

Search for Relic Axions with the CAST Dipole magnet at CERN

Lino Miceli

IBS Center for Axion and Precision Physics Research (CAPP/IBS)
At the Korea Advanced Institute of Science and Technology (KAIST)
Daejeon, Republic of Korea

University of Geneva
Physics Department
7 December 2016



CAPP

Center for Axion and Precision Physics Research at KAIST, Daejeon, South Korea



Korea
Advanced
Institute of
Science and
Technology

Established October 2013. Goal: ~ 60 people within the first 3-5 years

- 15 research fellows
- 20 grad. Students
- 10 junior/senior staff
- Visitor program
- Engineers, Technicians

CAPP, April 2016



Outline

- Axion search motivation
- A common DM axion detection method
- Status of cold dark matter axion searches
- Search in a dipole magnetic field
- The CAST-CAPP/IBS Project at CERN

CAPP Program

Focus is on two key issues of contemporary physics:

- Nature of dark matter
 - Comprehensive axion search program (~75%)
- Baryon asymmetry of the universe
 - R&D towards a proton EDM measurement in a storage ring (~25%)

CAPP main goals in axion search

- Establish state of the art axion dark matter experimental program at KAIST
- R&D program to improve on all experimentally accessible parameters

- **Promote/contribute-to international collaborations**

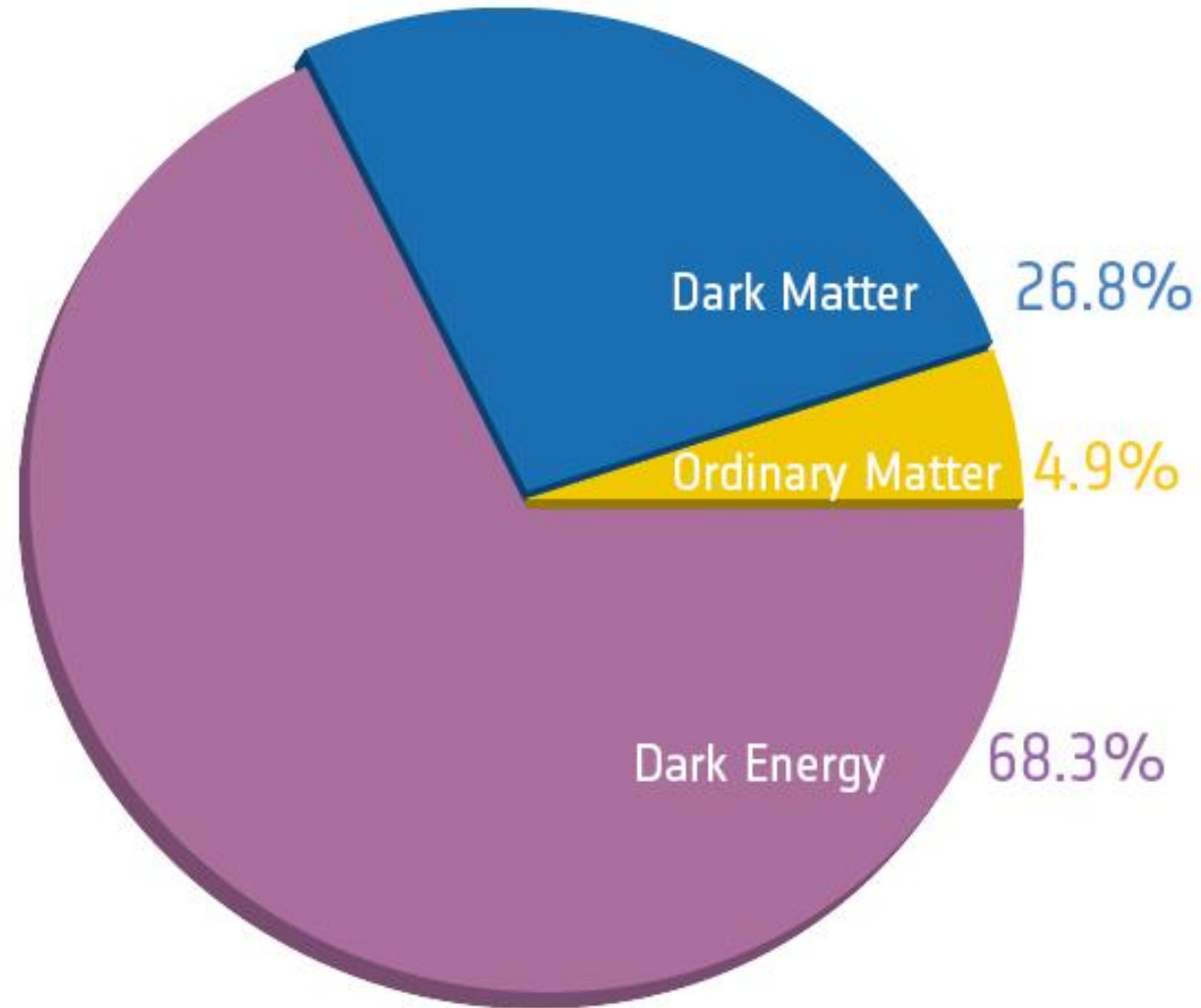
CAPP is leading a new experiment (CAST-CAPP/IBS project), within the CERN Solar Axion Telescope (CAST) collaboration at CERN, to search for cold dark matter axions with rectangular cavities in the CAST dipole magnet

Axion search motivation

The axion is a good cold dark matter candidate if its mass is in the range $\approx (1 - 100) \mu\text{eV}$.

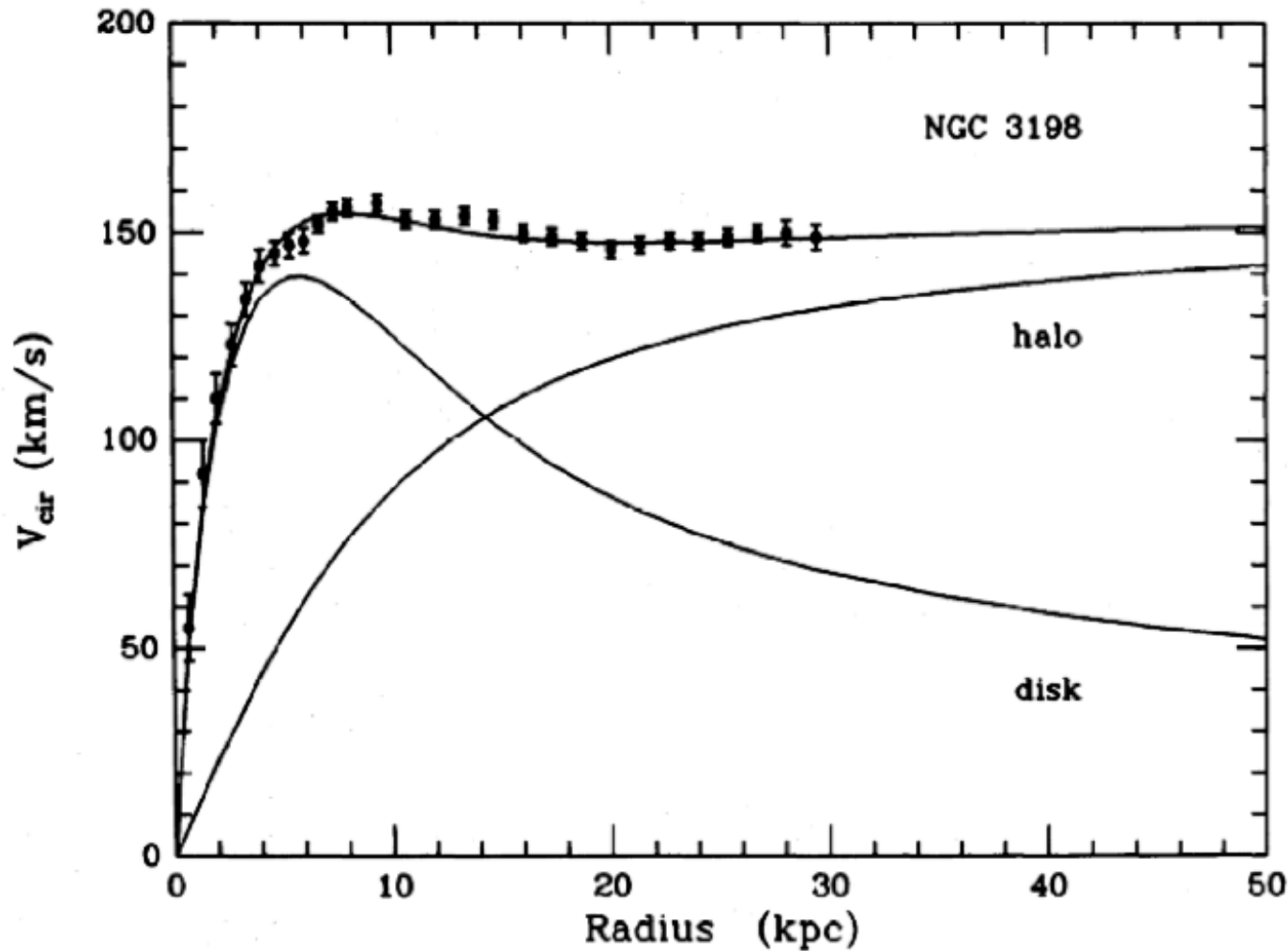
An axion discovery would solve the so called strong CP problem.

The composition of our universe



http://www.esa.int/spaceinimages/Images/2013/03/Planck_cosmic_recipe

Evidence for DM



Rotational curves in spiral galaxies.

At large distances from the center of galaxies, stars have higher than expected orbital velocities.

$$v = \left(\frac{G M(r)}{r} \right)^{1/2}$$

Circular orbit

F. Zwickey, Helv. Phys. Acta. 6, 110 (1933)

The strong CP problem

$$\mathcal{L}_{QCD} = \dots + \theta \frac{\alpha_s}{8\pi} G\tilde{G} \quad \leftarrow \quad \text{this term violates CP}$$

Measuring the EDM of a particle is a good test of CP violation

An estimate of the neutron EDM:

$$d_n \approx 3.6 \times 10^{-16} \theta \text{ e} \cdot \text{cm}$$

M. Pospelov, A. Ritz, Ann. Phys. 318 (2005) 119.

Whereas the experiment sets a limit at

C. A. Baker, et al, hep-ex/0602020

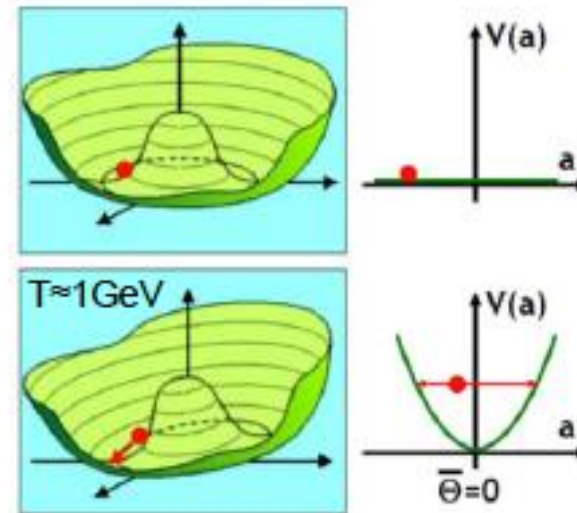
$$d_n < 2.9 \times 10^{-26} \text{ e} \cdot \text{cm} \quad \longrightarrow \quad \theta < 10^{-10}$$

The measured value of the θ parameter is ~ 10 orders of magnitude smaller than QCD predictions

Solution to the strong CP problem

- Peccei-Quinn : θ_{QCD} is a dynamical variable (1977).
 - Wilczek and Weinberg: axion particle (1977)
 - J.E. Kim: Hadronic (invisible) axions (1979)
- Axions: pseudoscalars, similar to pions, but much lighter

$$m_a \approx 6 \times 10^{-6} \text{ eV} \frac{10^{12} \text{ GeV}}{f_a}$$



The link between axion and DM

- Dark matter density: $0.3-0.5 \text{ GeV/cm}^3$
- Axions in the $1-100 \mu\text{eV}$ range: $10^{12}-10^{14}/\text{cm}^3$
 - Light axion would have condensed with the right density to be DM
- Lifetime $\sim 7 \times 10^{44} \text{ s} \left(100 \mu\text{eV} / m_a\right)^5$.
- Axions interact very weakly
- Kinetic energy $\sim 10^{-6} m_a$ (“cold”) very narrow line in spectrum.



can be treated as a classical system.

Common detection method

Based on the axion coupling to two photons.

In the presence of a strong magnetic field the conversion probability is enhanced (similar to Primakoff (*) effect)

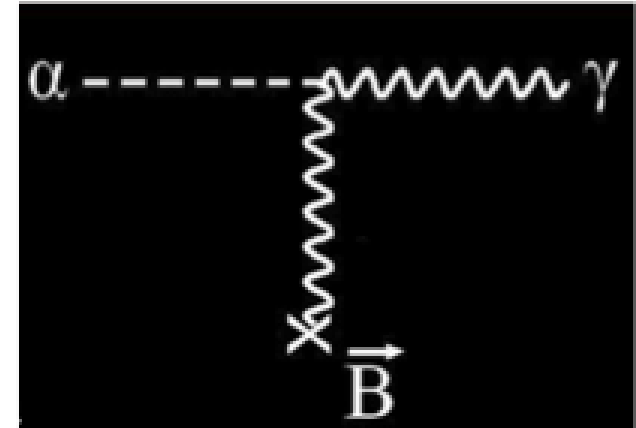
$$\mathcal{L} = -g_{a\gamma\gamma} a(t) \mathbf{E}(t) \cdot \mathbf{B}$$

$g_{a\gamma\gamma}$ coupling constant

$a(t)$ axion field

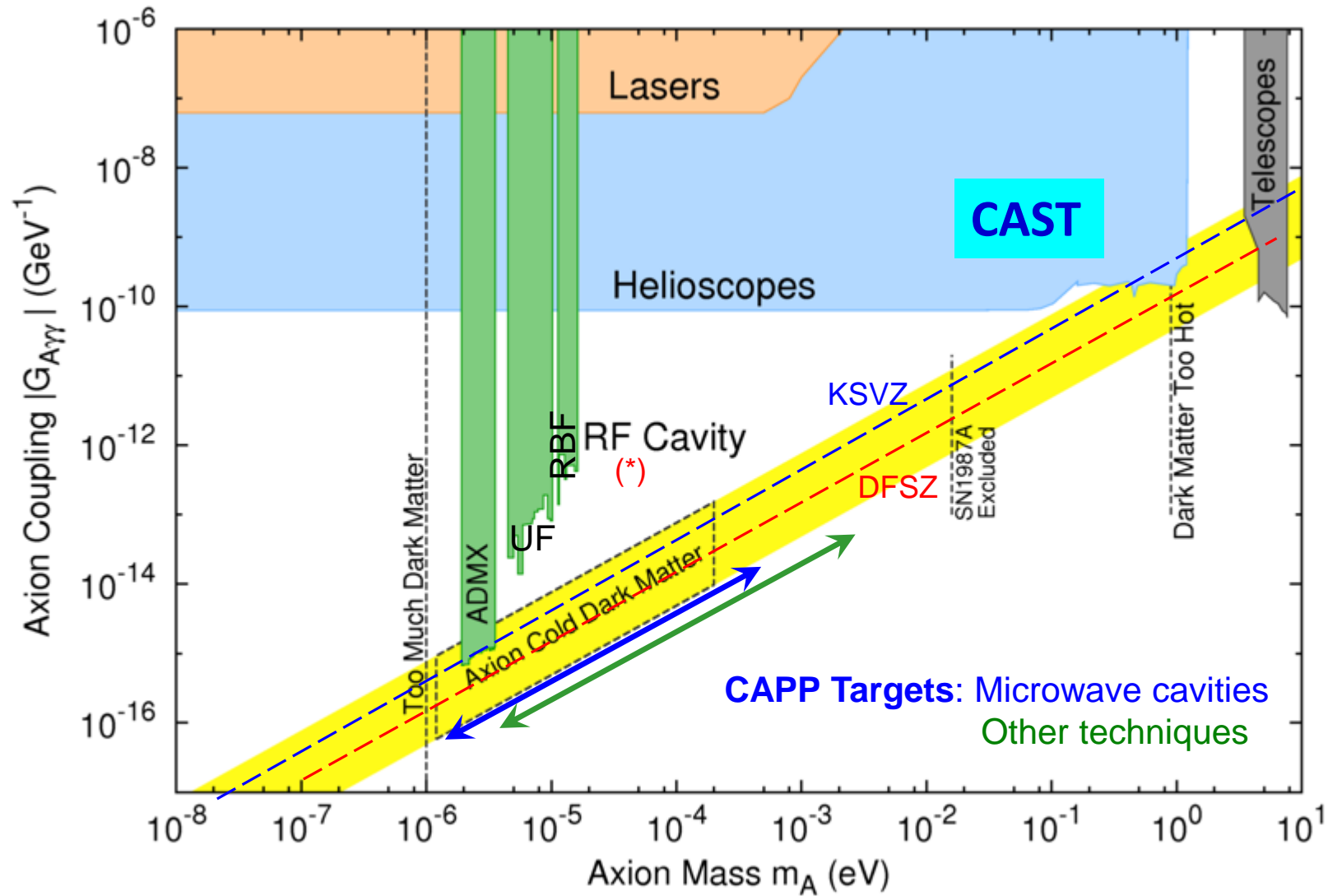
\mathbf{B} provides a virtual photon enhancing the conversion probability

$\mathbf{E}(t)$ electric field associated with the outgoing photon



(*) H. Primakoff, Phys. Rev. 81, 899 (1951)

Axion experiments: two parameter searches



(*) One missing experiment. See later.

Axion detection with microwave cavities

Axion-to-photon conversion probability further enhanced in a microwave cavity that resonates to the frequency of the axion mass (Sikivie (*)).

On-resonance axion conversion power in a microwave cavity:

$$P \approx g_{a\gamma\gamma}^2 \left(\frac{\rho_a}{m_a} \right) \mathbf{B}^2 \cdot \mathbf{Q} \cdot \mathbf{V} \cdot \mathbf{C}$$

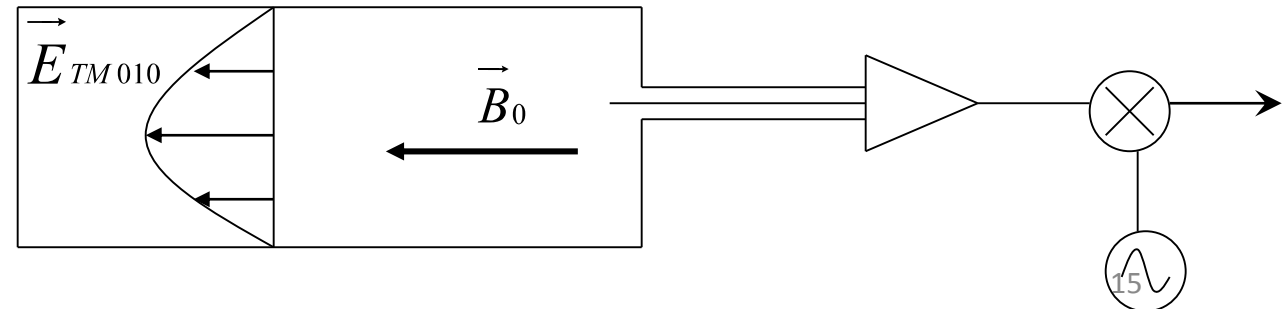
$$Q = 2\pi f \frac{\text{Stored Energy}}{\text{Power Loss}}$$

Quality factor

$$C = \frac{1}{B_0^2 V} \frac{|\int \mathbf{B} \cdot \mathbf{E} d^3 x|^2}{\int \mathbf{E} \cdot \mathbf{E} d^3 x}$$

Geometry factor

→ Tunable cavities



(*) P. Sikivie, Phys. Rev. Lett. 51, 1415 (1983).

CAPP/IBS axion target plan

- Major improvement elements:
 - High field solenoid magnets, $B: 9\text{T} \rightarrow 25\text{T} \rightarrow 40\text{T}$
 - High volume magnets/cavities, $V: 5\text{l} \rightarrow 50\text{l}$
 - High quality factor of cavity, $Q: 10^5 \rightarrow 10^6$
 - Low noise amplifiers, $T_N: 2\text{K} \rightarrow 0.25\text{K}$
 - Low physical temperature, $T_{\text{ph}}: 1\text{K} \rightarrow 0.1\text{K}$

Scanning rate improvement: 25×10^6

Improvement in coupling constant: 70

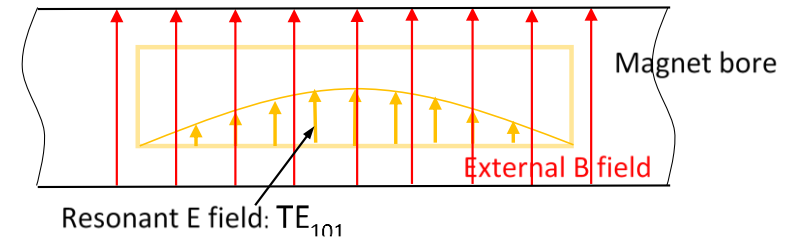
Axion dark matter in the mass range $\sim 1\mu\text{eV}$ to $100\mu\text{eV}$. Plan to either detect or exclude axions down to 10% of dark matter.

CAST-CAPP/IBS search: rectangular geometry (*)

First experiment using rectangular cavities in a dipole magnet

- Rectangular cavity resonant frequencies

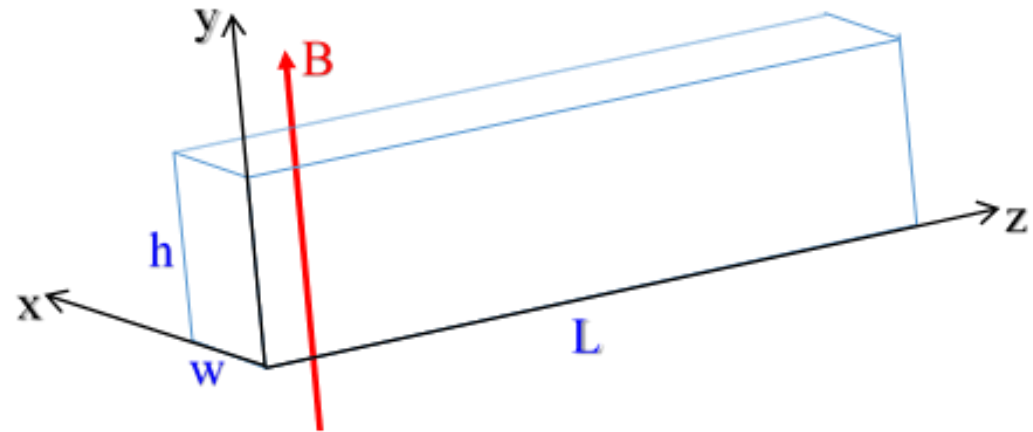
$$f_{lmn} = \frac{c}{2\pi\sqrt{\mu_r\epsilon_r}} \sqrt{\left(\frac{l}{w}\right)^2 + \left(\frac{m}{h}\right)^2 + \left(\frac{n}{L}\right)^2}$$



- Resonant E field aligned with the external B field: TE_{l0n} modes

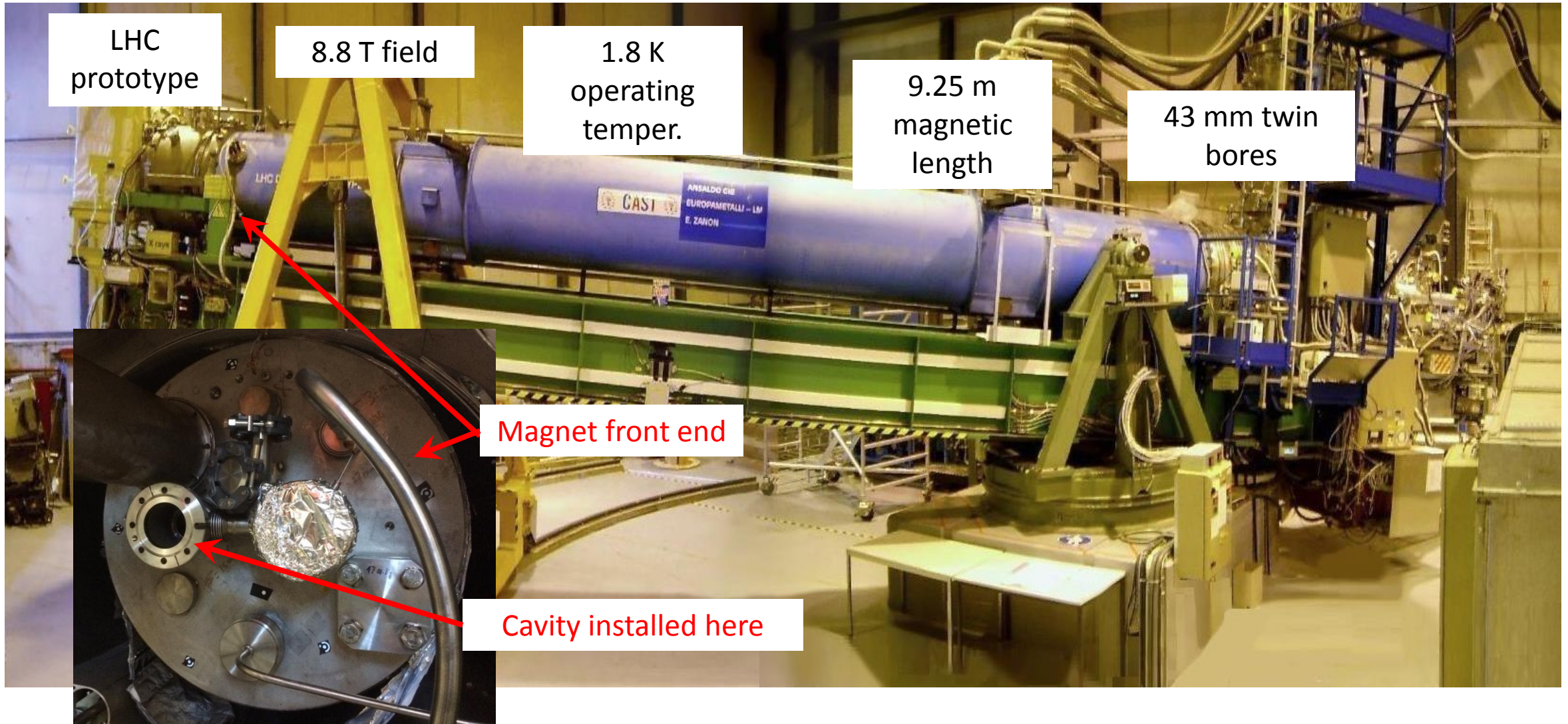
$$E_y \sim \sin\left(\frac{l\pi}{h}x\right) \sin\left(\frac{n\pi}{L}z\right)$$

$$E_x = E_z = 0$$



(*) O. Baker et al., Phys. Rev. D85, 035018 (2012)

CAST-CAPP/IBS Search: The CAST Dipole Magnet



Resonant power

$$\begin{aligned} \bullet P &= (g_{a\gamma\gamma})^2 \rho_a \frac{1}{m_a} B^2 C V \min[Q_c, Q_a] \\ &= 1.6 \times 10^{-23} \text{W} \times (g_{a\gamma\gamma} 10^{14} \text{GeV})^2 \left(\frac{\rho_a}{300 \text{ MeV/cm}^3} \right) \left(\frac{2.4 \times 10^{-5} \text{eV}}{m_a} \right) \\ &\quad \times \left(\frac{B}{9 \text{ T}} \right)^2 \left(\frac{C}{0.66} \right) \left(\frac{V}{5 \text{ l}} \right) \left(\frac{Q}{5 \times 10^3} \right) \end{aligned}$$

- $m_a = 24 \mu\text{eV}$ ($f = 5.8 \text{ GHz}$)
- $B = 9 \text{ T}$, CAST magnet
- $V = 5 \text{ liters}$
- $Q = \min[Q_c, Q_a] = Q_0/2 \sim 5,000$; critical coupling
 - Q_c loaded quality factor
 - Q_0 cavity quality factor

Time required for a single measurement (example)

$$SNR = \frac{P}{K_B T} \sqrt{\frac{t}{b}}$$

$$t = 9 \times 10^5 s \left(\frac{SNR}{4}\right)^2 \left(\frac{T}{3.8 \text{ K}}\right)^2 \left(\frac{C}{0.66}\right)^{-2} \left(\frac{B}{9 \text{ T}}\right)^{-4} \left(\frac{V}{5 \text{ l}}\right)^{-2}$$

$$\times (g_{a\gamma\gamma} 10^{14} \text{ GeV})^{-4} \left(\frac{\rho_a}{300 \text{ MeV/cm}^3}\right)^{-2} \left(\frac{2.4 \times 10^{-5} \text{ eV}}{m_a}\right)^{-3}$$

$$\times \left(\frac{Q}{5 \times 10^3}\right)^{-2} \left(\frac{10^6}{Q_a}\right) \sim 10 \text{ days, } g_{a\gamma\gamma} = 10^{-14} \text{ GeV}^{-1}$$

$m_a = 24 \mu\text{eV}$ ($f = 5.8 \text{ GHz}$) ; $B = 9 \text{ T}$, CAST magnet

$V = 5 \text{ liters}$

$Q = \min[Q_c, Q_a] = Q_0/2 \sim 5,000$; critical coupling

Q_c loaded quality factor

Q_0 cavity quality factor

$T =$ System Temperature = physical temperature + receiver&lifier-chain equivalent noise temperature. Commercial HEMT amplifiers.

Scanning rate (example)

$$\begin{aligned} \frac{df}{dt} = \frac{f}{Q} \frac{1}{t} &\approx \frac{3.4 \text{ KHz}}{\text{year}} (g_{a\gamma\gamma} 10^{15} \text{ GeV})^4 \left(\frac{5.8 \text{ GHz}}{f}\right)^2 \left(\frac{4}{\text{SNR}}\right)^2 \left(\frac{3.8 \text{ K}}{T}\right)^2 \\ &\times \left(\frac{B}{9 \text{ T}}\right)^4 \left(\frac{C}{0.66}\right)^2 \left(\frac{V}{5 \text{ l}}\right)^2 \left(\frac{Q}{5 \times 10^3}\right) \\ &\sim 90 \text{ MHz /year at } g_{a\gamma\gamma} 10^{-14} \text{ GeV}^{-1} \end{aligned}$$

$m_a = 24 \mu\text{eV}$ ($f = 5.8 \text{ GHz}$) ; $B = 9 \text{ T}$, CAST magnet;

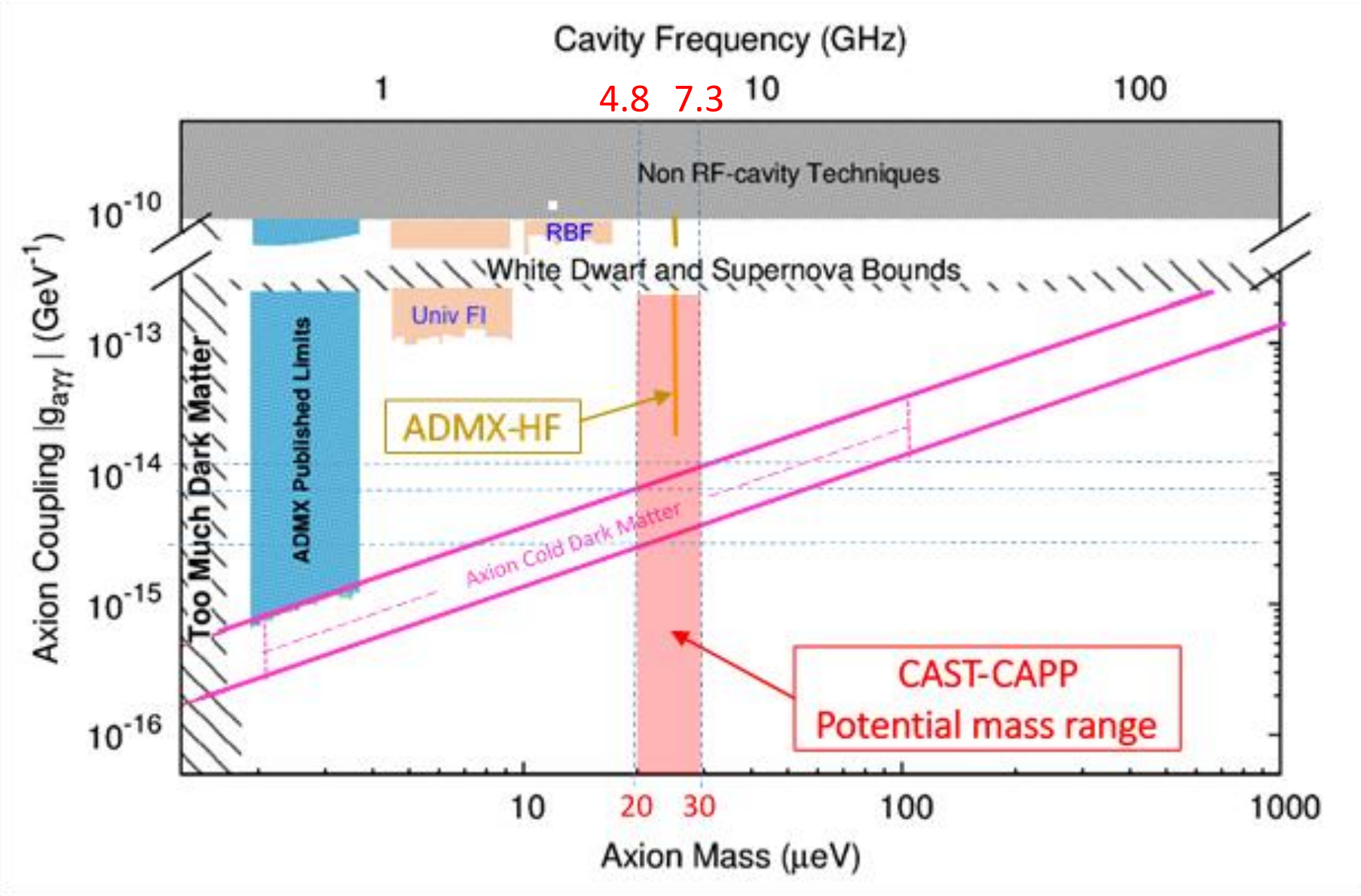
$V = 5 \text{ liters}$

$Q = \min[Q_c, Q_a] = Q_0/2 \sim 5,000$; critical coupling

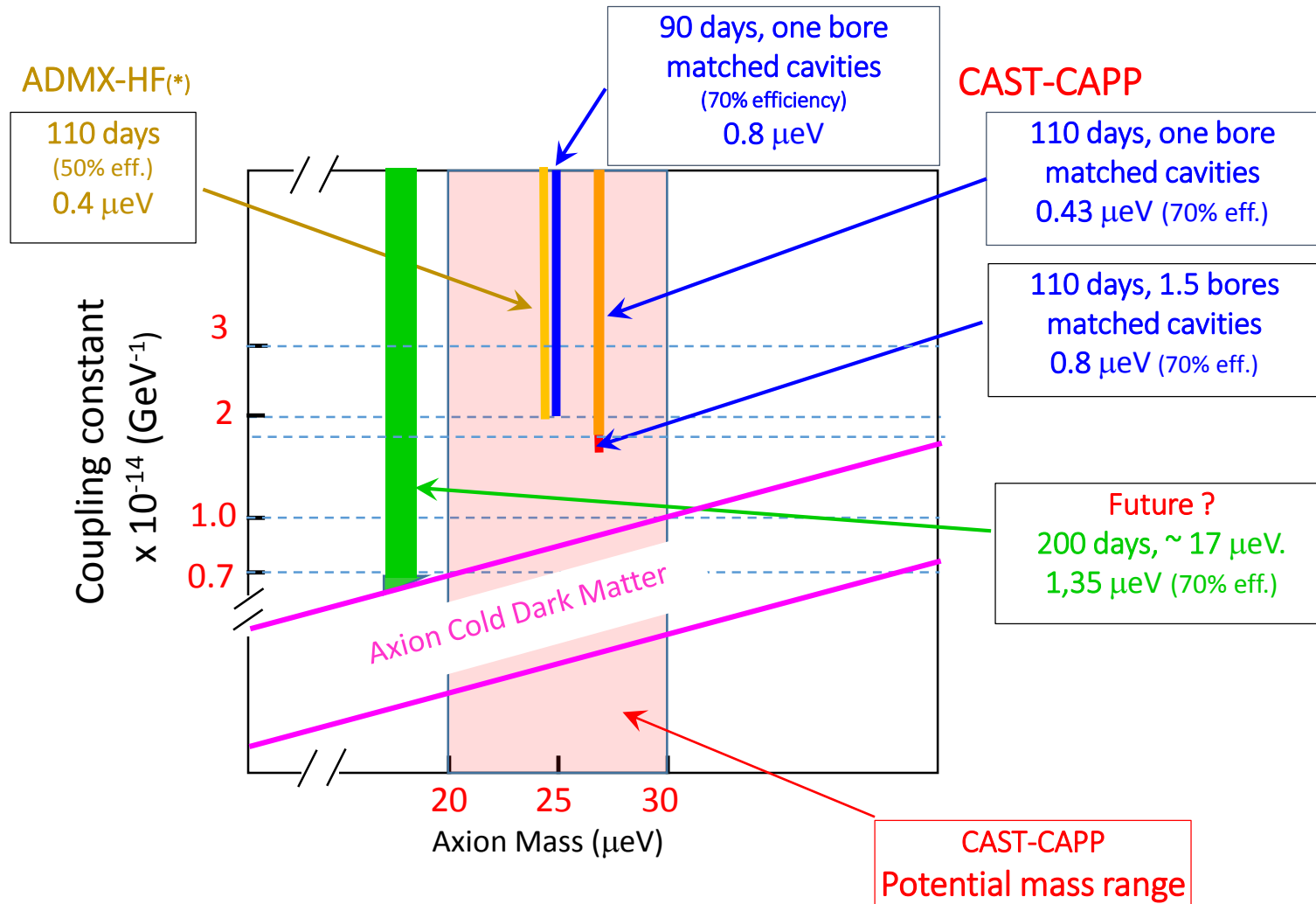
- Q_c loaded quality factor
- Q_0 cavity quality factor

T : System Temperature = physical temperature + receiver&lifier-chain equivalent noise temperature.

Sensitivity



CAST-CAPP sensitivity

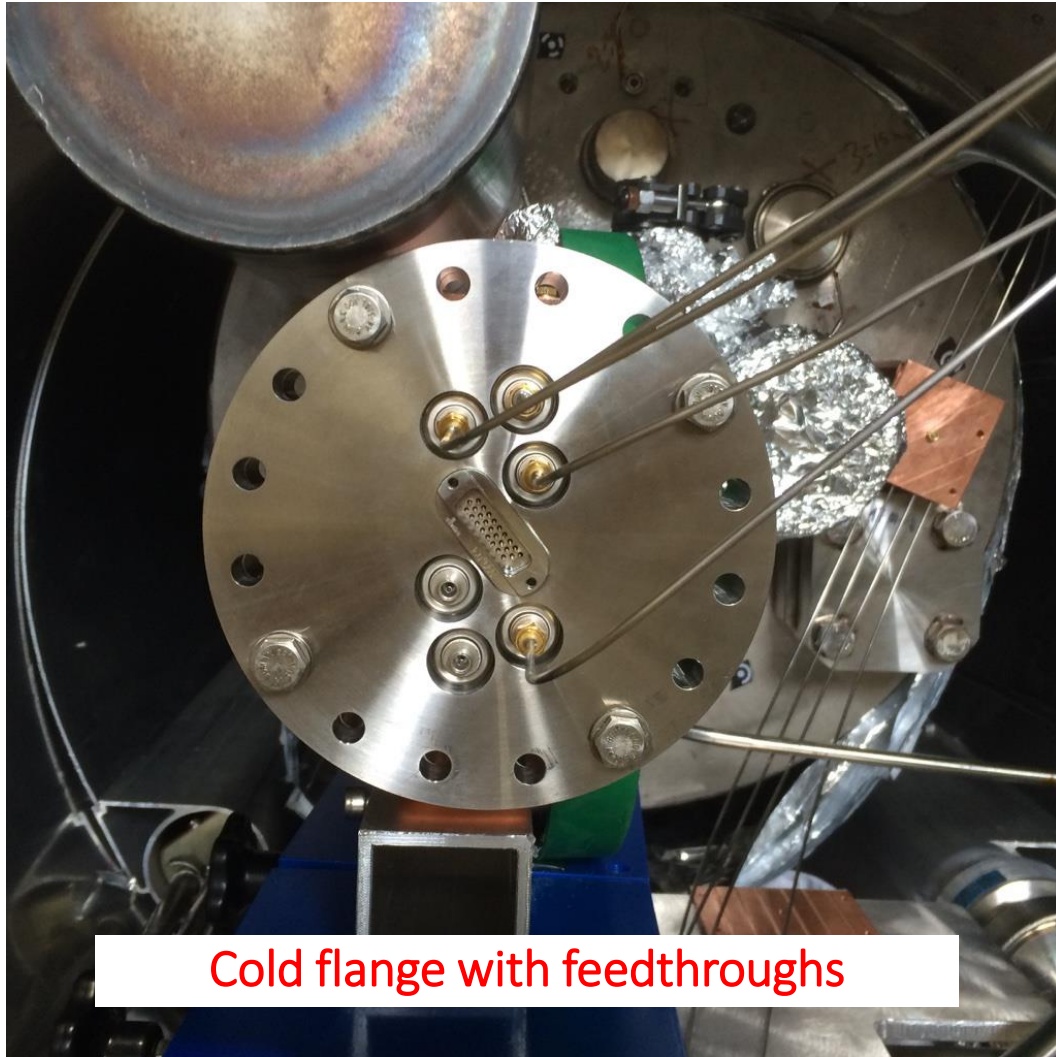


(*) arXiv:1610.02580v1 8 Oct 2016

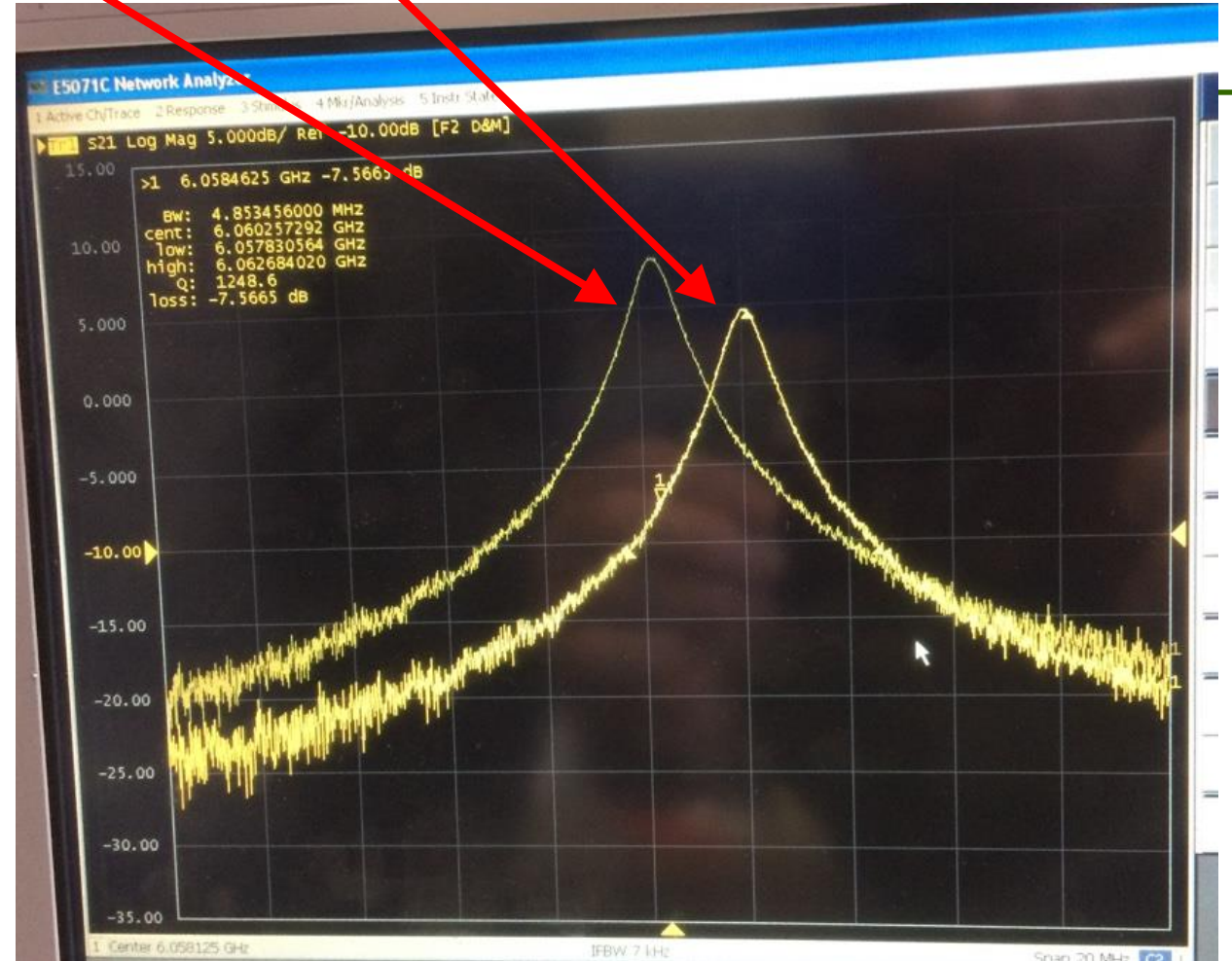
Status: One year since project started

- **November 2015:** Project start after CERN SPSC approval
- **Jan-May 2016:** cavity prep, integration with magnet, vessel design and construction, OK to install
- **June 2016:** cavity installation and testing
 - Resonance measured after installation, before and after start of vacuum pumping
 - Resonant frequency shifted up by ~ 1.8 MHz, roughly corresponding to the change in electrical permittivity of the medium, at room temperature.

Cavity resonance before and after vacuum pumping



Cold flange with feedthroughs



Resonant frequency shifted up by ~ 1.8 MHz, corresponding to the change in electrical permittivity of the medium, at room temperature.

Status: One year since project started

- August/September 2016:
 - Magnet cold (1.8 K)
 - Fundamental resonant frequency re-measured: shifted up from to ~ 6.078 GHz \rightarrow the cavity behaves as expected.
 - Magnet ramped to 13,000 A (operating field): no significant changes in the resonance
 - Quench tests:
 - Magnet quenched from low and operating field (~ 9 T): Cavity unaffected.
 - \rightarrow cavity mechanically and electrically stable
- November 28: ~ 1 week data acquisition before magnet shut down

Cavity unchanged and stable before and after magnet quench

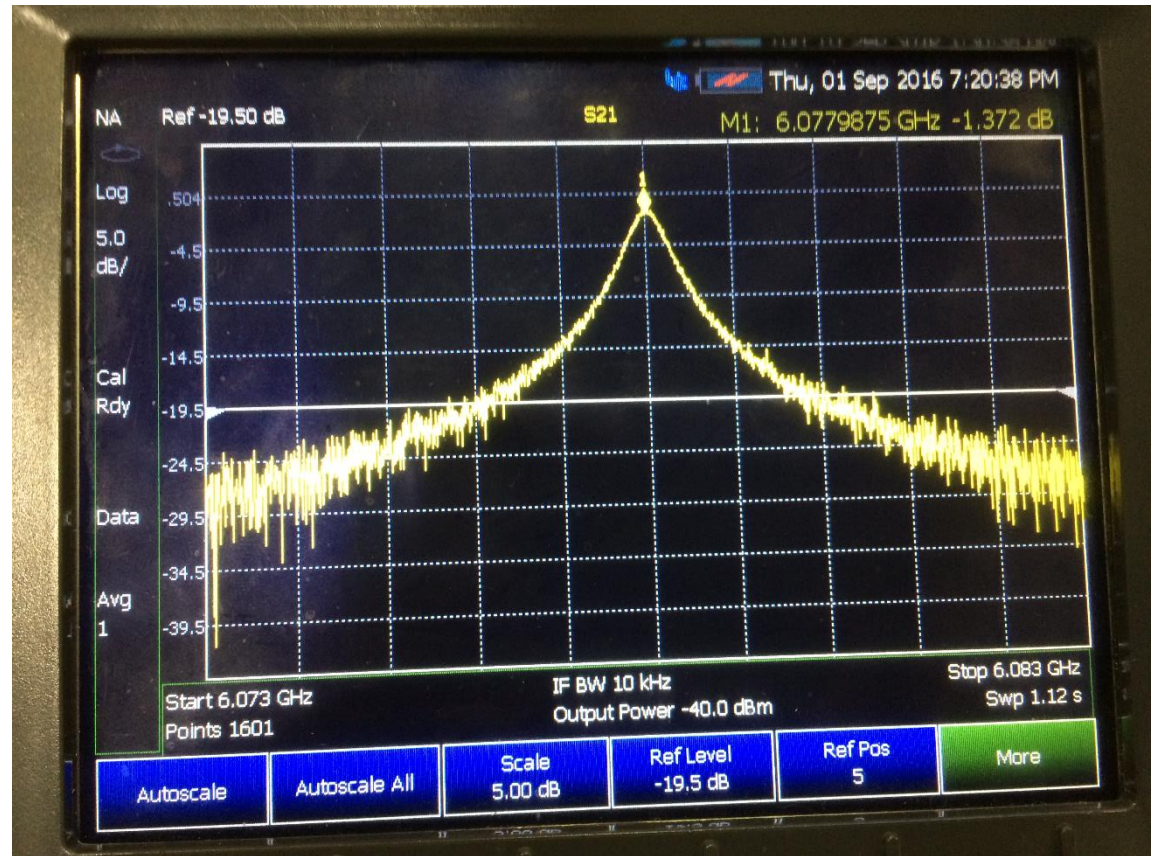
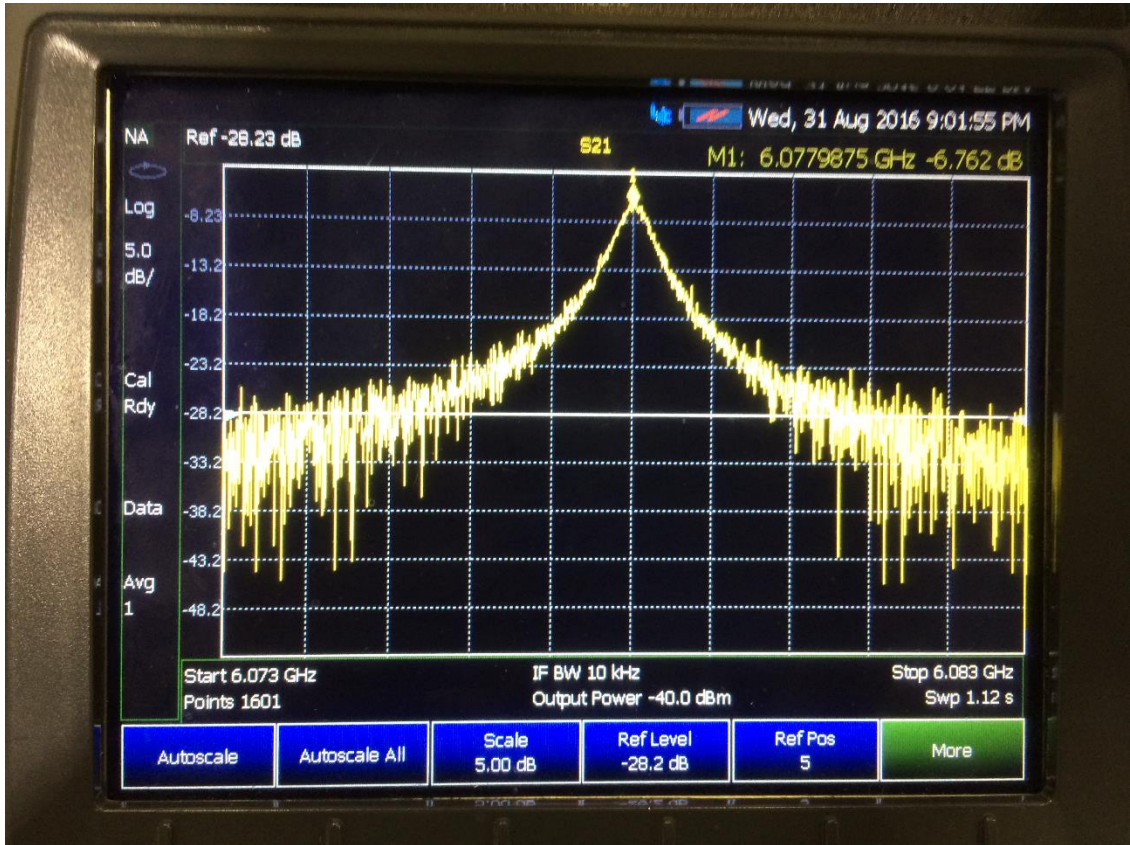
Center freq. = 6.078 GHz

Magnet cold

Before magnet quench (full field)

After magnet quench

Amplitude



Resonant frequency unchanged

Frequency

A closer look at the resonance with and without the magnetic field

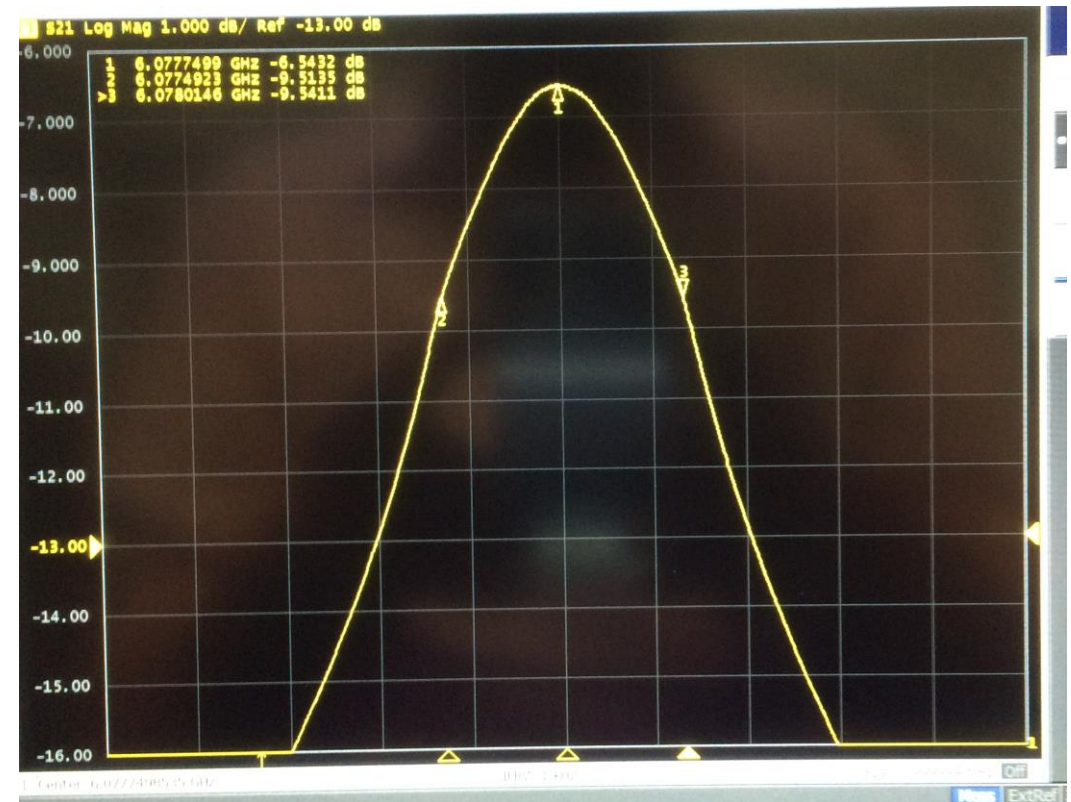
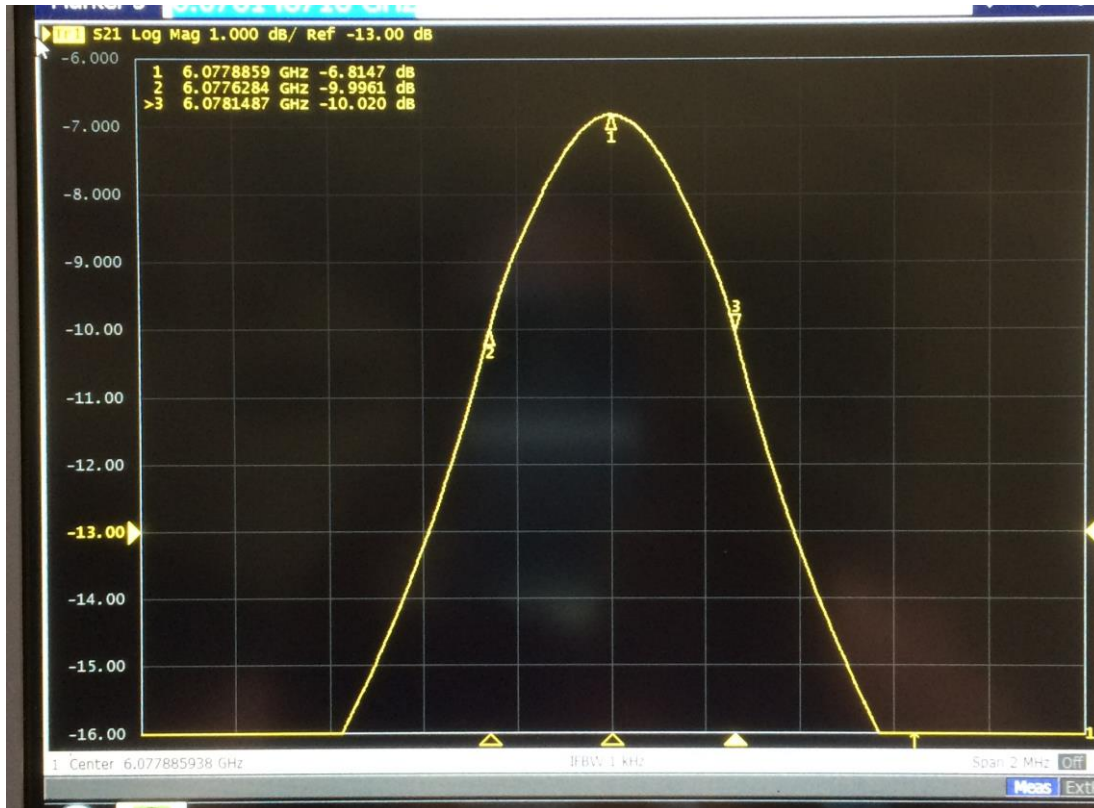
Center freq. = 6.078 GHz

Magnet cold

Magnet off

Magnet on

Amplitude



Frequency

Resonant frequency shifted by -136 KHz, 2 parts in 100,000.

Installed Cavity

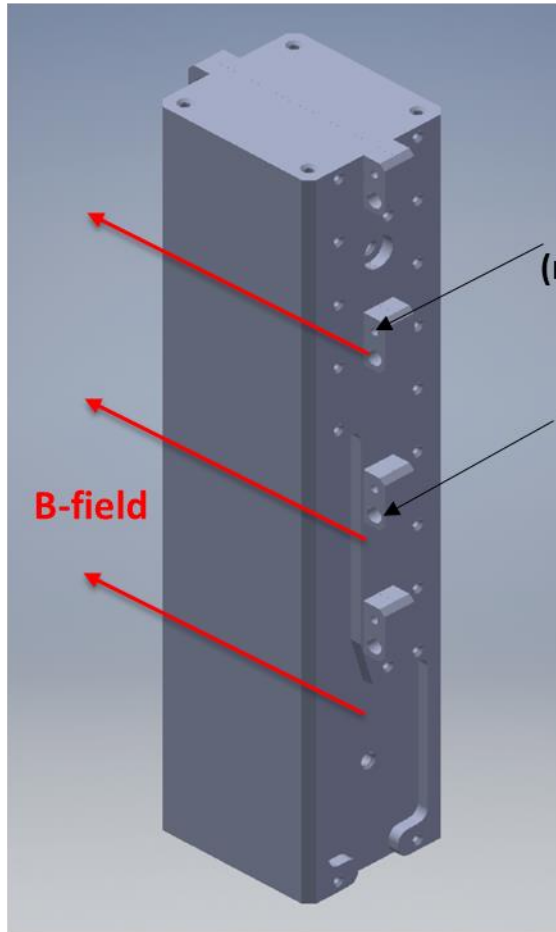
- Fundamental mode $f \cong 6.078$ GHz (low-T)
- $Q \cong 10,000$ (room-T)
- Dimensions: 138 mm X 25 mm X 23 mm.
- Material: stainless steel 10-micron thick electrodeposited copper layer
- Longitudinally split
- Magnetically coupled

Setup

- Cavity signals to a cryogenic low noise HEMT amplifier outside of the cold bore, inside a vacuum vessel
- Two temperature sensors: one on cavity, one near the cryogenic amplifier
- Cavity temperature 3.8 K
- Amplifier temperature 2.5 K
- 2-axis Hall probe near the cavity; 3-axis Hall probe near the amplifier

The RF cavity installed inside the bore

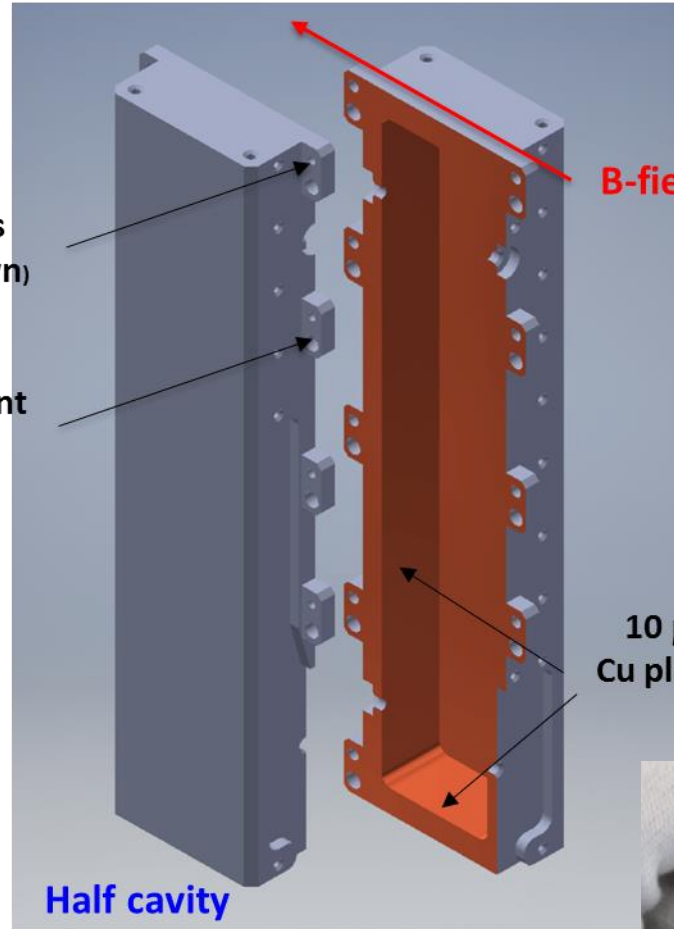
“Inspired” by LHC beam screen design and testing , but 7 times thinner copper plating



Assembled cavity

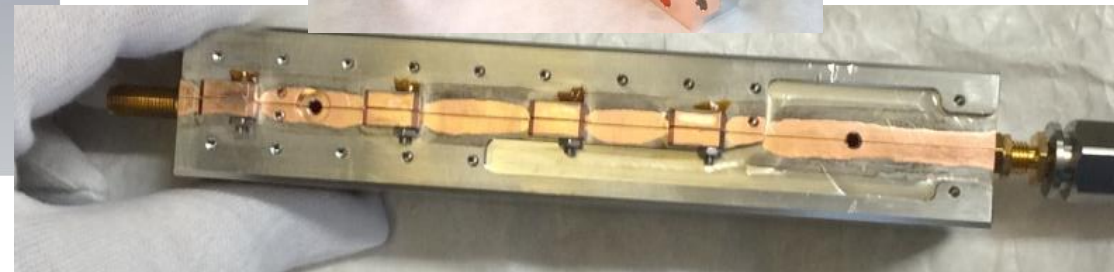
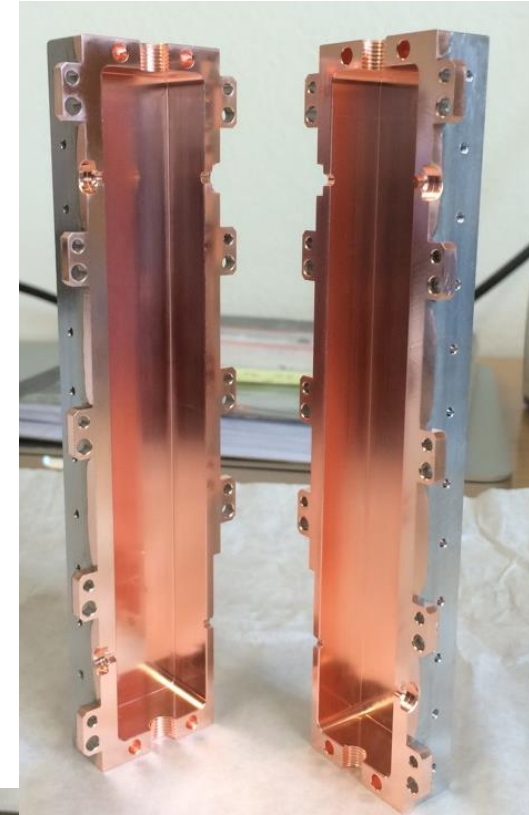
M2 bolts
(not shown)

Alignment
holes

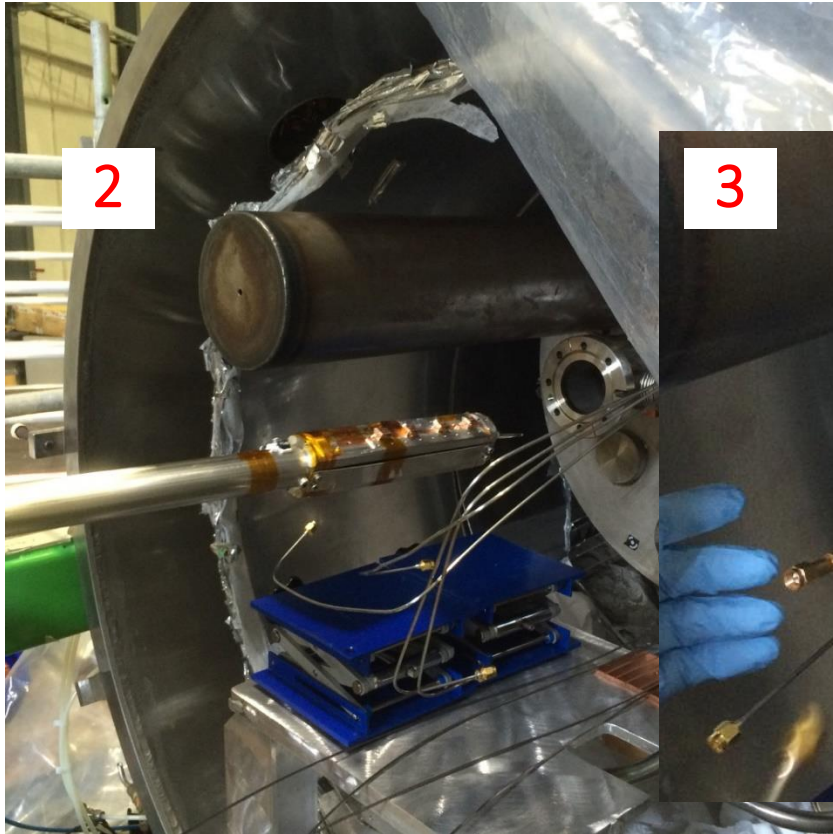


Half cavity

10 μm
Cu plating



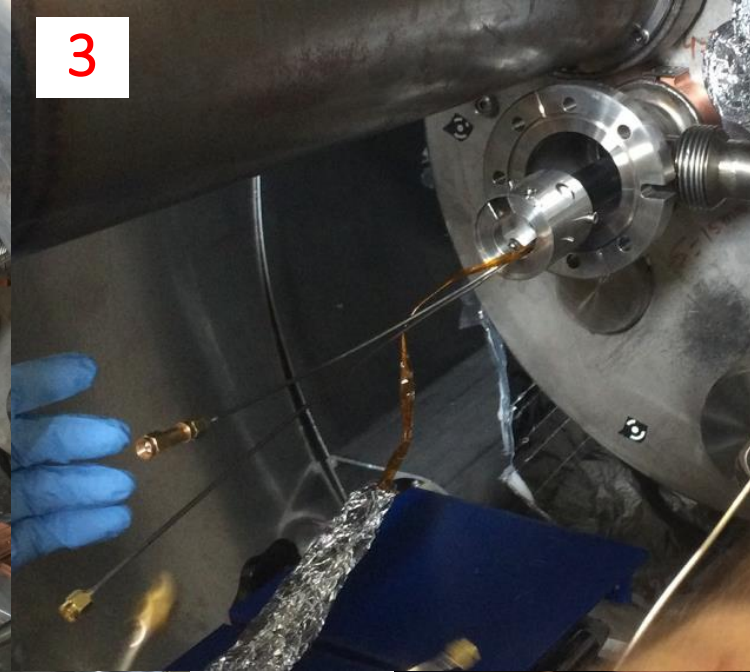
Installation



2

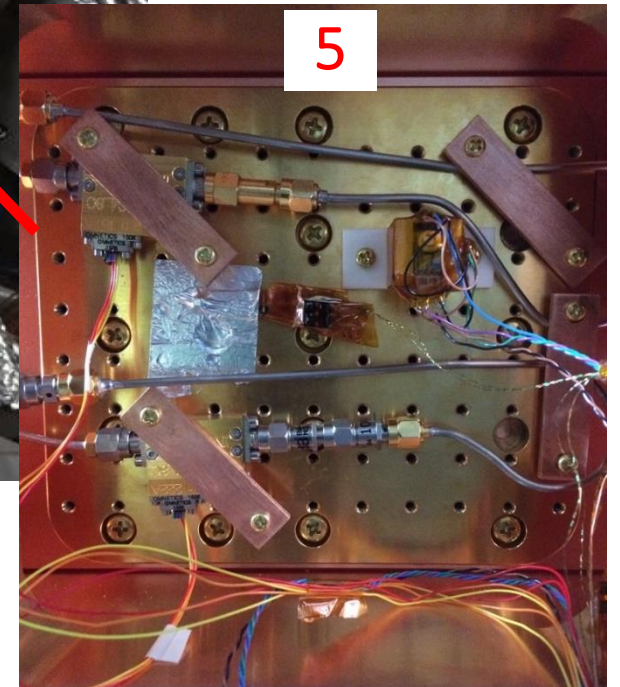


4



3

VACUUM VESSEL



5

Low-temperature electronics in vac. vessel



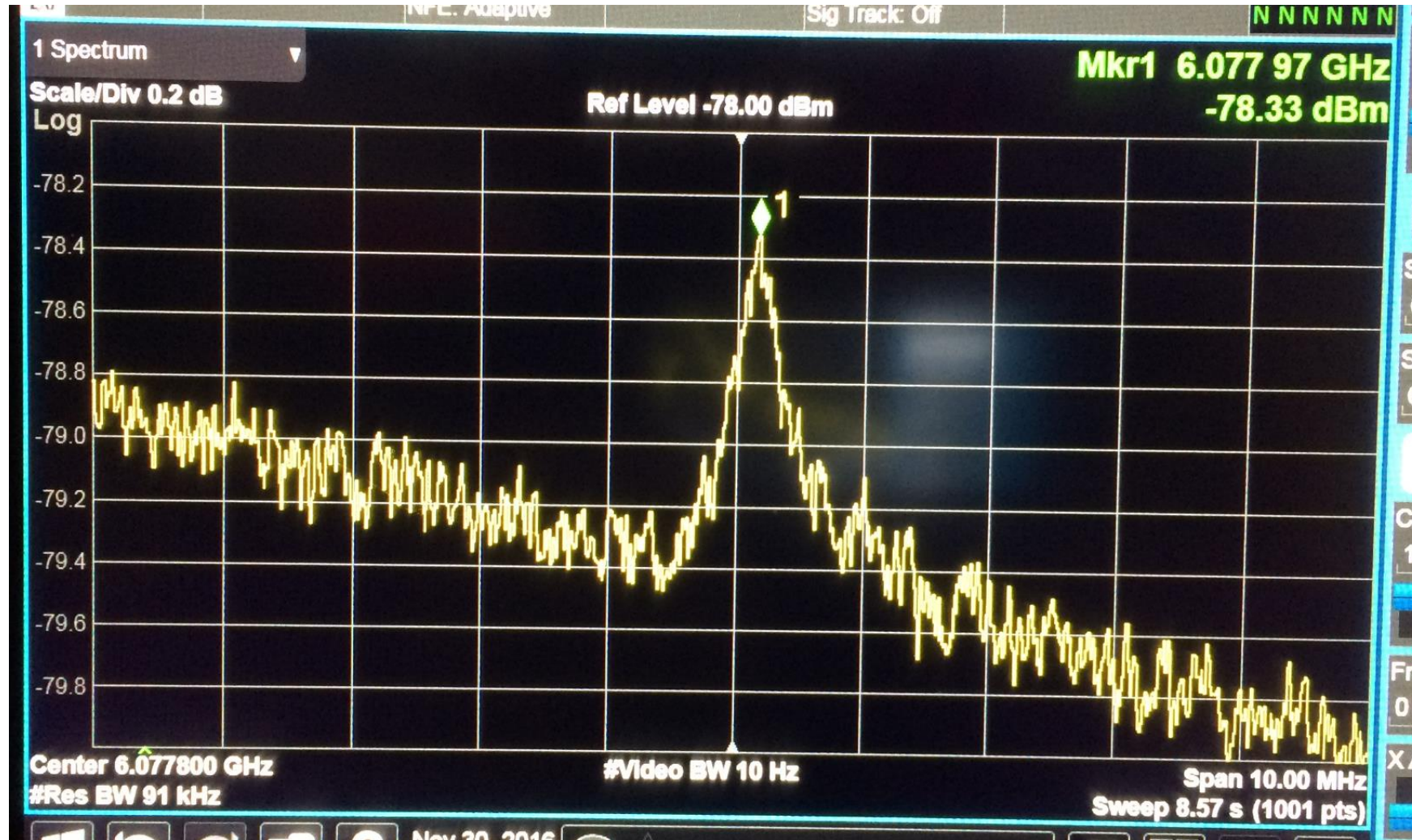
1

Insertion tube

Cavity

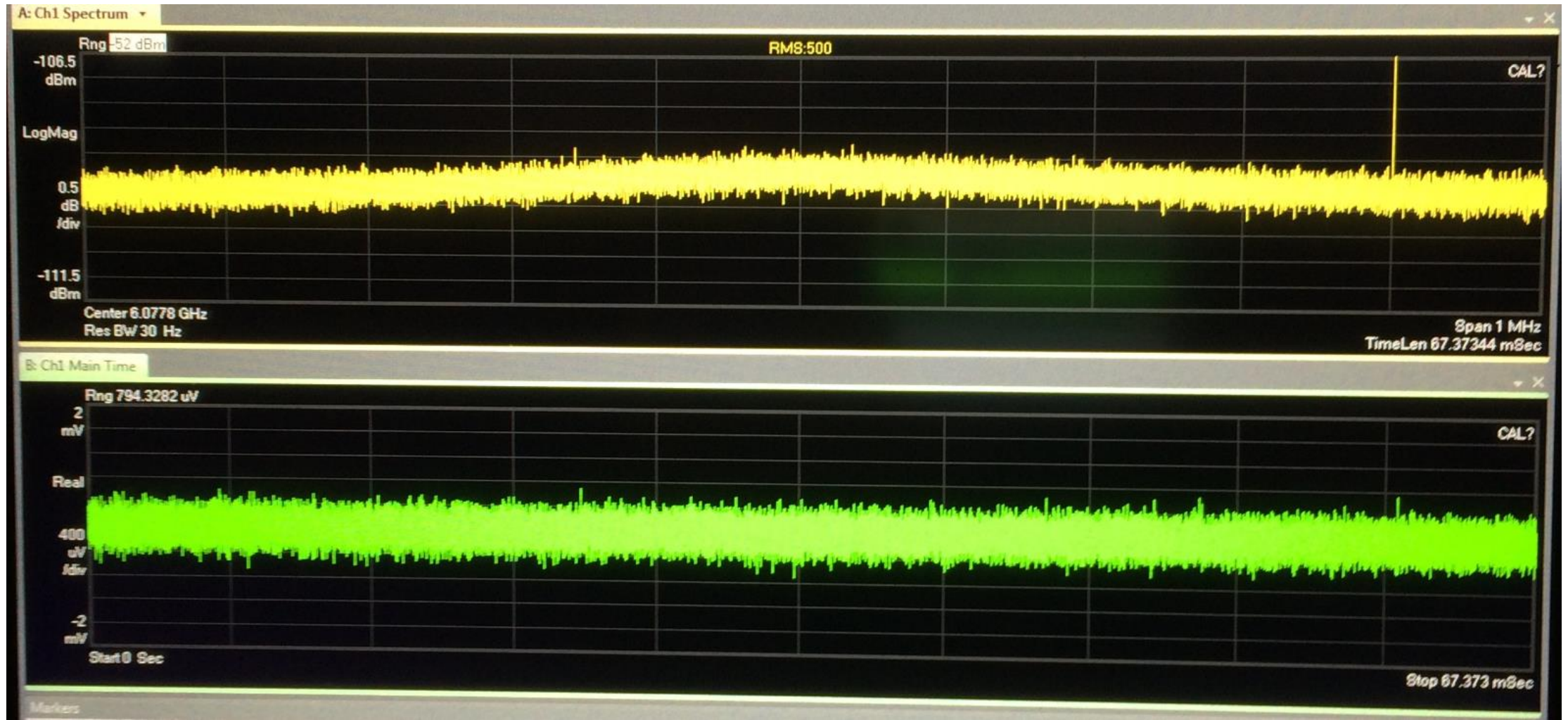
A sample from recent data acquisition

Amplitude



Frequency

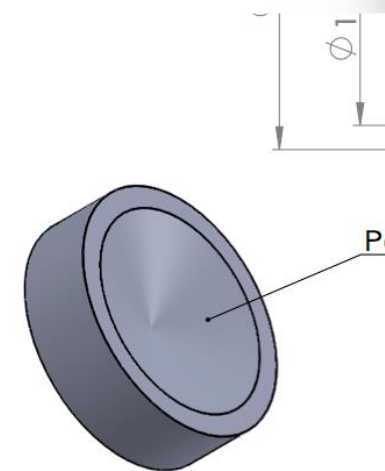
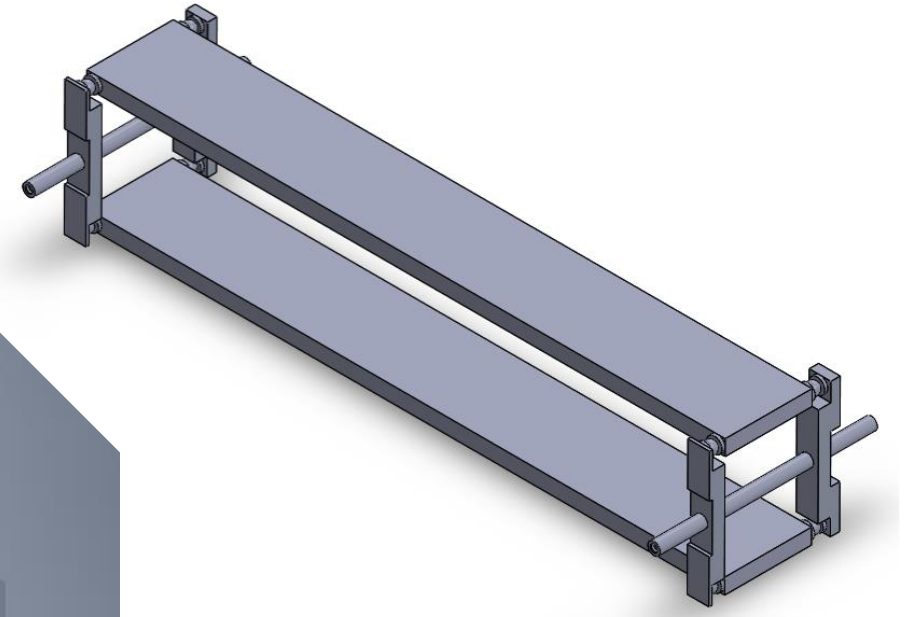
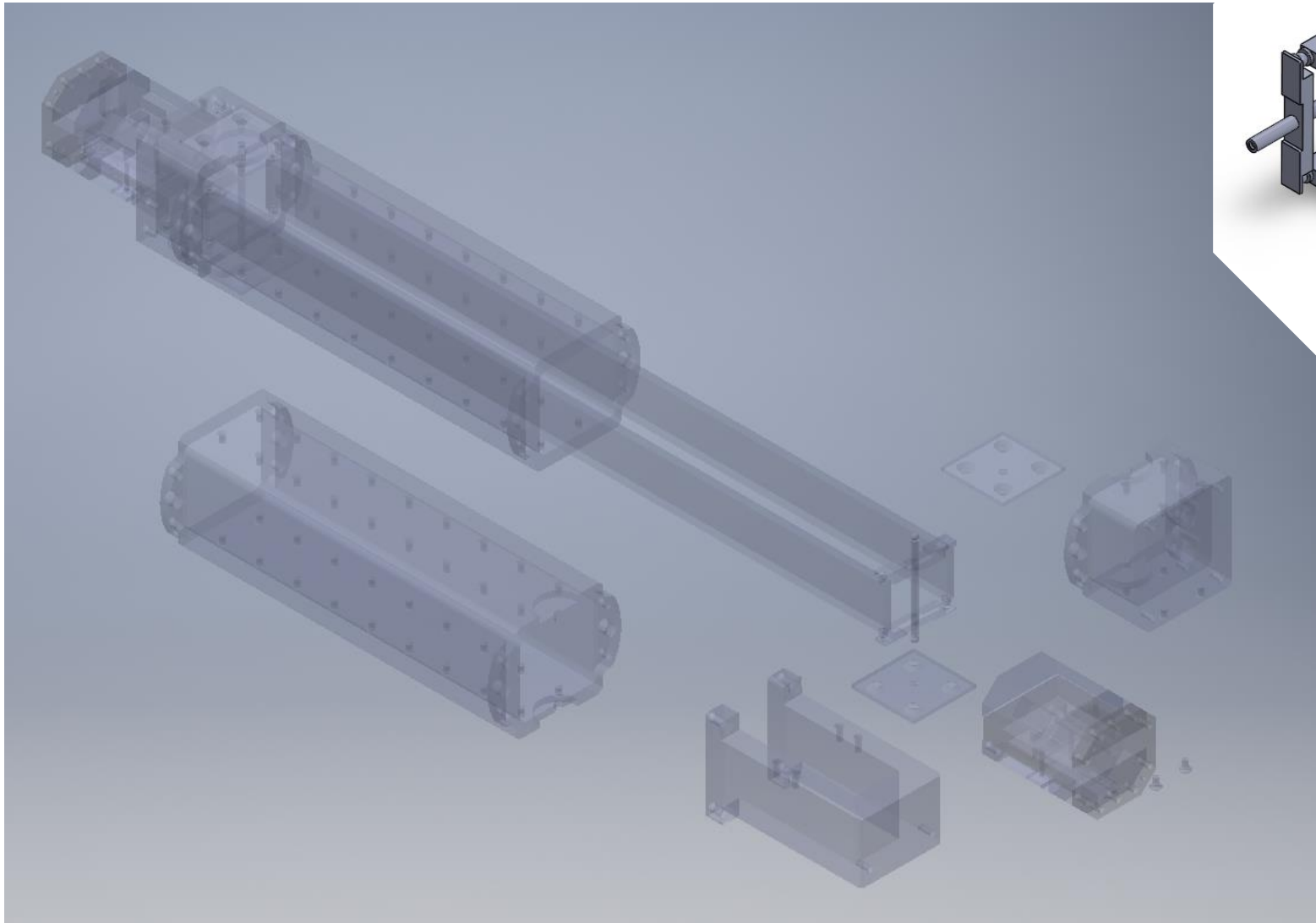
A sample from recent data acquisition



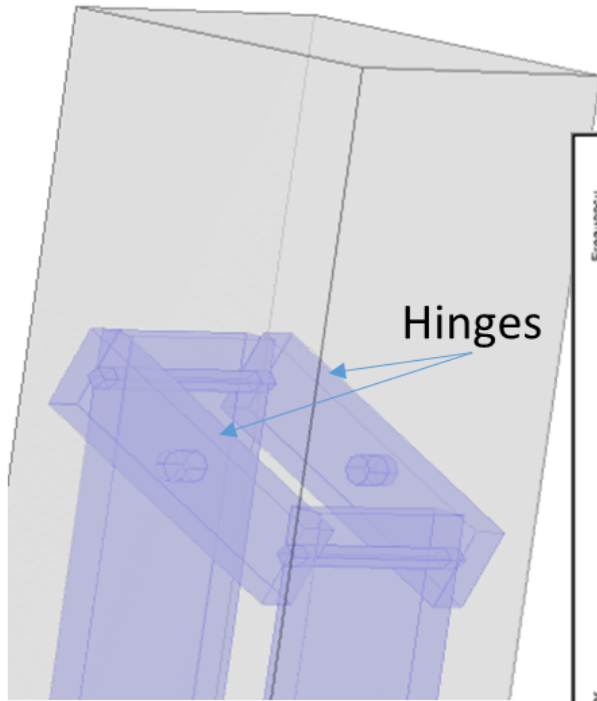
CAST-CAPP R&D

- Design and fabrication of longer rectangular cavities.
- Amplifiers in high magnetic fields.
- Cavity tuning and coupling for multi-cavity operation.
- Multi-cavity operation

Next: Cavity tuning with dielectric bars

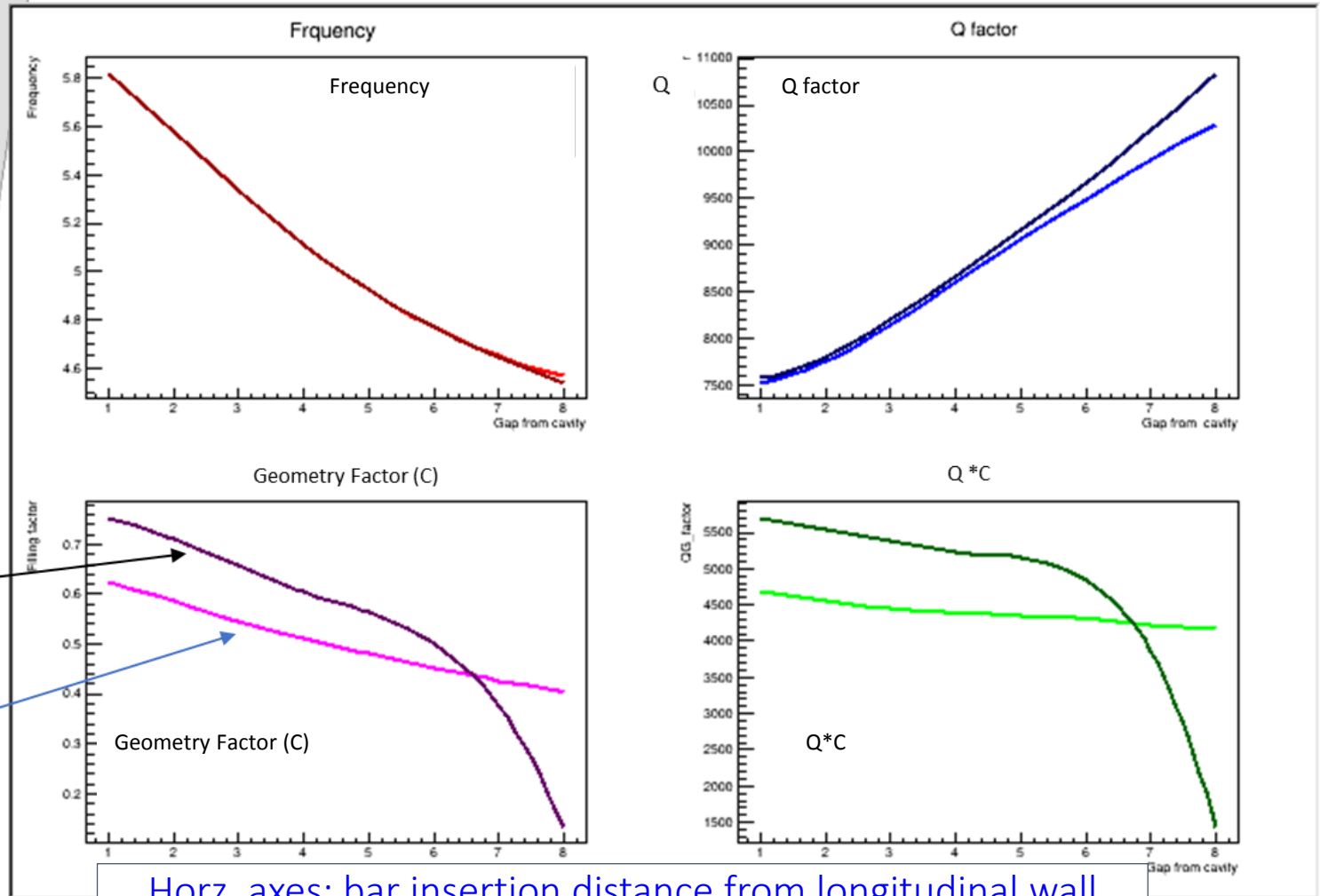


Cavity tuning with dielectric bars

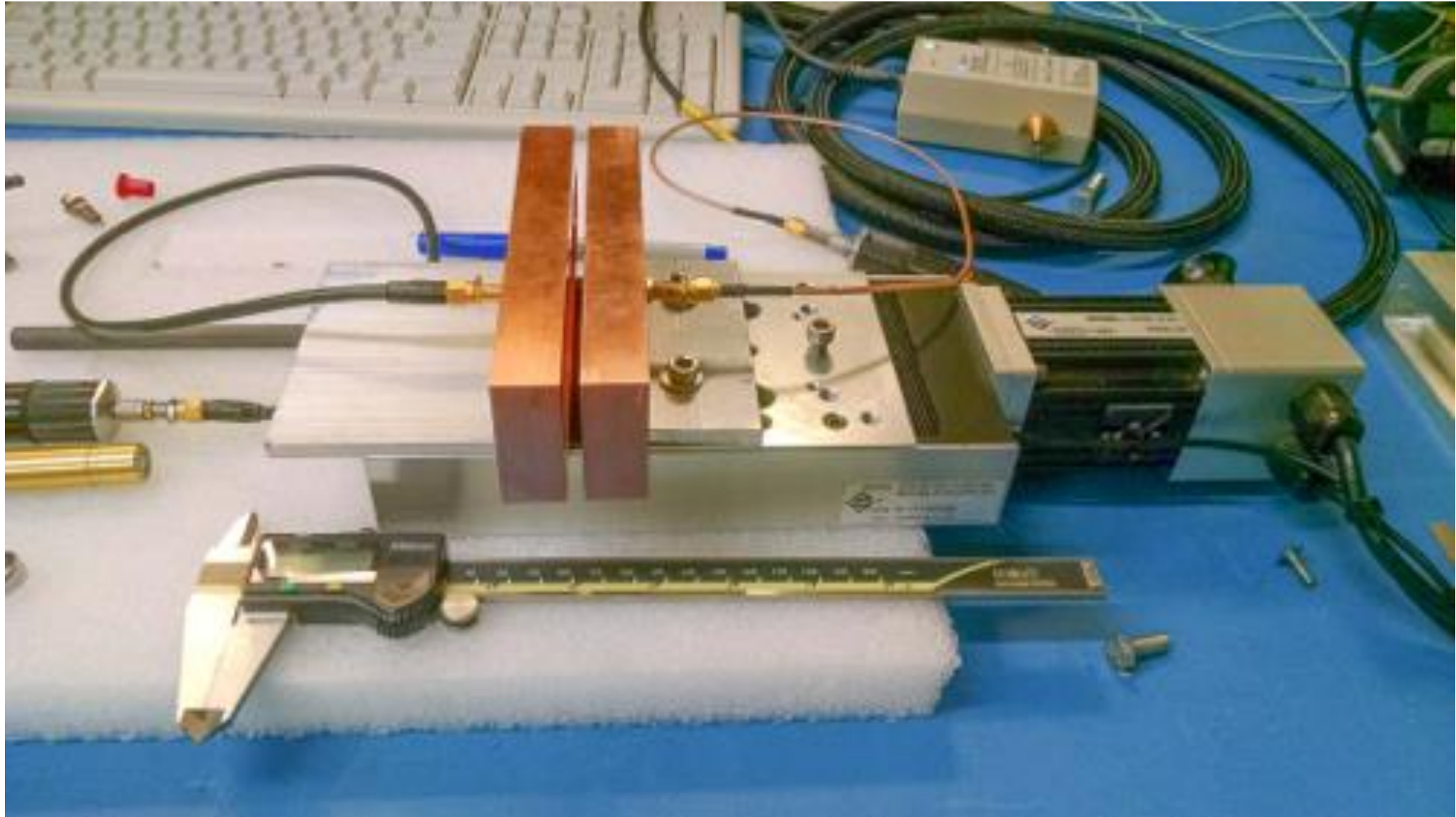


With hinges

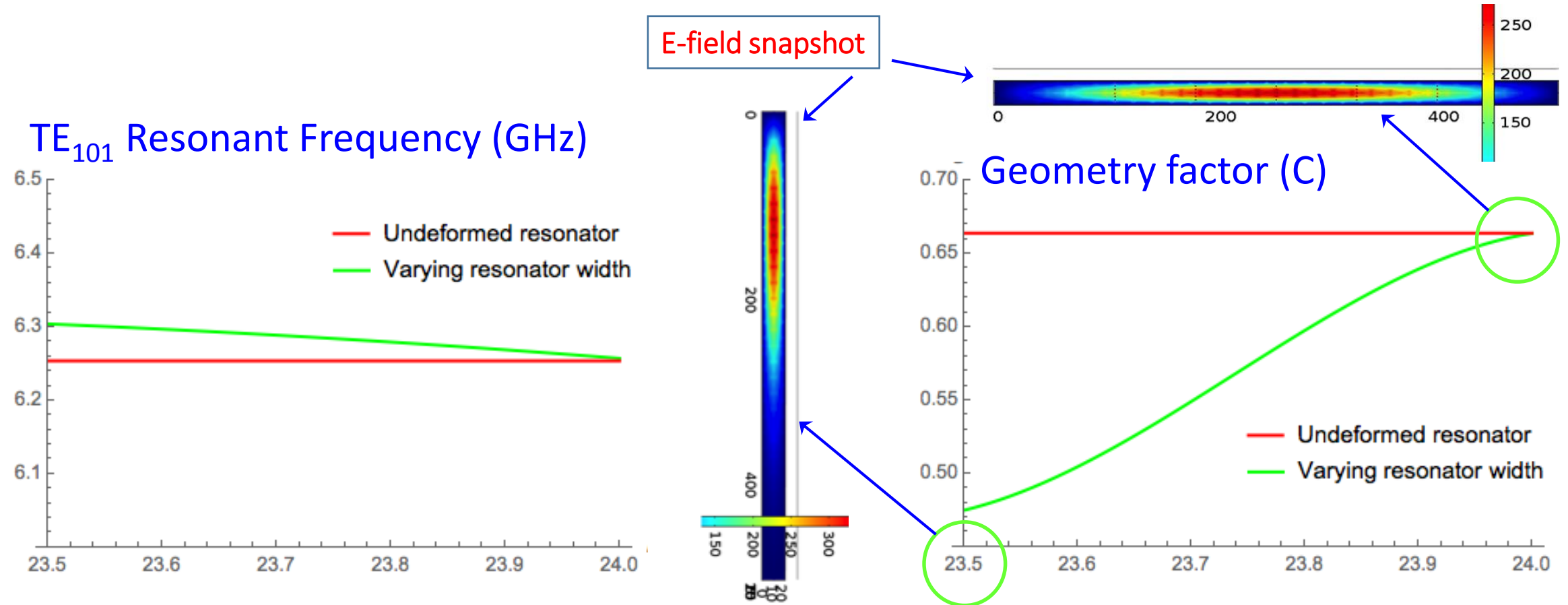
Without hinges



Alternative tuning: Split cavity



Cavity engineering: Mechanical tolerances vs. mode localization. Cavity: 2.5cm X 2.4cm X 50cm



Horizontal axis: varying resonator width on one side (mm)

Summary:

- For the first time, a microwave cavity has been installed in a dipole magnet for axion search
- The cavity is stable inside the magnet, unaffected by quench forces
- We have taken a few days of data
- This feasibility study shows that an axion search experiment using the Sikivie technique is sustainable in a large dipole magnet
- We plan to move to multiple cavities for higher sensitivity