

Thermal Particle Injection in Nonlinear Diffusive Shock Acceleration

Donald C. Ellison^a, Pasquale Blasi^b, Stefano Gabici^c

(a) *Physics Dept., North Carolina State Univ., Raleigh, NC 27695, U.S.A.*

(b) *High Energy Astrophysics Group, INAF, Osservatorio Astrofisico di Arcetri, Largo E. Fermi 5, I-50125, Firenze, Italy*

(c) *Humboldt Fellow, Max-Planck-Institut fuer Kernphysik, Saupfercheckweg 1, 69117 Heidelberg, Germany*

Presenter: Don Ellison (don_ellison@ncsu.edu), usa-ellison-D-abs3-og14-poster

Particle acceleration in collisionless astrophysical shocks, i.e., diffusive shock acceleration (DSA), is the most likely mechanism for producing cosmic rays, at least below 10^{15} eV. Despite the success of this theory, several key elements, including the injection of thermal particles, remains poorly understood. We investigate injection in strongly nonlinear shocks by comparing a semi-analytic model of DSA with a Monte Carlo model. These two models treat injection quite differently and we show, for a particular set of parameters, how these differences influence the overall acceleration efficiency and the shape of the broad-band distribution function.

1. Introduction

We compare a recent semi-analytic (SA) model of non-linear diffusive shock acceleration (DSA) [2][3][4] with a well-established Monte Carlo (MC) model (e.g., [12][15][11]). Both include a thermal leakage model for injection and the non-linear (NL) backreaction of accelerated particles on the shock structure, but they do these in very different ways. Also different is the way in which the particle diffusion in the background magnetic turbulence is modeled. Our limited comparison shows that the important NL effects of compression ratios $\gg 4$ and concave spectra do not depend strongly on the injection model as long as injection is efficient. A fuller understanding of the complex plasma processes involved, particularly if injection is weak, will require particle-in-cell (PIC) simulations (e.g., [13]), but these simulations, which must be done fully in three-dimensions[14], cannot yet be run long enough, in large enough simulation spaces, to accelerate particles from thermal to relativistic energies in order to show strong NL effects. For now, approximate methods must be used.

The Monte Carlo model is more general than the semi-analytic model,¹ but it is considerably slower computationally. Since it is important in many applications, such as hydro models of supernova remnants, to include the dynamic effects of nonlinear DSA in simulations which perform the calculation many times (e.g., [8]), a rapid, approximate calculation, such as the SA one discussed here, is useful.

2. Models

The main features of the two models that are relevant for our comparison here are:

Injection: In the semi-analytic model, a free injection parameter, ξ_{SA} , determines the fraction of total particles injected into the acceleration mechanism and the injection momentum, p_{inj} . Specifically, $p_{inj} = \xi_{SA} p_{th}$, where $p_{th} = \sqrt{2m_p k T_{DS}}$ (T_{DS} is the downstream temperature). The fraction, η_{SA} , of unshocked particles crossing the shock which become superthermal in the SA model is $\eta_{SA} = (4/3\pi^{1/2})(r_{sub} - 1)\xi_{SA}^3 e^{-\xi_{SA}^2}$, where r_{sub} is the subshock compression ratio. The fraction η_{SA} , which is approximately the number of particles in the Maxwellian defined by T_{DS} with momentum $p > p_{inj}$, is determined by requiring the continuity, at p_{inj} , of the Maxwellian and the superthermal distribution (see [4]). Since p_{th} depends on the injected fraction, the solution must be obtained by iteration.

In the Monte Carlo model, the injection depends on the scattering assumptions. We assume that particles pitch-angle scatter elastically and isotopically in the local plasma frame and that the mean free path (mfp)

¹For instance, the MC model can treat a specific momentum dependence for the scattering mean free path, particle acceleration in relativistic shocks[9][10], and NL effects in oblique shocks[11].

is proportional to the gyroradius, i.e., $\lambda \propto r_g$, where $r_g = pc/(qB)$. With these assumptions, the injection is purely statistical with those “thermal” particles which manage to diffuse back upstream gaining additional energy and becoming superthermal. Note that in this scheme the viscous subshock is assumed to be transparent to all particles, even thermal ones, and that any downstream particle with $v \geq u_2$ has a chance to be injected. Here, u_2 is the downstream flow speed.

For comparison with the SA model, we have included an additional parameter, $v_{\text{thres}}^{\text{MC}}$, to limit injection in the MC simulation. Only downstream particles with $v \geq v_{\text{thres}}^{\text{MC}}$ are injected, i.e., allowed to re-cross the shock into the upstream region and become superthermal. In our previous Monte Carlo results, with the sole exception of [6], we have taken $v_{\text{thres}}^{\text{MC}} = 0$.

Momentum dependent diffusion: Diffusion is treated very differently in the two models. As just mentioned, the MC simulation models pitch-angle diffusion by assuming a $\lambda(p)$ (see [10]). The semi-analytic model does not explicitly describe diffusion but assumes only that the diffusion is a strongly increasing function of particle momentum p so that particles of different p interact with different spatial regions of the upstream precursor. Eichler [5] used a similar procedure. With this assumption, particles of momentum p can be assumed to feel some average precursor fluid speed u_p and an average compression ratio $r_p \sim u_p/u_2$ (see [4]).

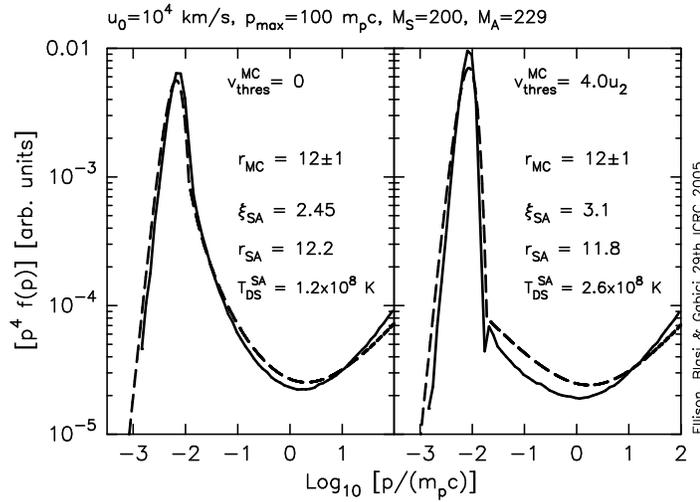


Figure 1. Phase space distributions $f(p)$ times p^4 for a MC model (solid curves) and a SA model (dashed curves). In the left panel $v_{\text{thres}}^{\text{MC}} = 0$, and in the right panel $v_{\text{thres}}^{\text{MC}} = 4u_2$ in the MC model. Here, $u_2 = u_0/r$ is the downstream plasma speed in the shock frame, u_0 is the shock speed, and the Mach numbers, compression ratios, and shocked temperatures are indicated. In all cases, the shock is parallel and Alfvén heating is assumed in the shock precursor [7].

The different way diffusion is treated influences not only injection, but also the shape of the distribution function $f(p)$ [$f(p)$ is the momentum phase space density, i.e., particles/(cm³-d³p)]. Both models give the characteristic concave $f(p)$ which hardens with increasing p and, since the overall shock compression ratio r can be greater than 4, this spectrum will be harder than p^{-4} at ultra-relativistic energies. In the results we show here, the acceleration is limited with a cutoff momentum so $f(p)$ cuts off abruptly at p_{max} . More realistic models will show the effects of escape from some spatial boundary (e.g., finite shock size) or from a finite acceleration time. In either case, the spectrum will show a quasi-exponential turnover, e.g., $f(p) \propto p^{-\sigma} \exp[-\alpha^{-1}(p/p_{\text{max}})^\alpha]$ [7], where α is included to emphasize that the detailed shape of the turnover depends on the momentum dependence of the diffusion coefficient near p_{max} .

Thermalization: There is no thermalization process in the MC simulation in the sense of particles exchanging energy between one another because particles scatter elastically in the local frame. However, a quasi-thermal low energy distribution is created as unshocked particles cross the shock from upstream to downstream at different angles and receive different fractions of the speed difference, $u_0 - u_2$. For the parameters used here, the low-energy peak is essentially a Maxwellian when $v_{\text{thres}}^{\text{MC}} = 0$.

In the semi-analytic model, the shocked thermal pressure and density are determined from the conservation relations and these are translated to a Maxwell-Boltzmann distribution.

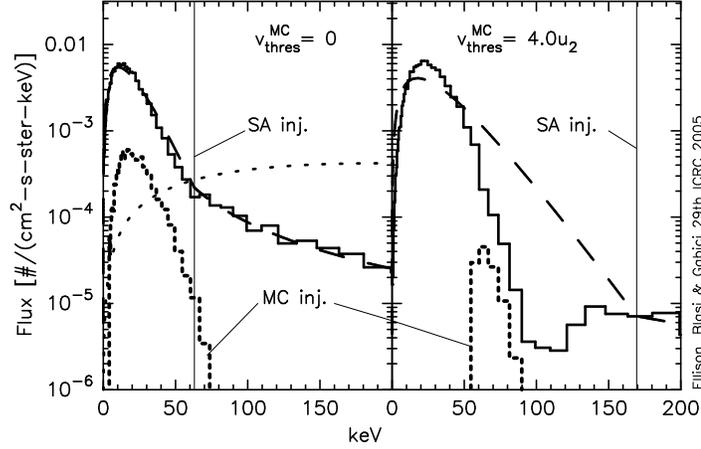


Figure 2. Low energy portions of the spectra shown in Fig. 1 where the solid histograms are the MC results and the dashed curves are the SA results. The heavy dotted histograms show the distribution of particles that were injected in the MC model. The solid vertical lines are drawn at p_{inj} , the momentum at which particles are injected in the SA model. For $v_{\text{thres}}^{\text{MC}} = 0$, $\eta_{\text{MC}} = 5.7 \times 10^{-2}$ and $\eta_{\text{SA}} = 4.6 \times 10^{-2}$, while for $v_{\text{thres}}^{\text{MC}} = 4u_2$, $\eta_{\text{MC}} = 4.1 \times 10^{-3}$ and $\eta_{\text{SA}} = 2.5 \times 10^{-3}$. The light-weight dotted curve shows the Maxwellian ($T_{\text{DS}} = 2.2 \times 10^9$ K) that would have resulted if no DSA occurred.

3. Results

In Fig. 1 we show distribution functions (times p^4) for a set of parameters and with $v_{\text{thres}}^{\text{MC}} = 0$ (left panel) and $v_{\text{thres}}^{\text{MC}} = 4u_2$ (right panel). In both panels, the solid curves are the MC results and the dashed curves are the SA results where ξ_{SA} has been chosen to provide the best match to the MC spectra.² In Fig. 2 we show the low energy portions of the spectra along with the distributions of injected particles in the MC model (dotted histograms). The vertical solid lines are at p_{inj} and they show the transition between the Maxwellian and the superthermal population in the SA results.

The important aspects of the plots are:

- (i) The broad-band match between the two very different calculations is quite good, particularly for $v_{\text{thres}}^{\text{MC}} = 0$. Both models show the important characteristics of nonlinear DSA, i.e., $r \gg 4$, concave spectra, and a sharp reduction in the shocked temperature from test-particle values.
- (ii) The shocked temperature depends on $v_{\text{thres}}^{\text{MC}}$ with weaker injection (i.e., larger $v_{\text{thres}}^{\text{MC}}$) giving a larger T_{DS} .³

²Note that once the input parameters (i.e., u_0 , p_{max} , $v_{\text{thres}}^{\text{MC}}$, and the sonic and Alfvén Mach numbers: M_S and M_A) are fixed, the only parameter varied to match the spectra shown in Figs. 1 and 2 is ξ_{SA} . The normalization is not adjusted.

³The light-weight dotted curve in the left panel of Fig. 2 is a Maxwellian at $T_{\text{DS}} = 2.2 \times 10^9$ K, the temperature of the shocked gas without DSA.

(iii) The distribution $f(p)$ is harder near p_{\max} in the MC results than with the SA calculations. A common prediction of SA models is that the shape of the particle spectra at p_{\max} has the form $p^{-3.5}$ if the shock is strongly modified and the diffusion coefficient grows fast enough in momentum. This can be demonstrated by solving the equations in the extreme case of a maximally modified shock and approximating the spectrum with a power law at high momenta. The MC model makes no such power-law assumption.

(iv) The minimum in the $p^4 f(p)$ plot occurs at $p > m_p c$ in both models. The transition between $f(p)$ softer than p^{-4} and $f(p)$ harder than p^{-4} varies with shock parameters and increases as p_{\max} increases. This is an important difference from the algebraic model of Berezhko & Ellison[1] where the minimum is fixed at $m_p c$.

(v) For the particular parameters used in these examples, the overall compression ratio is relatively insensitive to $v_{\text{thres}}^{\text{MC}}$, but shocks having other parameters may show a greater sensitivity. Note that Alfvén wave heating is assumed in all of the results presented here. If only adiabatic heating was assumed, the compression ratios would be much higher (see [7]).

(vi) In all cases, the MC model injects more particles than the SA model but the average energy of the injected particles is less, as indicated by the peak of the curve labelled ‘MC inj’ vs. the ‘SA inj’ energy in Fig. 2.

(vii) In contrast to $v_{\text{thres}}^{\text{MC}} = 0$, the ‘thermal’ part of $f(p)$ with $v_{\text{thres}}^{\text{MC}} = 4u_2$ (Fig. 2) shows large differences in the two models. While both conserve particle, momentum, and energy fluxes so that the broad-band $f(p)$ matches well for a wide range of parameters, the different treatments of the subshock lead to large differences in the critical energy range $2 \lesssim [E/(kT_{\text{DS}})] \lesssim 5$. This offers a way to distinguish these models observationally.

4. Discussion and Conclusions

Using two approximate acceleration models, we have shown that the most important features of NL DSA, i.e., $r \gg 4$ and concave spectra, are robust and do not strongly depend on the injection model as long as injection is efficient. If injection is weak, as might be the case in highly oblique shocks, accelerated spectra will depend more on the details of injection, at least in the transition range between thermal and superthermal energies. Also, the relative efficiencies for injecting and accelerating electrons vs. protons or protons vs. heavier ions may require a more detailed description of injection, as may be provided by future PIC simulations.

Acknowledgements: D.C.E. wishes to acknowledge support from a NASA grant (ATP02-0042-0006) and S.G. acknowledges support from the Humboldt foundation.

References

- [1] Berezhko, E. G., & Ellison, D. C. 1999, ApJ, 526, 385
- [2] Blasi, P. 2002, Astropart. Phys., 16, 429
- [3] Blasi, P. 2004, Astropart. Phys., 21, 45
- [4] Blasi, P., Gabici, S., & Vannoni, G. 2005, MNRAS, in press
- [5] Eichler, D. 1984, ApJ, 277, 429
- [6] Ellison, D. C. 1985, J.G.R., 90, 29
- [7] Ellison, D. C., Berezhko, E. G., & Baring, M. G. 2000, ApJ, 540, 292
- [8] Ellison, D. C., Decourchelle, A., & Ballet, J. 2004, A&A, 413, 189
- [9] Ellison, D. C. & Double, G. P. 2002, Astropart. Phys., 18, 213
- [10] Ellison, D. C. & Double, G. P. 2004, Astropart. Phys., 22, 323
- [11] Ellison, D. C., & Jones, F. C., & Baring, M. G. 1999, ApJ, 512, 403
- [12] Ellison, D. C., & Moebius, E., & Paschmann, G. 1990, ApJ, 352, 376
- [13] Giacalone, J. and Ellison, D. C. 2000, J.G.R., 105, 12,541
- [14] Jones, F. C., Jokipii, J. R. & Baring, M. G. 1998, ApJ, 509, 238
- [15] Jones, F.C., & Ellison, D.C. 1991, Space Sci. Rev., 58, 259