\nu \to \nu_{_B}$	$\pmb{\nabla} \rightarrow \pmb{v}_{\mathfrak{s}}$	NC	Global disappearance, sterile neutrinos	



# Water Cerenkov

- -- can be made in very large volumes (already SK =50kton)
- -- very well known technology

-- other applications: proton decay, low energy natural neutrinos, atmospheric, solar and SN neutrinos, Gadzooks, etc...

- -- cannot be magnetized easily
- -- pattern recognition ~limited to 1 ring events (--> sub GeV neutrinos)
- -- baseline detector for sub-GeV neutrinos.
- -- three projects around the world: HK, UNO, MEMPHYS
- -- community organized and coordinated in its own

cost estimates range from 0.5 G\$ (HK) to 1G  $\in$  (MEMPHYS) for 1 MTon





#### **The MEMPHYS Project**





## **R&D on electronics (ASICs)**

Integrated readout : "digital PM (bits out)" Charge measurement (12bits) Time measurement (1ns) Single photoelectron sensitivity

High counting rate capability (target 100 MHz)

Large area pixellised PM : "PMm<sup>2</sup>" 16 low cost PMs Centralized ASIC for DAQ Variable gain to have only one HV

Multichannel readout

Gain adjustment to compensate non uniformity Subsequent versions of OPERA\_ROC ASICs

aim at 200 euros/channel







**Mechanics & PMT tests** 

Taken in charge by IPNO: well experienced in photodetectors (last operation: Auger). With PHOTONIS tests of PMT 8", 9"  $\rightarrow$  12" and Hybrid-PMT and HPD



Electronic box water tight

Basic unit that we want to build and test under water



IPNO



A possible schedule for MEMPHYS at Frejus





# Superbeam + beta beam together



2 ways of testing CP, T and CPT : redundancy and check of systematics





#### Possible experimental set-up



# 'Standard' Next Neutrino Oscillation Experiments

- Aim to study CPV, Mass hierarchy
- Megaton Scale Detector + Upgraded Accelerator
- Typical → Detector 0.5 Mton (fiducial Volume)



- Other Subjects
  - Proton Decay (10<sup>35</sup> years for  $e\pi^0$ )
  - SN neutrinos



Suzuki Neutrino 2008



# fact

# **5 Mt Neutrino Oscillation Detector**

- Proton decay search ~10<sup>36</sup>yr
- SN neutrino detection: ~1 every year
  - Reaches 5Mpc w/ ~ 5 events
- ➔ PD and SN really add the value to the experiment
- Precise atmospheric neutrino measurements
- Flexible location of the detector for a long baseline neutrino oscillation experiment
- Effective investment: accelerator or detector
  - More on detector
- possibility to find unexpected
- Many technical challenges
- Need to start R&D now for a detector of more than 20 year from now



## -- Liquid Argon TPC:

DOE (detector of everything) it can do everything, can it do it BETTER than a dedicated standard technique? Case by case...

impressive progress from ICARUS T600

recent highlights

-- observation of operation in magnetic field

-- programme on-going to demonstrate long drift, or long wires

100 kton fid... ? When, how, how much?

Trade off between mass and detector quality

Main advantage (in superbeam like T2K) piO/electron separation Less advantageous for betabeam (low energy mu/pi separation)





# An ideal detector exploiting a Neutrino Factory should:

Identify and measure the charge of the muon ("golden channel") with high accuracy

Identify and measure the charge of the electron with high accuracy ("Platinum channel")

Identify the  $\tau$  decays ("silver channel")

Measure the complete kinematics of an event in order to increase the signal/back ratio



## -- Magnetized Iron Neutrinofactory Detector\*)

this is a typical NUFACT detector for  $E_v$ >1.5 GeV



GOLDEN CHANNEL

experience from MINOS & NOvA designs prepared for Monolith and INO

iron-scintillator sandwich with sci-fi + APD read-out

proposed straightforward design 90kton for ~175M\$ (Nelson)









Event rates for 10<sup>21</sup> muon decays for 50 GeV beam

Baseline	$\overline{\mathbf{v}}_{\!\mu}~\mathbf{C}\mathbf{C}$	$v_e$ CC	$ν_{\mu}$ signal (sin² θ <sub>13</sub> =0.01)		
732 Km	10 <sup>9</sup>	2 x 10 <sup>9</sup>	3.4 x 10 <sup>5</sup>	(J-PARC I $\rightarrow$ SK = 40)	
3500 Km	4 x 10 <sup>7</sup>	7.5 x 10 <sup>7</sup>	3 x 10 <sup>5</sup>	5	





#### Multi-Pixel-Photon-Counter Operation

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#### Kudenko



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#### BASELINE SILVER DETECTOR





# LARGE MAGNETIC VOLUME

MIND + emulsions provide golden + silver with low efficiency (muon decays)

these are feasible and of established performance.

Observing the platinum channel  $V_{\mu} \rightarrow V_{e}$ 

or the silver channel  $V_e \rightarrow V_\tau$  for more decay channels

requires a dedicated

Low Z and very fine grained detector immersed in a large magnetic volumete (CFS NOMAD) Alain Blok





Possible improvement: Totally Active Scintillating Detector (TASD) using Nova and Minerva concepts

3333 Modules (X and Y plane) Each plane contains 1000 slabs Total: 6.7M channels





Momenta between 100 MeV/c to 15 GeV/c Magnetic field considered: 0.5 T Reconstructed position resolution ~ 4.5 mm Reduction threshold: access second oscillation maximum and electron identification 3 WIN07 Kolkata Alain Blork





# However, possible magnetisation can be achieved using magnetic cavern concept (10 modules with 15m x 15 m diameter)



### A superconducting magnetized LAr TPC detector



Tentative Yoke parameters						
Cylindrical Fe yoke	10 kton LAr			100 kton LAr		
Magnetic induction (T)	0.1	0.4	1.0	0.1	0.4	1.0
Magnetic flux (Weber)	70	280	710	385	1540	3850
Assumed saturation field in Fe (T)	1.8			1.8		
Thickness (m)	0.4	1.6	3.7	1	3.7	8.7
Height (m)	10		20			
Mass (kton)	6.3	25	63	34	137	342



# Cylindrical Fe yoke. (Instrumented?)

NB: Superconducting Magnetic Energy Storage (SMES) systems were considered for underground storage of MJ energy without return yoke buried in tunnels in bedrock (see e.g. Eyssa and Hilal, J. Phys. D: Appl. Phys 13 (1980) 69). Avoid using a yoke?



# μ end electron momentum resolution: 3 gaps (3cm thick) and 0.5 T

![](_page_25_Figure_1.jpeg)

For the electron only hits associated to the primary electrons used in the parabolic fit (Kalman not used) Given the non negligible energy loss in the target, the electron energy is taken downstream for the comparison of true against reconstructed

FIRST INDICATION THAT THE PLATINUM CHANNEL COULD BE USED!

![](_page_25_Picture_4.jpeg)

ISS-detectors WIN07 Kolkata

![](_page_26_Picture_0.jpeg)

# Near detectors and flux instrumentation

![](_page_26_Picture_2.jpeg)

![](_page_26_Picture_3.jpeg)

![](_page_27_Picture_0.jpeg)

#### near detector constraints for CP violation

ex. beta-beam or nufact (interchange role of  $V_e$  and  $V_u$  for superbeam)

$$\frac{P(v_e \rightarrow v_\mu) - P(\overline{v}_e \rightarrow \overline{v}_\mu)}{P(v_e \rightarrow v_\mu) + P(\overline{v}_e \rightarrow \overline{v}_\mu)} = A_{CP} \alpha \frac{\sin \delta \sin (\Delta m_{12}^2 L/4E) \sin \theta_{12} \sin \theta_{13}}{\sin^2 \theta_{13} + \text{solar term...}}$$

Near detector gives  $V_e$  diff. cross-section\*detection-eff \*flux and ibid for bkg BUT: need to know  $V_{\mu}$  and  $\overline{V}_{\mu}$  diff. cross-section\* detection-eff

with small (relative) systematic errors.

→knowledge of cross-sections (relative to each-other) required
→knowledge of flux!

![](_page_27_Picture_7.jpeg)

European Effort

- EuroNu: four year EU Design Study for "A High Intensity Neutrino Oscillation Facility in Europe" (Super-beam, Neutrino Factory, Betabeam, neutrino detectors and physics performance).
  - Neutrino detectors: study Magnetised Neutrino Iron Detector (MIND) performance for golden measurement at a neutrino factory, water Cherenkov detector for beta and super beams and near detectors for all facilities EUROv APPROVED for 4M€, begins now
- DevDet is a new Integrating Activity proposal across Europe to coordinate "Detector Development Infrastructures for Particle Physics Experiments"
  - It is a 37.8 M€ proposal to the European Union (EU) with a requested EU contribution of 11.0 M€. It has 87 participants from 21 different countries
  - It includes the luminosity-upgraded LHC (SLHC), future Linear Colliders (ILC/CLIC), future accelerator-driven neutrino facilities and B-physics facilities
  - Funding for R&D and test beams (including neutrino test beam)

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![](_page_28_Picture_8.jpeg)

![](_page_29_Picture_1.jpeg)

## SIZE matters

- -- Liquid Argon drift
- -- scintillator/fibers light transmission
- -- price per channel

PID e/mu, piO/e, (SB) pi+-/muon (BB) two track resolution Muon and electron Charge assignement (NUFACT)

==> redundant combined low energy test beam

![](_page_29_Picture_8.jpeg)

#### large area Mmegas chambers

![](_page_30_Figure_1.jpeg)